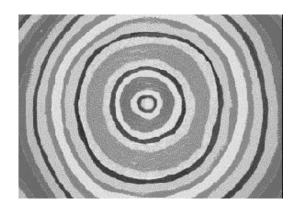


Listening to Gravitational Waves:

Einstein's Songlines from the Universe

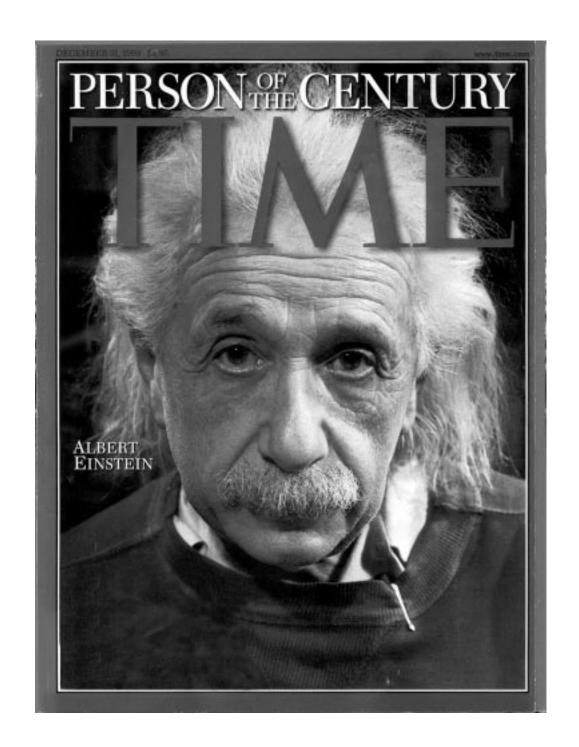
Barry C. Barish

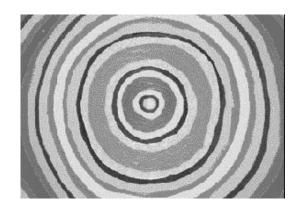




Albert Einstein



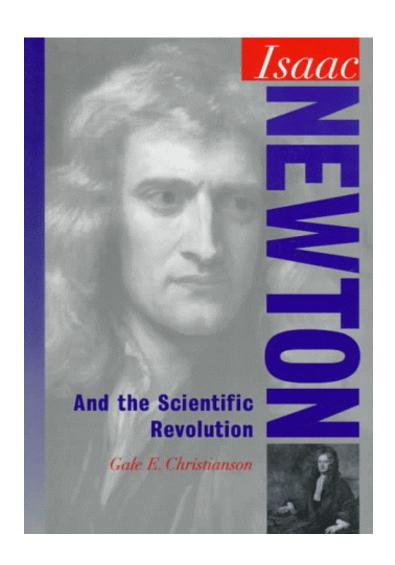


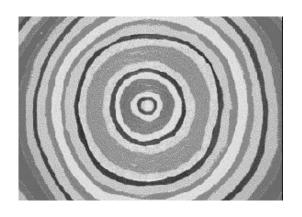


- Perhaps the most important scientist of all time!
- Invented the scientific method in *Principia*
- Greatest scientific achievement: universal gravitation



Sir Isaac Newton



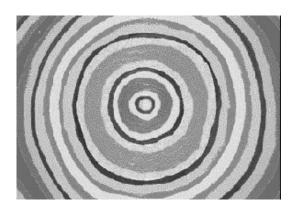


Scientific Method *Principia*

- We are to admit no more causes of natural things such as are both true and sufficient to explain their appearances
- the same natural effects must be assigned to the same causes
- **18** qualities of bodies are to be esteemed as universal
- propositions deduced from observation of phenomena should be viewed as accurate until other phenomena contradict them.

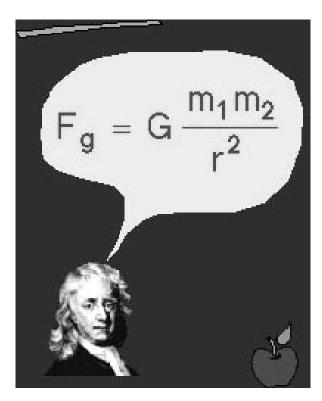




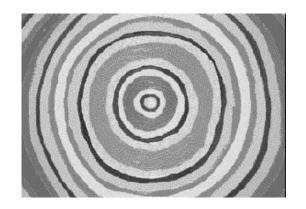


Newton Universal Gravitation

- Three laws of motion and law of gravitation (centripetal force) disparate phenomena
 - » eccentric orbits of comets
 - » cause of tides and their variations
 - » the precession of the earth's axis
 - » the perturbation of the motion of the moon by gravity of the sun
- Solved most known problems of astronomy and terrestrial physics
 - » Work of Galileo, Copernicus and Kepler unified.



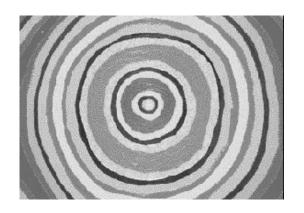






Albert Einstein

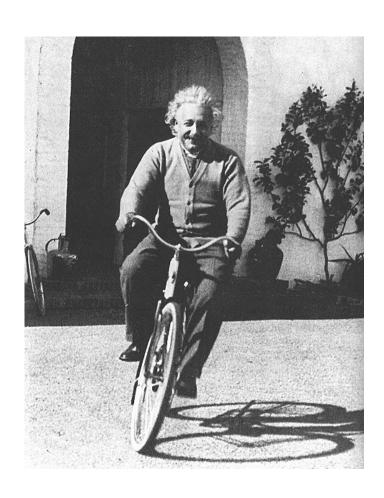
- The Special Theory of Relativity (1905) overthrew commonsense assumptions about space and time.
 Relative to an observer, near the speed of light, both are altered
 - » distances appear to stretch
 - » clocks tick more slowly
- The General Theory of Relativity and theory of Gravity (1916)

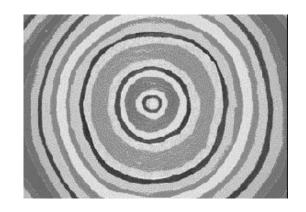


Einstein's *Spacetime Wrinkles*

- Discards concept of absolute motion; instead treats only relative motion between systems
- space and time no longer viewed as separate; rather as four dimensional space-time
- gravity described as a warpage of space-time, not a force acting at a distance







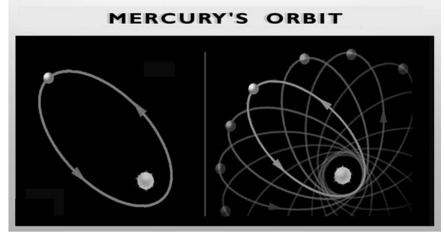
Einstein's Theory of Gravitation experimental tests

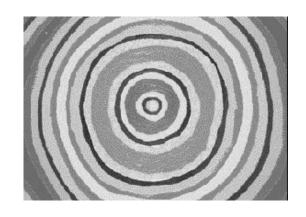
"Einstein Cross"
The bending of light rays
gravitational lensing



Mercury's orbit
perihelion shifts forward
twice Newton's theory



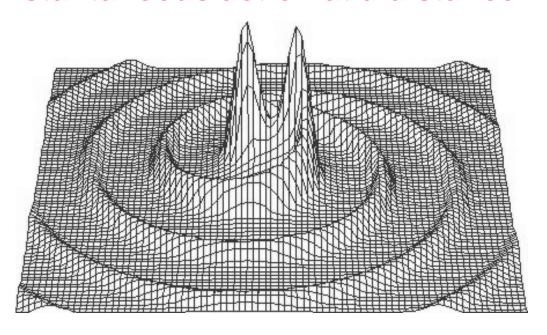




Einstein's Theory of Gravitation

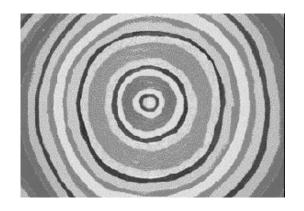
experimental tests

Newton's Theory "instantaneous action at a distance"

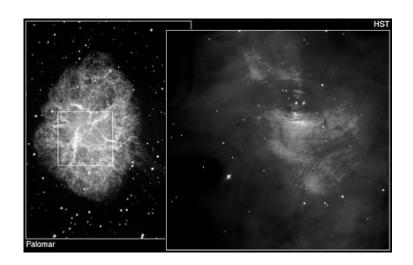


Einstein's Theory information carried by gravitational radiation at the speed of light



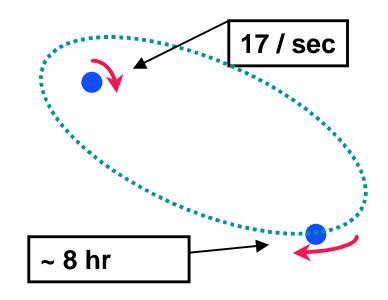


Gravitational Waves the evidence

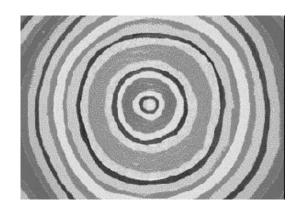


Neutron Binary System

PSR 1913 + 16 -- Timing of pulsars





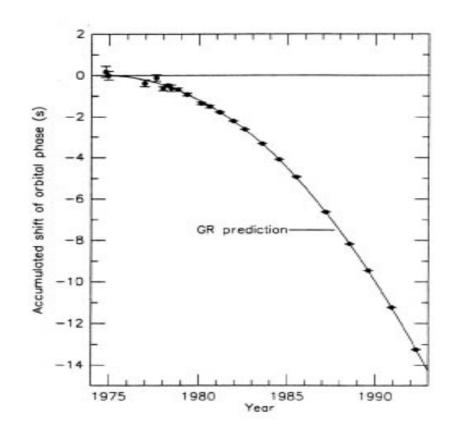


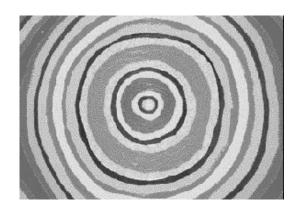
Hulse and Taylor results

emission of gravitational waves

- due to loss of orbital energy
- period speeds up 14 sec from 1975-94
- measured to ~50 msec accuracy
- deviation grows quadratically with time







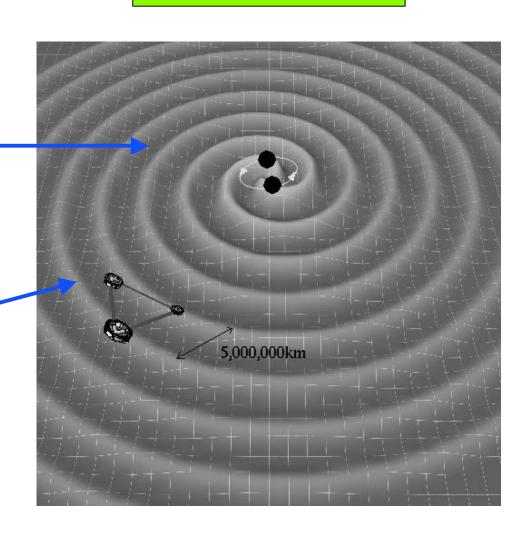
Einstein's

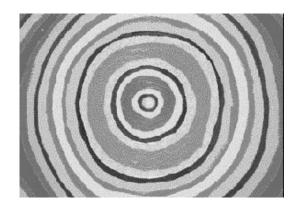
Songlines

Radiation of Gravitational Waves from binary inspiral system

LISA

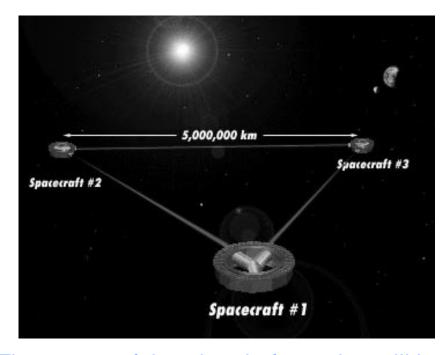






Interferometers

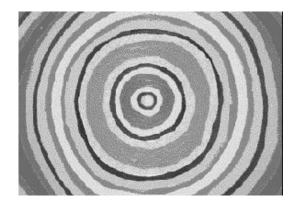
The Laser
Interferometer
Space
Antenna
(LISA)



The center of the triangle formation will be in the ecliptic plane

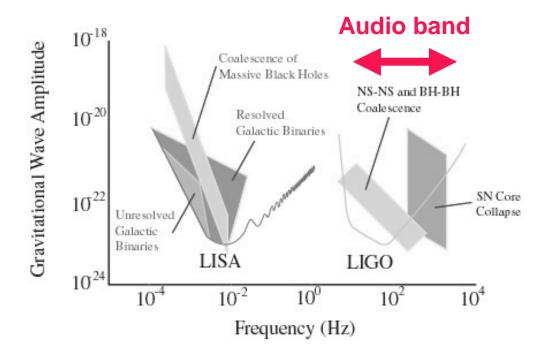
1 AU from the Sun and 20 degrees behind the Earth.



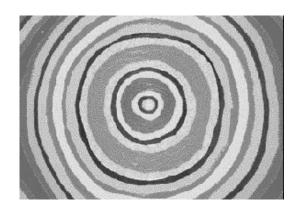


- EM waves are studied over ~20 orders of magnitude
 - **»** (ULF radio \rightarrow HE γ rays)
- Gravitational Waves over
 ~10 orders of magnitude
 - » (terrestrial + space)

Astrophysics Sources frequency range





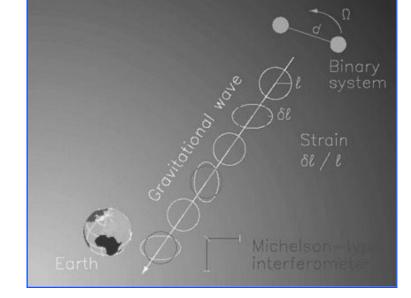


Interferometers terrestrial

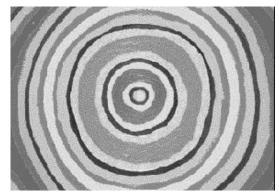
Suspended mass Michelson-type interferometers on earth's surface detect distant astrophysical sources

International network (LIGO, Virgo, GEO, TAMA) enable locating sources and decomposing polarization of

gravitational waves.

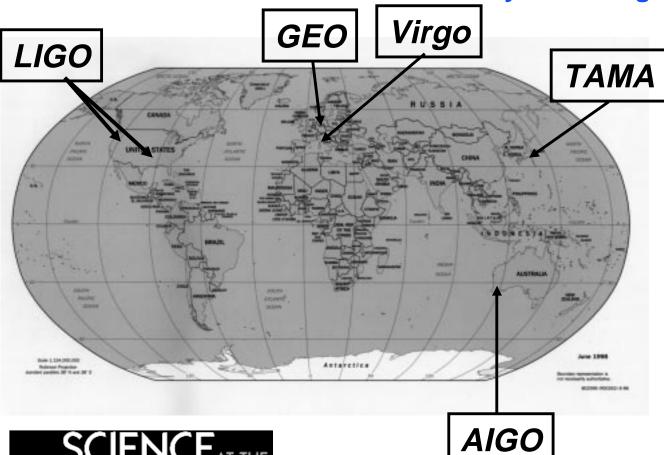






International Network

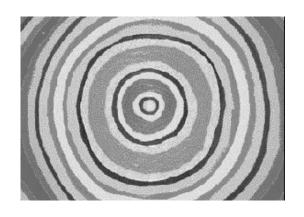
Simultaneously detect signal (within msec)



detection confidence

locate the sources

decompose the polarization of gravitational waves



Gravitational Waves the effect

Leonardo da Vinci's Vitruvian man







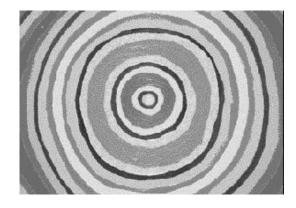


 stretch and squash in perpendicular directions at the frequency of the gravitational waves

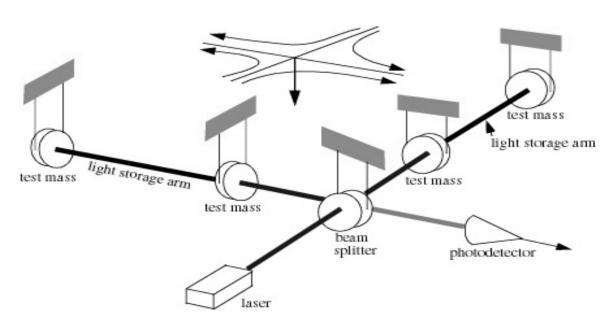


The effect is greatly exaggerated!!

If the man was 4.5 light years high, he would grow by only a 'hairs width'
LIGO (4 km), stretch (squash) = 10⁻¹⁸ m will be detected at frequencies of 10 Hz to 10⁴ Hz. It can detect waves from a distance of 600 10⁶ light years

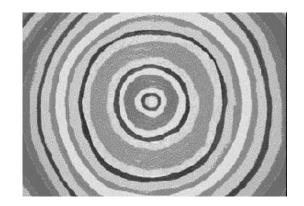




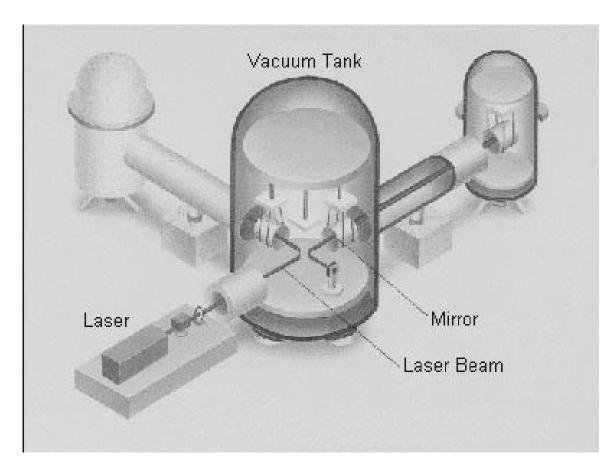


- The concept is to compare the time it takes light to travel in two orthogonal directions transverse to the gravitational waves.
- The gravitational wave causes the time difference to vary by stretching one arm and compressing the other.
- The interference pattern is measured (or the fringe is split) to one part in 10¹⁰, in order to obtain the required sensitivity.





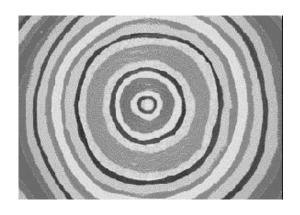




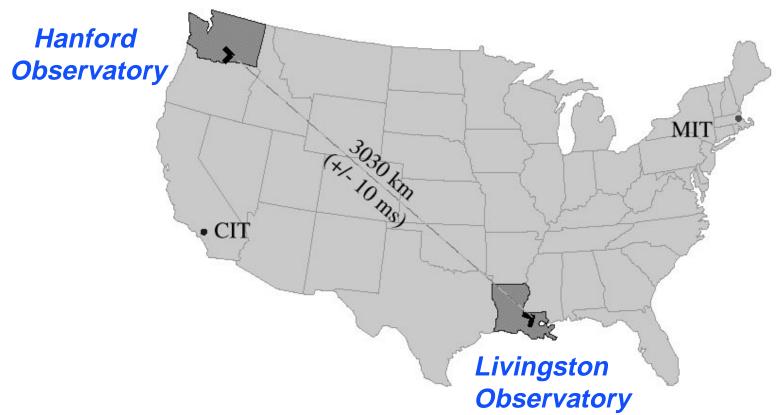
Schematic of TAMA 300m interferometer

The effects of gravitational waves appear as a fluctuation in the phase differences between two orthogonal light paths of an interferometer.

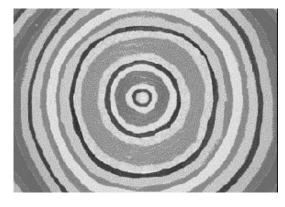




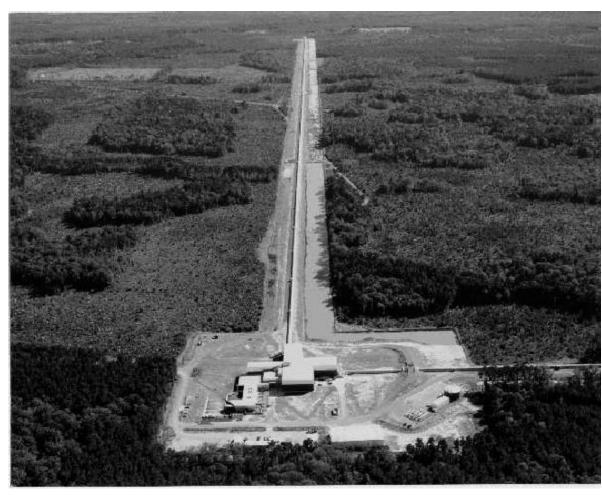




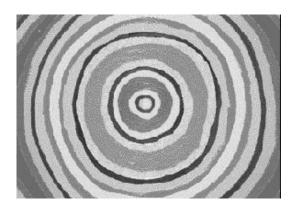




LIGO Livingston

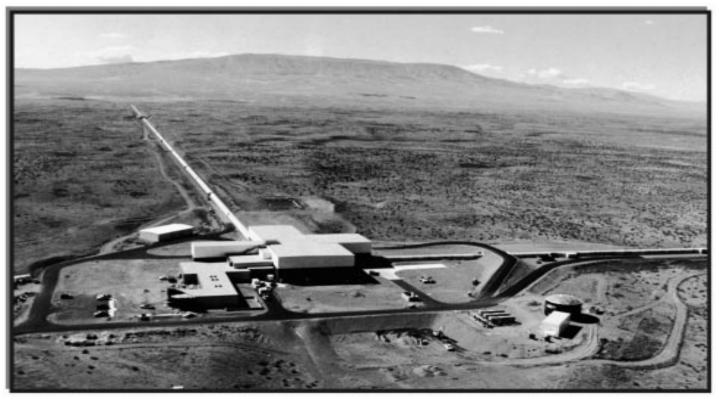




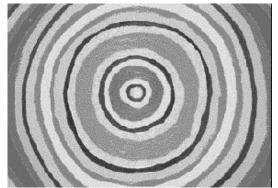


LIGO

Hanford





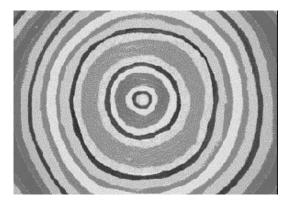






- LIGO beam tube under construction in January 1998
- 65 ft spiral welded sections
- girth welded in portable clean room in the field



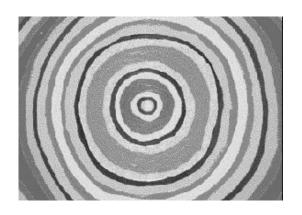


LIGO

vacuum equipment



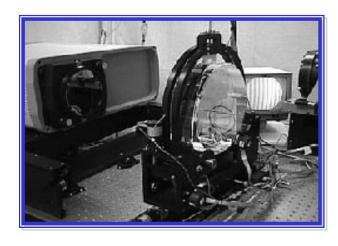


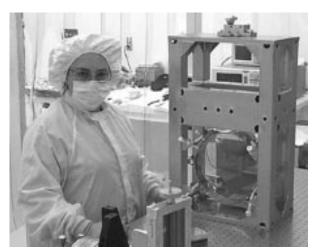


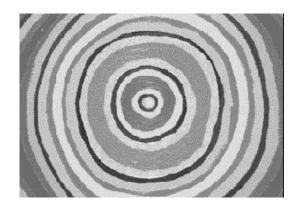
- Optics polished & coated
 - Microroughness within spec. (<10 ppm scatter)
 - » ROC within spec. ($\delta R/R < 5\%$, except for BS)
 - » Coating defects within spec. (pt. defects < 2 ppm, 10 optics tested)</p>
 - Coating absorption within spec. (<1 ppm, 40 optics tested)
- Optics polished at CSIRO in Australia











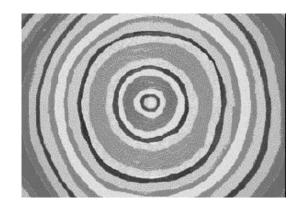




Einstein's Songlines



- AIGO will soon 'listen' for Einstein's Songlines with gravitational waves
- Basic tests of General Relativity will be possible (eg. Black holes)

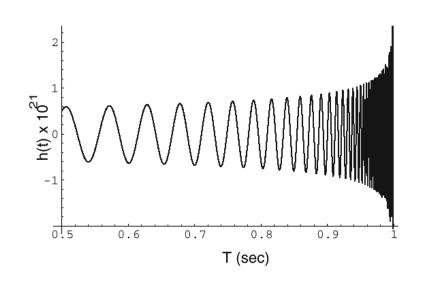


Sources of Gravitational Waves

Inspiral of Neutron Stars

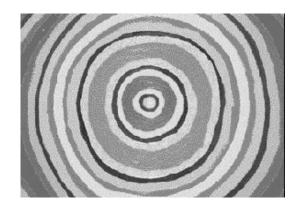


"Chirp Signal"

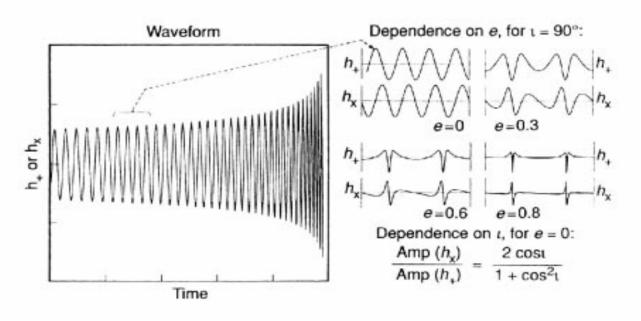








Chirp Signal binary inspiral

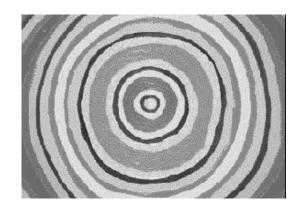


determine



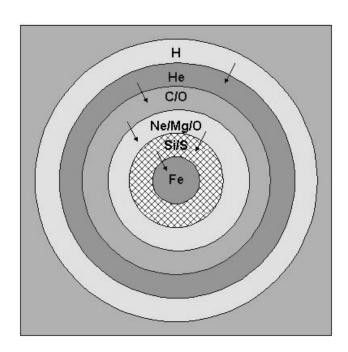
- masses of the two bodies
- •orbital eccentricity e and orbital inclination i





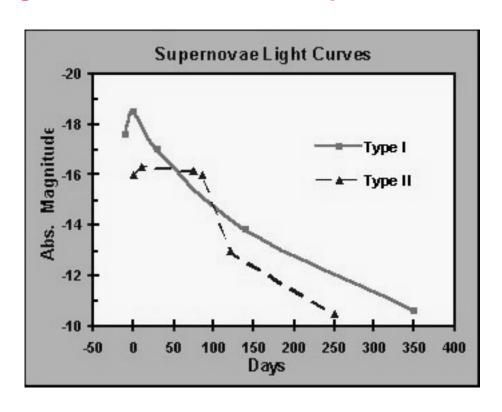
Sources of Gravitational Waves

Supernovae gravitational stellar collapse

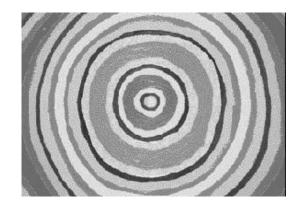


The Collapse



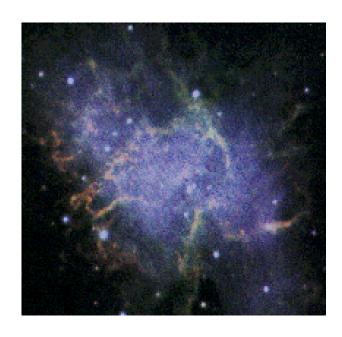


Optical Light Curve



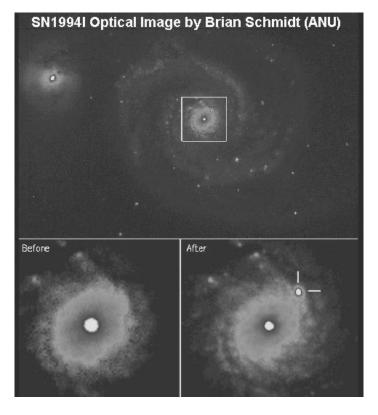
Sources of Gravitational Waves

Supernovae optical observations

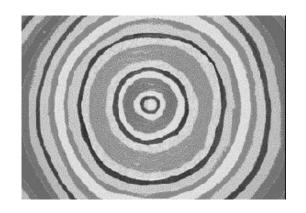


Crab Nebula 1054 AD



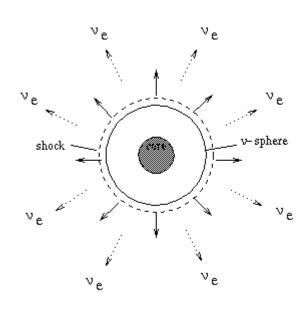


Supernovae - SN1994I

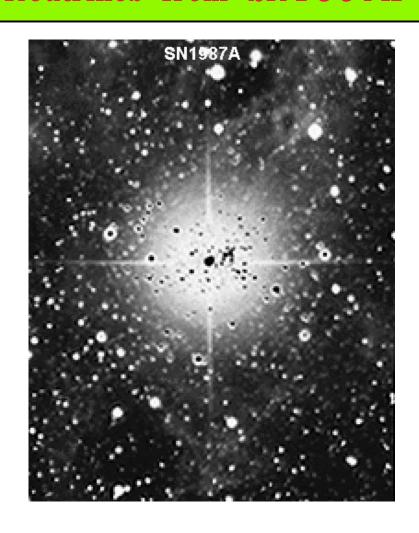


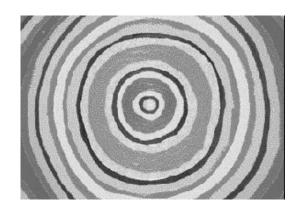
Supernovae

Neutrinos from SN1987A





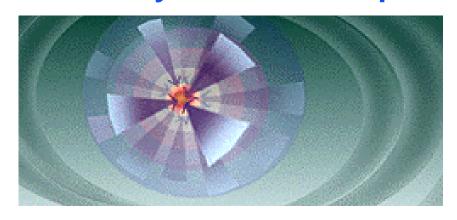




Supernovae

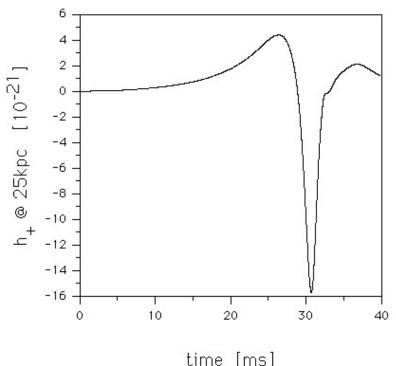
Gravitational Waves

Non axisymmetric collapse

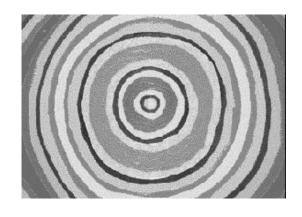


Rate
1/50 yr - our galaxy
3/yr - Virgo cluster

'burst' signal



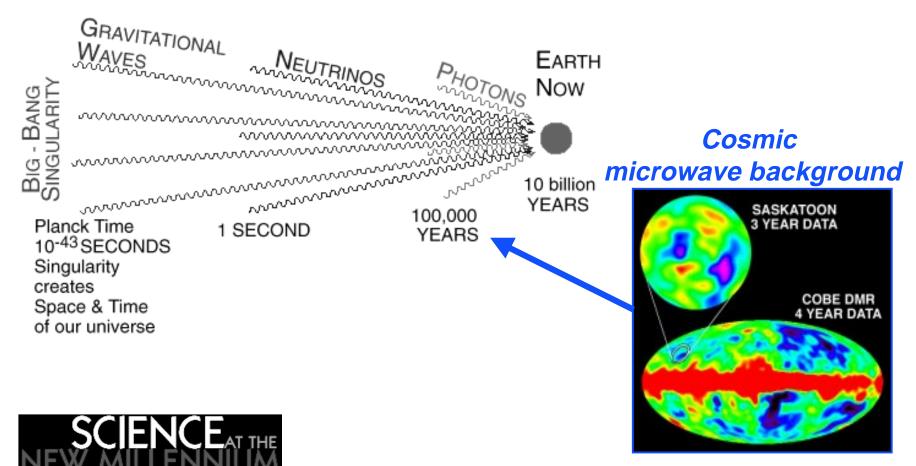


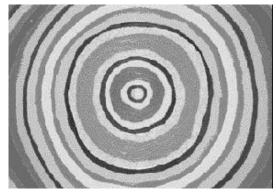


Sources of Gravitational Waves

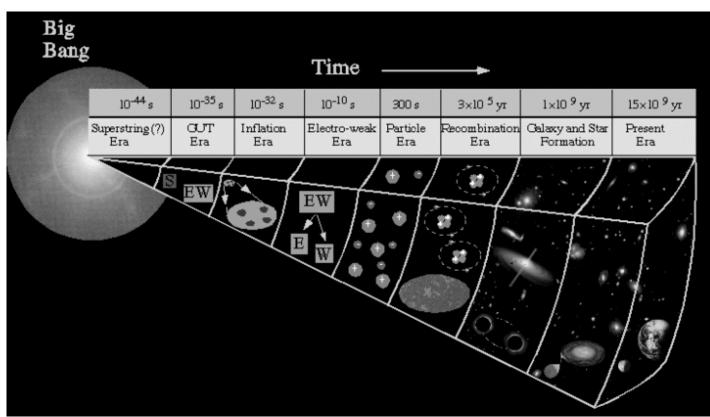
'Murmurs' from the Big Bang

signals from the early universe

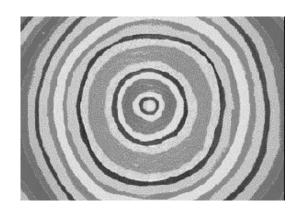




Connect the Beginning of the Universe to Fundamental Physics

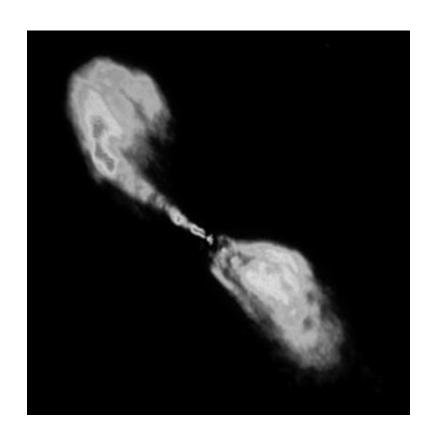




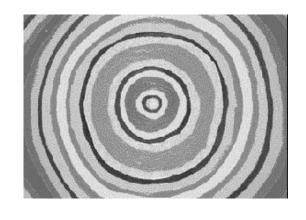


Astrophysics and Cosmology

- More than 95% of the Universe is non luminous matter (dark matter)
- Gravitational waves will open up an entirely new window on the Universe

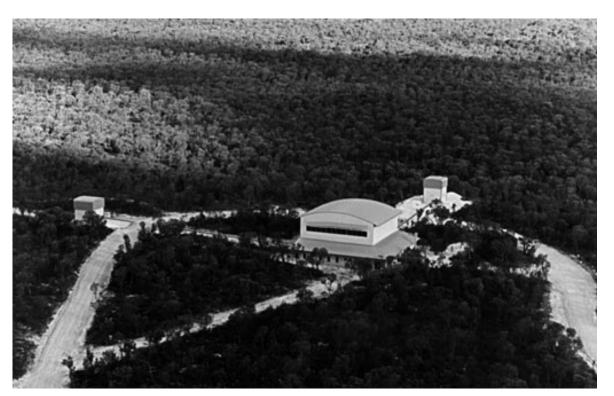








Australian Consortium for Interferometric Gravitational Astronomy (ACIGA)



The University of Western
Australia
The Australian National
University
The University of Adelaide
CSIRO Lindfield
Monash University



The First Stage of the Laser Interferometer Gravitational Wave Observatory