

**New Folder Name** Shot Noise Formulas

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# Comparison of Shot Noise Formulas

R. E. Spero

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## ASSUMPTIONS AND DEFINITIONS

1. Sensitivity is relative to calibration obtained by moving one of the test masses;  $\tilde{x}(f) = L\tilde{h}(f)$ ,  $L$  = length of one arm.
2.  $\mathcal{L} \equiv$  total of losses in one arm (sum of scattering, transmission, absorption for both mirrors).
3.  $\tilde{x}(f) = \tilde{x}(0)\sqrt{1 + (f/f_k)^2}$ ,  $2\pi f_k = \omega_k = c\mathcal{L}/4L$ .
4. Formulas below do not include recycling, which would improve sensitivity by factor  $R^{-1/2}$ .
5.  $\lambda$  = optical wavelength;  $\tilde{\lambda} = \lambda/2\pi$ ;  $\dot{N} \equiv P/h\nu$ ;  $P$  = Laser power incident on beamsplitter, corrected for photodiode quantum efficiency.

## NOT RECOMBINED (40 m Mark I; included for reference)

The displacement sensitivity of the present (Mark I, not recombined) interferometer is<sup>1</sup>

$$\tilde{x}_{\text{Mk I}}(f) = \frac{L}{\pi\tau_e} \sqrt{\frac{3}{8}} \sqrt{\mathcal{M}} \left[ \frac{\lambda h}{cP} \left( 1 + [f/f_k]^2 \right) \right]^{1/2} \left( \frac{1}{\sqrt{2}} \right).$$

The cavity energy storage time is  $\tau_e = 2L/c\mathcal{L}$ , and the modulation function  $\mathcal{M}$  has the minimum value of 1. Define

$$\tilde{x}_0 = \lambda\mathcal{L} \frac{1}{\sqrt{\dot{N}}} \frac{\sqrt{3}}{4}$$

Then in the limit of low frequency  $f$ , optimum mirror parameters, and optimum (that is vanishingly small) modulation,  $\tilde{x}_{\text{Mk I}}(f) = x_0$ . For  $P = 1$  W and  $\mathcal{L} = 100$  ppm,  $\tilde{x}_0 = 2.2 \cdot 10^{-21}$  m/ $\sqrt{\text{Hz}}$ .

## 1. PROPOSAL ASSUMPTION

The December '89 proposal curves were based on the formula in *300 Years of Gravitation*, Kip's equation 115:

$$S_{h_{\text{FR}}}(f) = \frac{2\hbar c\lambda}{P} \left( \frac{1}{2BL} \right)^2 \left[ 1 + (2\pi BLf/c)^2 \right]$$

For the Fabry-Perot cavity  $B = 4/\mathcal{L}$ , and the low-frequency limit is

$$x_{300} = x_0/\sqrt{6} = 0.4x_0 \tag{1}$$

For knee-frequency much less than 1 kHz, 4 km arms, and  $P = 2$ W, this gives  $x_{300}(f = 1\text{kHz}) = 2.13 \cdot 10^{-18}$  m/ $\sqrt{\text{Hz}}$  (independent of  $\mathcal{L}$ ).

<sup>1</sup>This includes the  $1/\sqrt{2}$  correction to Stan's formula proposed by Harry and checked by Stan.

## 2. APPLIED OPTICS, 1991

The non-recycled fixed mass interferometer at MIT was described by David et. al. in *Applied Optics* 30, No. 22, August 1991, p. 3133. The low-frequency sensitivity is given as

$$\tilde{x} = \left( \frac{\lambda}{8F} \right) \left( \frac{2e}{I_{\text{Max+?}}} \right)^{1/2} \times \mathcal{M}$$

The text states that the minimum value of  $\mathcal{M}$  is 2.  $F = 2\pi/\mathcal{L}$ , and the second term is  $\sqrt{2/\dot{N}}$ ; Therefore

$$x_{\text{FM1}} = \sqrt{\frac{2}{3}} x_0 = 0.8x_0 \quad (2)$$

## 3. MARTIN, JANUARY, 1992

Martin's preliminary calculation of the shot noise in the fixed-mass interferometer he is setting up:

$$S_x^{\frac{1}{2}}(\omega) = \lambda \frac{\sqrt{|E_{\text{DC}}|^2 + 3|E_+|^2} (1 - r_3 r_4)^2}{|E_2||E_+| t_3^2 r_4} \sqrt{1 + (\omega/\omega_c)^2}$$

where  $E_{\text{DC}}$  = amplitude of extraneous light at photodiode (e.g. due to contrast effect), in  $\sqrt{\text{photoelectrons/sec}}$ ,  $E_+ = J_1(\Gamma) D E_L$  amplitude of field due to one modulation sideband at photodiode,  $D$  = transmission factor, typically 0.85,  $E_L$  = laser field,  $E_2 = \sqrt{R} J_0(\Gamma) E_L$  = field due to light circulating inside recycling cavity, travelling from recycling mirror to beamsplitter,  $r_3, t_3$  = cavity input mirror amplitude reflectivity, transmission, and  $r_4$  is for end mirror. The term with all the E's has optimum value  $\sqrt{3}/\sqrt{R} E_L$ , and the term with r's and t's is  $\mathcal{L}/4$ , so

$$x_{\text{FM2}} = x_0 \quad (3)$$

## 4. VINET et. al.

In *Phys. Rev. D* 38(2), p. 433 (1988), Vinet, Meers, Man and Brillet calculated the shot noise for lots of interferometer configurations (all without modulation). Their equation (12) implies (with  $\tau'' = 1/(4\pi f_k)$ )

$$\tilde{h}(0) = \frac{1}{\sqrt{N}} \sqrt{2f_k^2} \frac{1}{c/\lambda}$$

so that

$$x_{\text{VMMB}} = \sqrt{\frac{2}{3}} x_0 = 0.8x_0 \quad (4)$$