

**LASER INTERFEROMETER GRAVITATIONAL WAVE OBSERVATORY
--LIGO--**

California Institute of Technology
Massachusetts Institute of Technology

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**Thermal Noise in a Violin Mode Polarized Along
the Optic Axis for a Single-Loop Suspension**

*This is an internal working note
of the LIGO Laboratory.*

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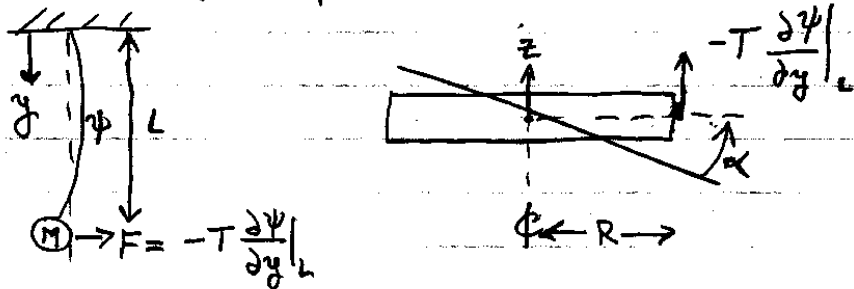
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Thermal Noise in Violin Mode Along Optic Axis for a Single Loop Suspension

FJR
8/27/02



Center of Mass Motion } $-T \frac{\partial \psi}{\partial y} \Big|_L = M \ddot{z}$

Rotation About Center of Mass } $-R T \frac{\partial \psi}{\partial y} \Big|_L = I \ddot{\alpha} = \xi M R^2 \ddot{\alpha}$

Solution for Suspension Wire } $\psi_n(y, t) = A_n \sin(k_n y) e^{i\omega_n t}$
 clamped at $y=0$
 $\frac{\partial \psi_n}{\partial y} \Big|_L = k_n A_n \cos(k_n L) e^{i\omega_n t}$

Boundary Condition } $z_n + \alpha_n R = A_n \sin(k_n L)$

Let $z = z_n e^{i\omega_n t}$
 $\alpha = \alpha_n e^{i\omega_n t}$

$$-T \frac{\partial \psi}{\partial y} \Big|_L = -M \omega_n^2 z_n e^{i\omega_n t}$$

$$-T \frac{\partial \psi}{\partial y} \Big|_L = -\xi M R \omega_n^2 \alpha_n e^{i\omega_n t}$$

$$\Rightarrow \alpha_n = \frac{z_n}{\xi R}$$

①

$$\Rightarrow z_n \left(1 + \frac{1}{\xi}\right) = A_n \sin(k_n L) = z_n \xi$$

$$-T \frac{\partial \Psi}{\partial z} \Big|_L = -M \omega_n^2 \frac{A_n \sin(k_n L)}{J}$$

$$T = \frac{1}{2} Mg$$

$$\frac{1}{2} Mg \cdot k_n A_n \cos(k_n L) = M \omega_n^2 \frac{A_n \sin(k_n L)}{J}$$

$$k_n L = n\pi + \delta k_n \cdot L$$

$$\frac{J}{2} g k_n \cos(n\pi) \approx \cos(n\pi) \cdot \delta k_n \cdot L$$

$$\frac{\delta k_n}{k_n} = \frac{J}{2} \cdot \frac{g}{L} \cdot \frac{1}{\omega_n^2} = \frac{\delta \omega_n}{\omega_n}$$

$$\Rightarrow \left(\frac{\delta \omega_n}{\delta k_n} \right) = \frac{J}{2} \cdot \frac{\omega_p^2}{\omega_n^2} \cdot \left(\frac{\omega_n}{k_n} \right)$$

Frequency shift due to finite-mass termination

$$z_n J = A_n \sin(k_n L) = A_n \cos(n\pi) \delta k_n \cdot L$$

$$= -A_n \cdot \frac{J}{2} \cdot \frac{\omega_p^2}{\omega_n^2} \cdot k_n L$$

Recoil of Mirror from Vibrating Wire

$$\Rightarrow z_n = -A_n \cdot \frac{n\pi}{2} \cdot \left(\frac{\omega_p}{\omega_n} \right)^2$$

RMS Thermal Noise in Wire Violin Mode

$$\frac{1}{2} m_w \omega_n^2 \langle A_n^2 \rangle = \frac{1}{2} kT$$

where m_w is mass of free segment of wire

(2)

$$\Rightarrow \langle A_m^2 \rangle^{1/2} = \left(\frac{kT}{m_w} \right)^{1/2} \frac{1}{\omega_m}$$

$$\langle z_m^2 \rangle^{1/2} = \frac{n\pi}{2\omega_m} \left(\frac{\omega_p}{\omega_m} \right)^2 \left(\frac{kT}{m_w} \right)^{1/2} = \frac{n}{f_m} \left(\frac{f_p}{f_m} \right)^2 \left(\frac{kT}{m_w} \right)^{1/2}$$

For LIGO Arm-Cavity Mirrors

wire density $\rho = 7.9 \text{ g/cm}^3$

$d = 0.012 \text{ in} \cdot 2.54 \text{ cm/in}$

$L = 45 \text{ cm}$

$$m_w = \rho \cdot \frac{\pi}{4} \cdot d^2 \cdot L = 0.26 \text{ g} = 2.6 \times 10^{-4} \text{ kg}$$

for 1st violin mode $m=1$, $f_1 = 344 \text{ Hz}$, $f_p = 0.74 \text{ Hz}$

$$\Rightarrow \langle z_1^2 \rangle^{1/2} = 1.34 \times 10^{-17} \text{ m}$$

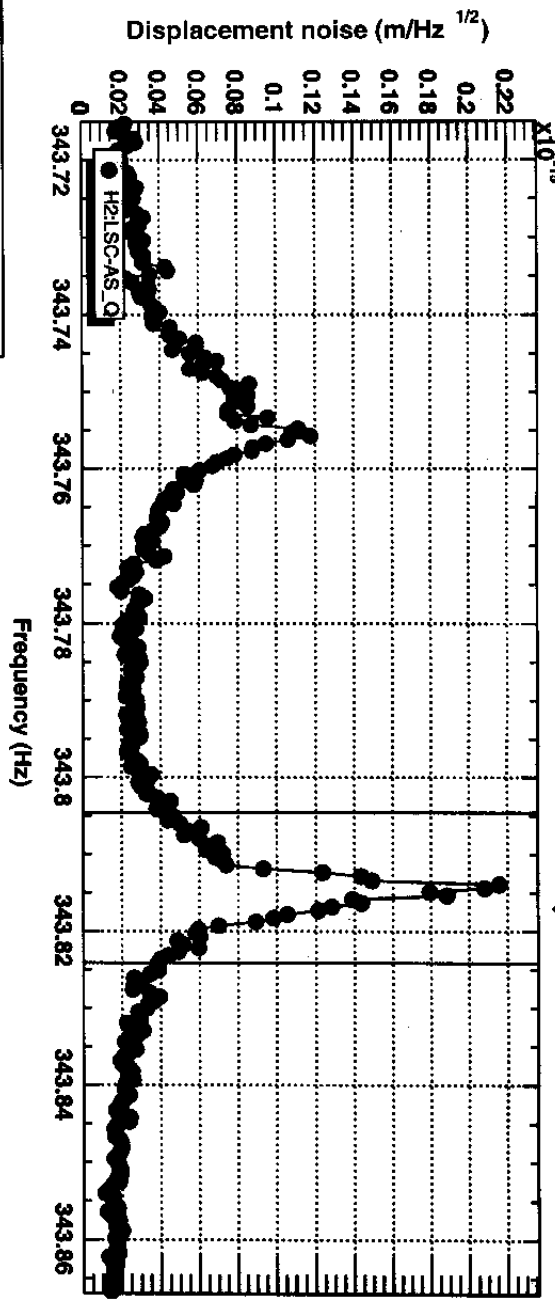
H2: ETMX Observations * (Identification by Raab/Vorvick, LHO e-log, 10/16/01)

f_1 (Hz)	$\langle z_1^2 \rangle^{1/2}$ (m)	Q
343.814	1.4×10^{-17}	$\sim 80,000$
344.051	1.0×10^{-17}	$\sim 70,000$

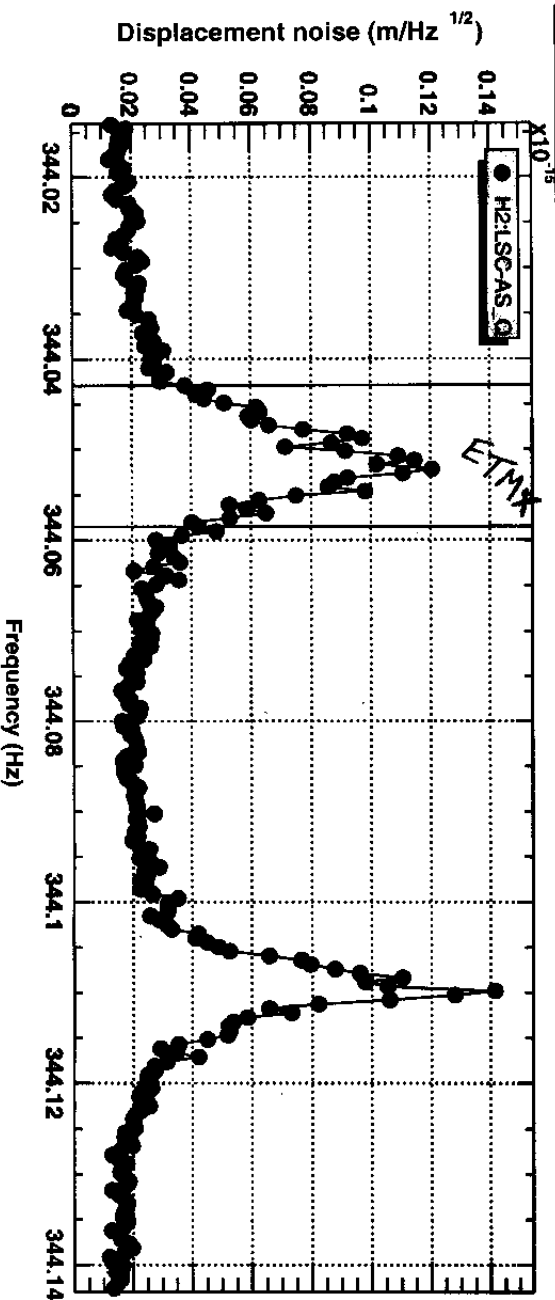
* Calibrated AS-Q spectrum; $T_0 = 26 \text{ Aug } 2002$
02:05:00 UTC; Channel Separation = 0.0005 Hz;

③ BW = 0.000732 Hz; 16 Averages; Refer to
LHO e-log Mon, 26 Aug 02, 17:21:23 PDT.

Calibrated AS-Q spectrum - Mon Aug 26 2002



Power spectrum



T0=26/08/2002 02:05:00

Avg=16

BW=0.000732361