

The Glasgow 10m Interferometer

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LIGO-G100925

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A bit of History + Overview of the facility

Experimental Results

- Control of coupled cavities
- Direct thermal noise measurement
- Diffractive cavity (2nd Littrow)
- Optical Springs with a 100g mirror

Planed Experiments

- Waveguide in a suspended cavity
- Interferometry with Laguerre Gauss modes
- Optical Rigidity: e.g. Optical Bar configuration, Double Optical Spring, Quantum Control, Negative Inertia ...





- The Glasgow 10m Prototype was first set up in 1978.
- Plenty of pioneering work:
 - 1st ever coincidence run of two interferometers (together with Garching 30m)
 - Multiple-stage suspensions
 - > etc ...
- Complete refurbishment of lab after 2000 …
- Since then we have the NEW Glasgow 10m Interferometer.







Overview of the facility (1)

GEO-like Infrastructure:

- Similar vacuum systems
- Triple Stage GEO-suspensions
- Same local control
- Class 100 cleanroom + additional cleanroom tent around every tank.
- Analog and Digital Control
 - dSPACE + aLIGO CDS



 Originally planned to accommodate a full aLIGO configuration (FP-MI with Dual-Recycling)





Overview of the facility (2)

- Maximise the potential of the Glasgow 10m by carrying out several strands of experiments in parallel:
- Direct measurement of thermal noise
- 3-mirror coupled cavity systems:
 - Control strategies
 - Radiation pressure experiments
 - Frequency reference for TN experiment
- Diffractive interferometry
 - Grating as cavity incoupler
 - Waveguide mirror





The Glasgow 10m Prototype Interferometer

- Ideal test bed for advanced interferometry concepts.
- Fast turn around for rapid, small-scale tests
- Timely validation of various innovative technologies
- Excellent training for students and Postdocs







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• Aim: Develop Control scheme for a linear 3-mirror cavity:

- Providing independent error-signals for the 2 degrees of freedom
- Allowing us to easily detune one of the cavities, without disturbing the other (needed for optical spring experiments).



Method makes use of 3 RF modulation frequencies:

- > PM sidebands @ 18MHz used to derive AC length-sensing signal.
- PM sidebands @ 10MHz, and AM sidebands @ 14.525MHz used to derive recycling cavity length-sensing signal.



- Concept works very nicely.
- More details on the experimental results can be found in:
 - S.Huttner et al: "Novel sensing and control schemes for a three-mirror coupled cavity". CQG, 2007, 24, 3825
 - B. Barr et al: "Optical modulation techniques for length sensing and control of optical cavities" Appl. Opt., OSA, 2007, 46
 - S. Huttner et al: "Techniques in the optimization of length sensing and control systems for a three-mirror coupled cavity", CQG, 2008, 25, 235003
- Concept can also be useful locking Khalili cavities (as ETM), for instance in AEI-10m sub-SQL experiment ...





- Aim: To directly measure coating Brownian noise at frequencies around 100 to 200Hz.
- Short (10cm) high-finesse (about 5000) measurement cavity.
- 10m cavity serves as frequency reference.
- Very sophisticated test mass suspensions (3 stages, last stage is monolithic)
- Target sensitivity: about 3x10⁻¹⁸m/sqrt(Hz)



More details on the thermal noise experiment: J.Taylor, PhD-thesis





- Current status: Haven't reached target sensitivity => More noise-hunting is needed. Potential noise candidates:
 - Electronic noise in frequency-stabilisation loop
 - Cross-coupling between 10cm and 10m cavities
- What could we do with the experiment once we reached target sensitivity?
 - Measurement of coatings at the frequencies within the detection band of the large GWDs.
 - Measurement of thermal noise of a hydroxid bond.
 - Other fun stuff such as LG-modes or Waveguide coatings....







- Motivation: Gratings can be used to replace transmissive optics => can decrease thermal problems and some thermal noise contributions.
- First demonstration of a suspended cavity using a 3-port grating as cavity incoupler (2nd order Littrow).
- Collaboration with Universities of Hannover, Jena and Birmingham.
- Using an grating etched into a fused silica substrate then overcoated with multi-layers of Ta₂O₅ and SiO₂.
- Achieved Cavity Finesse of about 1100.
- Detailed characterisation of signal ports completed.







Diffractive coupled cavity





• Two more examples of interesting investigations:







- Experimental confirmation of phase noise from side-motion of the grating.
- Prediction that in contrast to a normal mirror, sidemotion of a grating, introduces additional phase noise. (Freise et al, New J Phys, 2007, 9, 433-+)



More details: B.Barr et al, "Translational, rotational and vibrational coupling into phase in diffractivelycoupled optical cavities", soon in LSC-review ...

S.Hild



- Aim: Measure optical spring of a 100g mirror.
- Motivation: precursor for other optical rigidity experiments (also at AEI-10m sub-SQL-interferometer).
- We replaced the endmirror (2.7kg) of the 3-mirror cavity by a very light mirror (100g) in order to enhance the radiation pressure effects.







Optical spring with a 100g mirror

- Triple-stage suspension + reaction chain.
- Much of this design serves as reference for the suspension for the AEI-10m







55um radius wire





Some preliminary results



Illustration of OLTF

Measured Data (CLTF)



OVERVIEW

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Waveguide coating

Motivation: Waveguide coatings might provide very low thermal noise.

- A.Bunkowski et al: CQG, 2006, 23, 7297-7303
- F.Brückner et al: Opt. Lett., OSA, 2008, 33, 264-266
- F.Brückner et al: Optics Express, 2008, 17, 163-
- Aim: Demonstrate suspended cavity with waveguide mirror
- Collaboration with AEI-Hannover and Jena University. Main actors: D.Friedrich (Hannover) and F.Brueckner (Jena).
- Jena produced a batch of waveguide mirrors (1064nm) for the Glasgow 10m Prototype.
- Reflectivity of 99.5% (preliminary result)
- Installation will start in 10 days.







Waveguide coating

Install the waveguide mirrors as ITM. Planned experiments:

- Investigate control signals
- Demonstrate the absence of sidemotion noise
- Investigate homogeneity and and scattered light







LG-modes

- Motivation: LG modes can provide a reduction in thermal noise.
- Simulations and experiment of LGinterferometry seem promissing.
 - S.Chelkowski et al: "Prospects of higher-order Laguerre-Gauss modes in future gravitational wave detectors", *PRD*, **2009**, *7*9, 122002
 - P.Fulda et al: "Experimental demonstration of higher-order Laguerre-Gauss mode interferometry", *PRD*, 2010, 82, 012002
- Time for testing in an environment similar to a GW detector.
- Collaboration with University of Birmingham.
- Main aims:
 - Study of control signals
 - Investigate effects from mode-degeneracy

Examples of higher order LG modes produced with a spatial phase modulator





Optical rigidity

- Will continue on investigations on interferometer concepts based on optical rigidity.
- One potential experiment would be to go back to a heavy ETM and install very light ITM + a local readout for the ITM. => Close to an optical bar / optical lever configuration.
- Other routes currently considered and theoretically investigated are:
 - Double optical springs
 - Negative inertia
- Collaboration with Moscow (S.Vyatchanin and A.Rakhubovsky)





- The Glasgow 10m Prototype is a fully operational Multi-Experiment facility.
- Strong link with AEI-10m sub-SQL interferometer.
- Lots of interesting interferometry concepts under investigation.
- Lively exchange of PhD-students. Everybody is welcome to join ...