

# Discharging Fused Silica Test Masses with Ultraviolet Light

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Test mass charging is a potentially limiting noise source for gravitational-wave interferometers such as LIGO. We used a Kelvin probe and tunable light source to study the effectiveness of UV illumination for discharging test masses that are positively charged through contact electrification with viton. We found that the discharge rate is a linear function of the charge magnitude and the illumination intensity. The discharge rate is also a function of UV wavelength, with optimal discharging at 215 nm. A deposited energy on the order of 0.1 J/cm<sup>2</sup> is required to discharge the test mass by 90%.

## 1. Introduction

Charging is a potentially limiting noise source for both gravitational-wave interferometers<sup>1-3</sup> and other precision measurements of gravitational effects<sup>4,5</sup>. At the Laser Interferometer Gravitational-Wave Observatory (LIGO)<sup>6</sup>, charge may build up on the surface of suspended fused silica test masses. Charging can occur due to friction between dust and the test mass surface (particularly when the system is being pumped to vacuum), the deposition of excess electrons from a cosmic ray striking the beam pipe<sup>7</sup>, or from contact between the test mass and other materials such as viton-tipped earthquake stops, designed to protect the test mass by limiting its range of motion. Measurements of fused silica optics in vacuum have shown a substantial charging rate of  $\sim 10^5 e^-/\text{cm}^2/\text{month}$ <sup>8</sup>.

There are several potential noise contributions from charging<sup>9</sup>. Surface charge would generate electric fields that induce a force between a test mass and its metallic suspension frame, displacing the mass beyond the range of its positioning magnets. Sudden changes in charge magnitude or position would discontinuously change this force, moving the test mass in a way that mimics a gravitational-wave burst signal. Moving charges would generate fluctuating electric fields that could displace the test mass at frequencies in the interferometer's sensitive band. And static charges could attract dust to the surface of a test mass, reducing optical reflectance and increasing absorption, making thermal compensation more difficult<sup>10</sup>.

One possible solution for test mass charging is to discharge through UV illumination. This technique was used for Gravity Probe B, in which the charged surface and an adjacent "charge control electrode" were illuminated with UV light in order to discharge electrons by the photoelectric effect<sup>11</sup>. The net direction of charge flow could then be controlled by adjusting the voltage of the control electrode – a positively-charged surface could be discharged by receiving electrons liberated from the electrode. Researchers at the University of Glasgow found that UV radiation from an ion pump was causing negative charging of an optic by liberating electrons from the walls of a vacuum chamber; illuminating the optic with a UV lamp reversed the effect<sup>2</sup>. Measurements at the GEO 600 gravitational-wave observatory have also shown a substantial reduction in positive charge on a test mass by shining light on a control cathode<sup>12</sup>.

A concern with this technique is that exposure to UV light over time may damage the test mass reflective coatings; experimenters at Stanford University are currently measuring the

absorption of a fused silica test mass exposed to a UV LED to quantify this effect<sup>13</sup>. The goal of this project was to measure discharging rate as a function of UV intensity and wavelength. This will allow a determination of the minimum power level required for discharging, which can then be compared to the Stanford absorption measurements to see if UV illumination is a viable discharging technique. Fused silica optics with titania-doped tantala/silica coatings, the baseline coating planned for use in Advanced LIGO<sup>14</sup>, were charged through contact with viton and studied in vacuum, all to maintain fidelity with charging in a LIGO-like environment.

## 2. Experimental Setup

The charge measurements were made with a capacitive device called a Kelvin probe. A charge layer on a sample induces opposite charge to flow to the surface of the probe element. Modulating the capacitance creates an alternating current signal proportional to the potential difference between the probe and sample, which depends on the magnitude of charge on the sample<sup>15</sup>. The modulation can be achieved by vibrating the probe tip with a piezoelectric transducer (PZT) or voice coil, or by periodically occluding the sample with an optical chopper<sup>16</sup>. The AC signal is read with a lock-in amplifier set to the modulation frequency; the sign of the charge can be determined from the phase difference between the modulation and the probe signal.

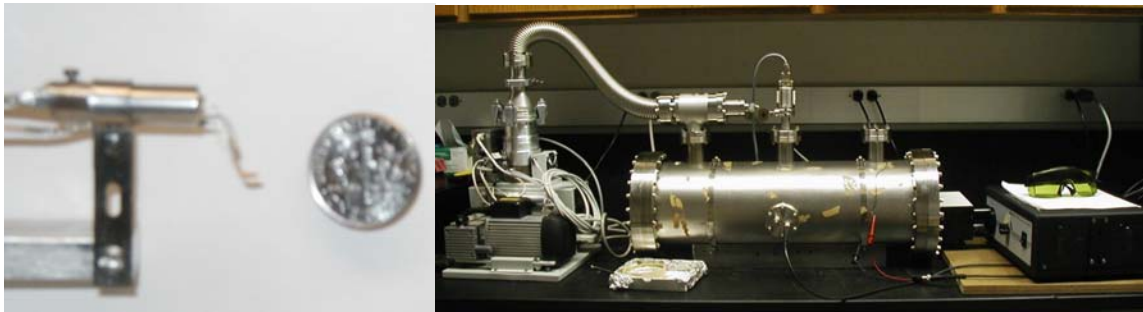


Figure 1: (Left) Besocke Kelvin probe. (Right) Vacuum system and xenon lamp with monochromator.

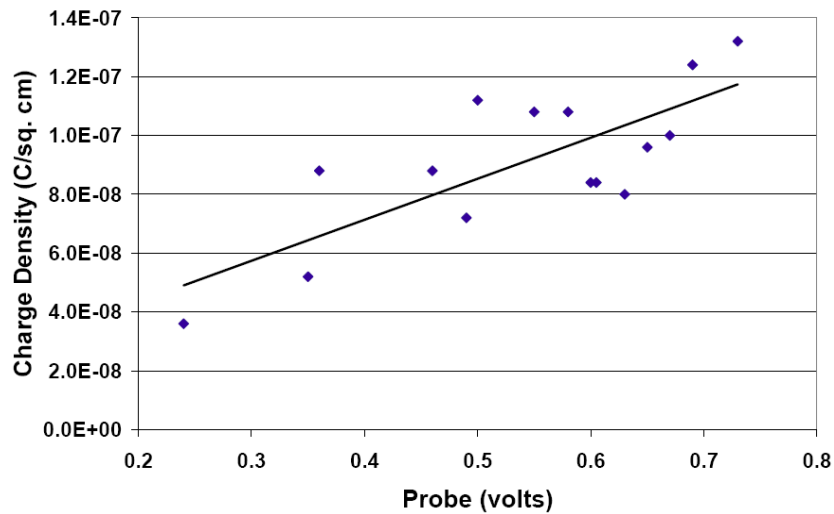


Figure 2: Calibration of Besocke Kelvin probe with Alphalab surface DC voltmeter

We used the Kelvin Probe S from Besocke delta phi GmbH, shown at left in Figure 1. The probe has a 2.5 mm diameter electrode which is vibrated vertically by a PZT, and is sensitive to 0.1 mV. We calibrated the probe by rubbing a large acrylic sample with felt and measuring the resulting charge with both the probe and an Alphalab surface DC voltmeter, the latter of which could be converted to a surface charge density. Figure 2 shows the results of many such measurements. The scatter is likely a result of the voltmeter averaging the charge over a larger area than the probe. A linear fit gives a calibration of  $(1.4 \pm 0.3) \times 10^{-7} \text{ C/m}^2/\text{volt} = (8 \pm 2) \times 10^7 \text{ e}^-/\text{cm}^2/\text{volt}$ . Thus the probe sensitivity of 0.1 mV corresponds to a minimum charge resolution of  $(8 \pm 2) \times 10^3 \text{ e}^-/\text{cm}^2$ .

The UV light source was a 175W Xenon based lamp from Spectral Products. At the output is a monochromator with a 2400 line/mm grating, allowing wavelength selection from 190-680 nm. Figure 3 shows measurements of the lamp intensity at UV wavelengths with 0.6 mm wide apertures at the input and output of the monochromator. The spot size at the charged sample is roughly  $4 \text{ cm}^2$  in area, so we are working with intensities of approximately  $0.5 \mu\text{W}/\text{cm}^2$  over this range of wavelengths.

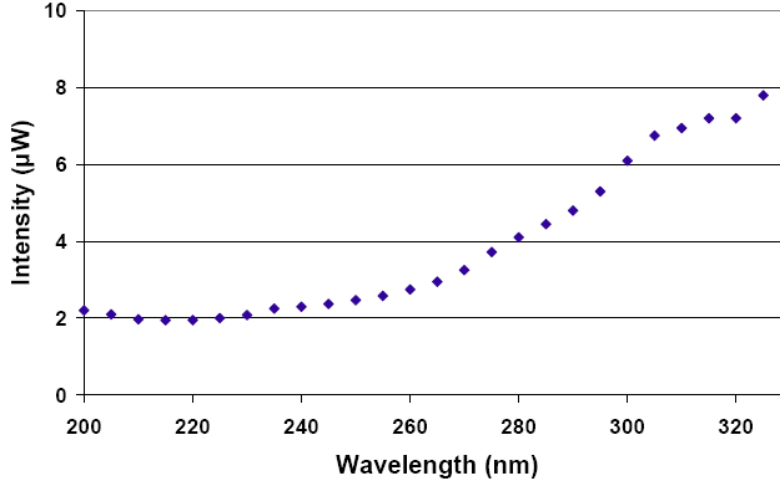


Figure 3: Measured intensity of Xenon based light source at UV wavelengths

The vacuum chamber is shown at right in Figure 1. A Pfeiffer Vacuum TSH-071E turbomolecular drag pumping station allows pressures down to  $4 \times 10^{-7}$  torr. At one end of the chamber, facing our UV light source, is an MDC Vacuum quartz viewport with a transmission of nearly 95% at wavelengths down to 200 nm. The Kelvin probe power and signal wires run through separate feedthroughs at opposite sides of the chamber to minimize noise pickup.

### 3. Discharging Measurements

The samples were two fused silica substrates, 7.6 cm in diameter and 0.25 cm thick. One of the samples had a thirty layer, quarter wave stack of titania-doped tantala/silica with silica as the top layer as a high reflection coating on one side. The samples were mounted at a 15 degree angle to the horizontal, facing the UV light source, and were held in place with two viton-tipped clamps. The Kelvin probe is mounted on a post on the opposite side of the sample from the light source, with the probe element 1 mm above the sample. The samples were charged through

contact with a viton O-ring, then brought down to vacuum and illuminated with UV light while the charge level was measured with the Kelvin probe. Note that the Kelvin probe measurements indicated that the sample was positively charged after contact with the viton; this will be further discussed at the end of this section.

Figure 4 shows the rate of discharge (given by the change in the Kelvin probe reading per second) versus total charge, measured over 15 hours with a single sample. There is a clear linear relationship between the discharge rate and total charge. This allows us to compare future discharge measurements made at different starting charge levels, by correcting all measurements to a standard initial charge. It also implies that UV illumination may be adequate for dealing with uneven distributions of charge across the test mass surface. Since the discharge rate scales

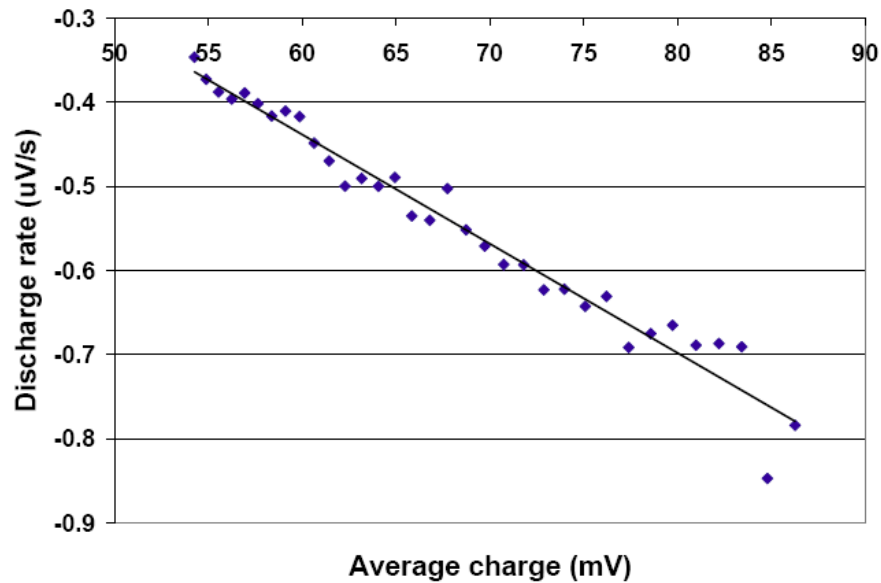


Figure 4: Relationship between discharge rate and charge on sample

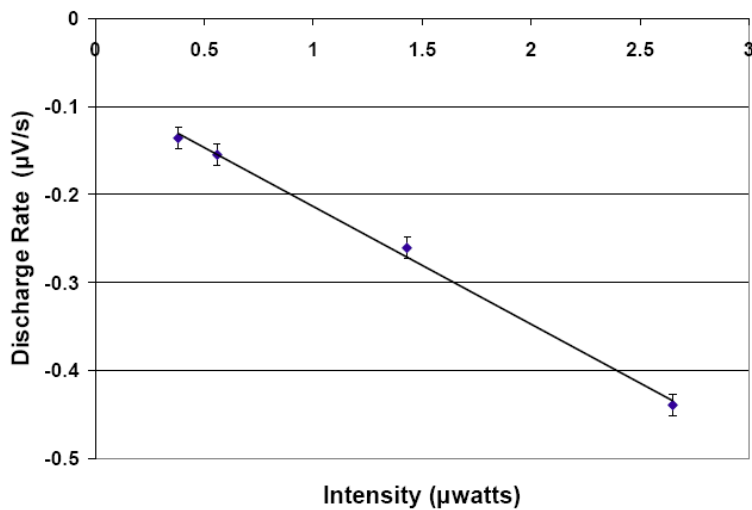
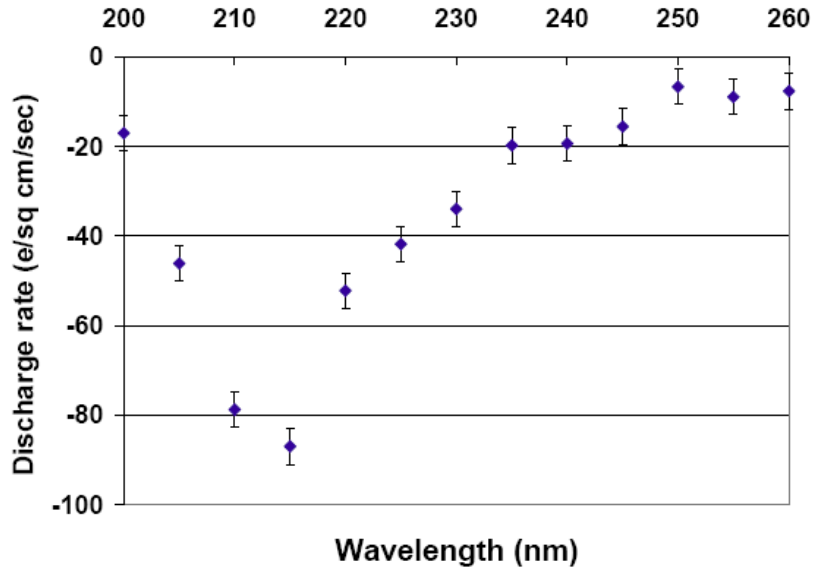


Figure 5: Relationship between discharge rate and UV illumination intensity



**Figure 6: Discharge rate versus UV illumination wavelength for uncoated fused silica sample**

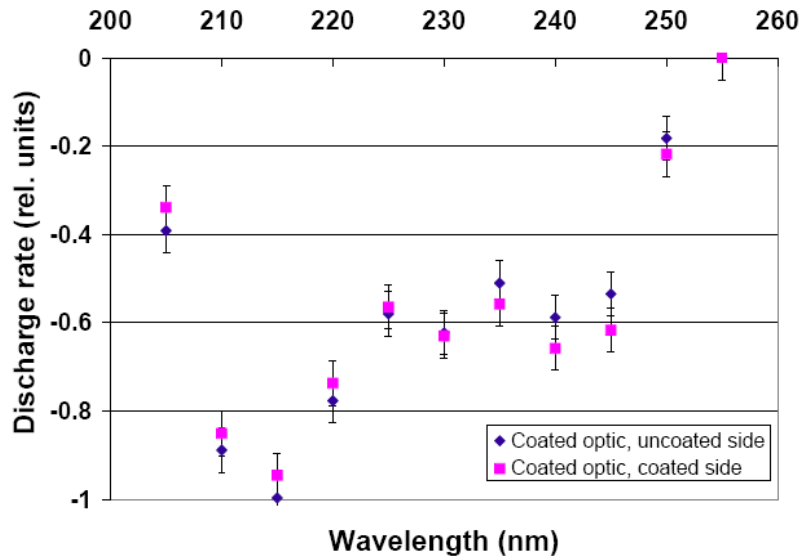
with charge density, areas of high charge density will discharge more rapidly than areas of low charge density, resulting in a more even distribution over time.

Similarly, Figure 5 shows the discharge rates for different UV light intensities at the same wavelength, accomplished by inserting different aperture widths at the monochromator output. This allows us to compensate for the varying intensities of the light source at different wavelengths, by correcting all measurements to a standard intensity.

A representative measurement of the uncoated sample discharge rate is shown in Figure 6, normalized to an illumination intensity of  $0.5 \mu\text{W}/\text{cm}^2$  and a charge magnitude of  $8 \times 10^6 e^-/\text{cm}^2$  (equal to a probe reading of 100 mV). The 200 nm discharge rate is rechecked every fourth measurement, to ensure that there is no consistent trend in our results that would indicate systematic error. The statistical fluctuation in the measurements at 200 nm are also used to determine error bars. When the sample is brought back to atmospheric pressure, recharged, and remeasured, the shape of the curve versus wavelength remains the same, but the absolute discharge rate can vary by as much as a factor of two. Since the spot size, illumination intensity, and charge magnitude do not vary by this much, the likely reason is that the area on the sample taken up by the charge distribution varies from one charging event to the next, resulting in a variation in the amount of light hitting a charged part of the sample surface.

Figure 6 allows an order-of-magnitude estimate of the UV light energy necessary to discharge the sample at the optimal wavelength of 215 nm. In one second the charge magnitude is reduced by a fraction of approximately  $80 / (8 \times 10^6) = 10^{-5}$  of its original value. The linear relationship between charge magnitude and discharge rate implies an exponential discharge over time, which implies a time constant  $\tau = 10^5 \text{ s}$ . This corresponds to a time of  $2.3 \times 10^5$  seconds for 90% discharging, and multiplying by the illumination intensity gives a total incident energy of  $0.11 \text{ J}/\text{cm}^2$  for 90% discharging.

Figure 7 shows the discharge rate versus wavelength for both the coated and uncoated side of the tantala/silica coated sample. Both curves have been normalized to the same discharge rate at 230 nm to allow comparison of the shapes. The curves match to within the measurement uncertainty, and both show a greater response between 230 nm and 250 nm than the uncoated



**Figure 7: Discharge rate versus UV illumination wavelength for the coated and uncoated sides of a single fused silica sample**

sample; the reason for this is not yet understood. The optimal discharging wavelength is again found to be 215 nm.

A reduction in positive charge implies that either positive ions, possibly deposited from contact with viton, are being removed from the sample surface, or electrons are being liberated from nearby material and are neutralizing the sample. Electrons could originate from the probe or from the surrounding vacuum chamber. We have found that discharging only occurs when the UV light is aimed directly at the optic, which eliminates the second possibility. We see similar discharging rates when the light is aimed slightly to the side of the probe element, but it is unclear how much stray light still reaches the probe; a better test would require motor control over the probe position, allowing us to move it aside, illuminate the optic, and bring it back into position without making vacuum. But our favored interpretation is that the light is liberating positive ions from the sample surface. The transfer of positive ions during contact electrification has been seen in other experiments<sup>17</sup>, and the peak discharge rate at 215 nm (or a photon energy of 5.8 eV) corresponds to a previously measured absorption peak for fused silica<sup>18</sup>. But the mechanism by which positive ions are removed from the surface is unknown.

#### 4. Conclusions

We have demonstrated that fused silica test masses acquire positive charge through contact electrification with viton. Illumination with UV light can remove this positive charging, with peak response at a wavelength of 215 nm. At this wavelength, on the order of  $0.1 \text{ J/cm}^2$  will result in a 90% discharge. We also found that the discharge rate varies linearly with charge density, suggesting that over time UV illumination may be able to damp out uneven distributions of charge.

Given that different mechanisms can result in either positive or negative charging of a test mass, it appears that the Gravity Probe B “control electrode” strategy described in the introduction may also be an appropriate plan for LIGO. Future use of the Kelvin probe setup

will include testing such a strategy, as well as adding motion control for the probe in order to map the charge distribution on samples and how it is affected by UV illumination.

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## References

- <sup>1</sup> R. Weiss, *Note on Electrostatics in the LIGO Suspensions*, LIGO technical note, (1996). <http://www.ligo.caltech.edu/docs/T/T960137-00.pdf>.
- <sup>2</sup> S. Rowan, S. M. Twyford, R. Hutchins, and J. Hough, *Investigations into the effects of electrostatic charge on the  $Q$  factor of a prototype fused silica suspension for use in gravitational wave detectors*, CQG **14**, 1537-1541 (1997).
- <sup>3</sup> P. L. Bender *et al.*, *LISA Pre-Phase A Report; Second Edition*, MPQ Report No. 233 (1998).
- <sup>4</sup> S. Buchman, C. W. F. Everitt, B. Parkinson, J. P. Turneare, R. Brumley, D. Gill, G. M. Keiser, and Y. Xiao, *The Gravity Probe B Relativity Mission*, Adv. Space Res. **25**, 1177-1180 (2000).
- <sup>5</sup> E. G. Adelberger, B. R. Heckel, and A. E. Nelson, *Tests of the Gravitational Inverse-Square Law*, Annu. Rev. Nucl. Part. Sci. **53**, 77-121 (2003).
- <sup>6</sup> S. J. Waldman, *Status of LIGO at the start of the fifth science run*, Class. Quantum Grav. **23**, S653 (2006).
- <sup>7</sup> V. B. Braginsky, O. G. Ryazhskaya, and S. P. Vyatchanin, *Notes about noise in gravitational wave antennas created by cosmic rays*, Phys. Lett. **A350**, 1-4 (2006).
- <sup>8</sup> V. P. Mitrofanov, L. G. Prokhorov, and K. V. Tokmakov, *Variation of electric charge on prototype of fused silica test mass of gravitational wave antenna*, Phys. Lett. **A300**, 370-374 (2002).
- <sup>9</sup> D. Ugolini, R. Amin, G. Harry, J. Hough, I. Martin, V. Mitrofanov, S. Reid, S. Rowan, and K-X. Sun, *Charging Issues in LIGO*, Proceedings of the 30<sup>th</sup> International Cosmic Ray Conference (2007).
- <sup>10</sup> R. Lawrence, D. Ottaway, P. Fritschel, and M. Zucker, *Active correction of thermal lensing through external radiative thermal actuation*, Opt. Lett. **29**, 2635-2637 (2004).
- <sup>11</sup> S. Buchman, T. Quinn, G. M. Keiser, D. Gill, and T. J. Sumner, *Charge measurement and control for the Gravity Probe B gyroscopes*, Rev. Sci. Instrum. **66**, 120-129 (1995).
- <sup>12</sup> M. Hewitson, K. Danzmann, H. Grote, S. Hild, J. Hough, H. Luck, S. Rowan, J.R. Smith, K.A. Strain, and B. Willke, *Charge measurement and mitigation for the main test masses of the GEO 600 gravitational wave observatory*, Class. Quantum Grav. **24**, 6379-6391 (2007).
- <sup>13</sup> K-X Sun, N. Leindecker, S. Higuchi, A. Markosyan, R. Route, M. Hennessey, J. Perales, S. Buchman, M. Fejer, R. Byer, and G. Harry, *Update of LIGO Test Mass Charging Mitigation Activities at Stanford*, LIGO document G070208-00-Z.
- <sup>14</sup> G. M. Harry, M. R. Abernathy, A. E. Becerra Toledo, H. Armandula, E. Black, K. Dooley, C. Nwabugwu, A. Villar, D. R. M. Crooks, G. Cagnoli, J. Hough, I. MacLaren, P. Murray, S. Reid, S. Rowan, P. H. Sneddon, M. M. Fejer, R. Route, S. D. Penn, P. Ganau, J.-M. Mackowski, C.

Michel, L. Pinard, and A. Remillieux, *Titania-doped tantala/silica coatings for gravitational wave detection*, CQG **24**, 405-415 (2006).

<sup>15</sup> L. Kronik and Y. Shapira, *Surface photovoltage phenomena: Theory, experiment, and applications*, Surf. Sci. Rep. **37**, 1-206 (1999).

<sup>16</sup> D. Ugolini, R. McKinney, and G. M. Harry, *Developing an optical chopper-modulated capacitive probe for measuring surface charge*, Rev. Sci. Instrum. **78**, 046102 (2007).

<sup>17</sup> J. A. Wiles, B. A. Grzybowski, A. Winkleman, and G. M. Whitesides, *A Tool for Studying Contact Electrification in Systems Comprising Metals and Insulating Polymers*, Anal. Chem. **2003**, 75, 4859-4867 (2003).

<sup>18</sup> F. Messina and M. Cannas, *In situ observation of the generation and annealing kinetic of E' centres induced in amorphous SiO<sub>2</sub> by 4.7 eV laser irradiation*, J. Phys.: Condens. Matter **17**, 3837-3842 (2005).