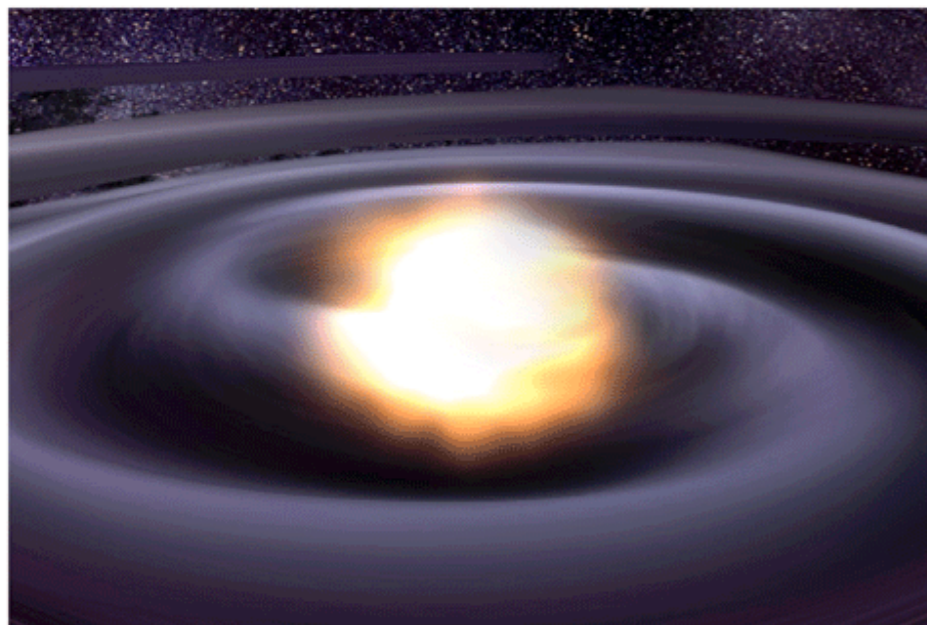
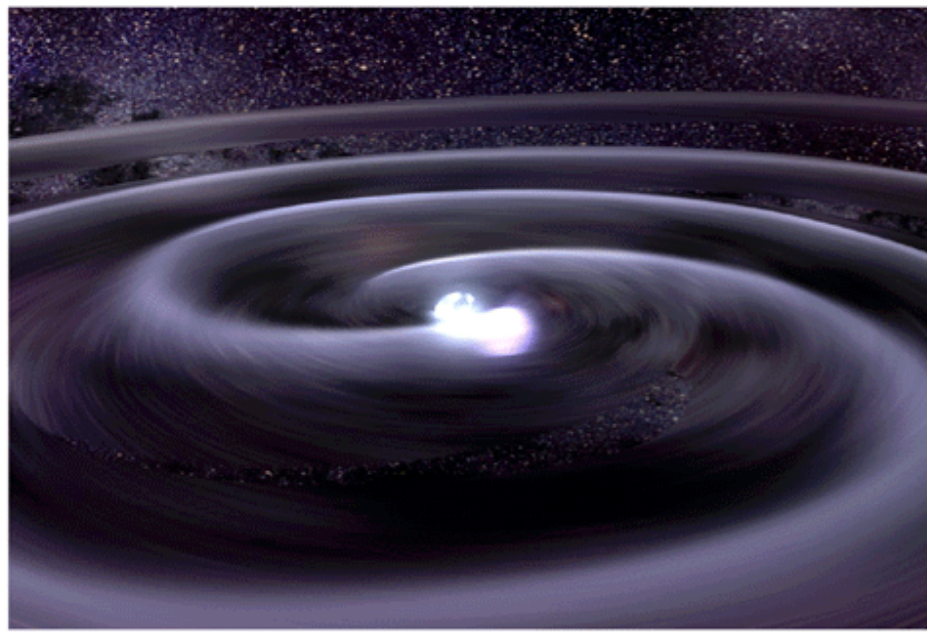
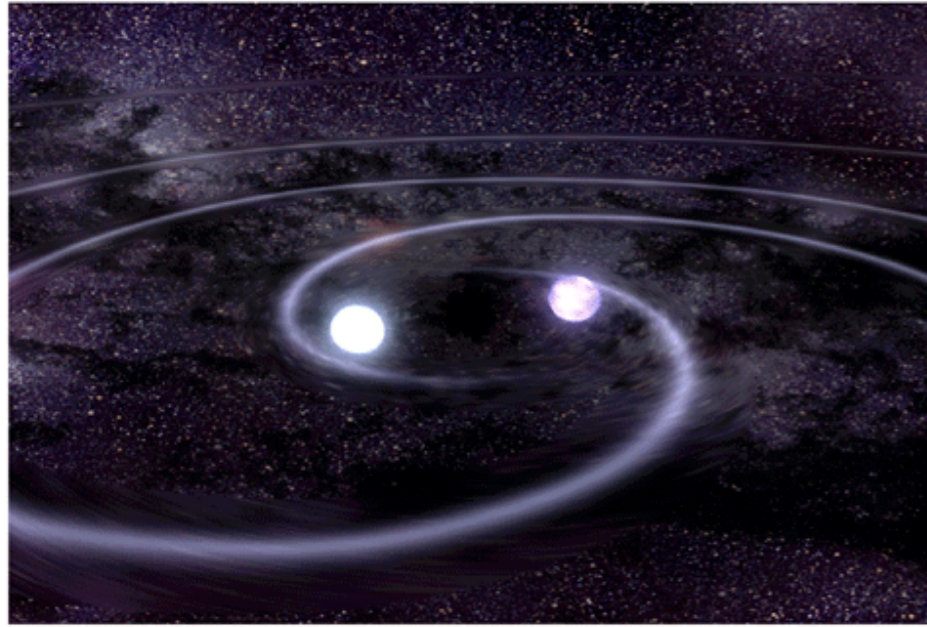
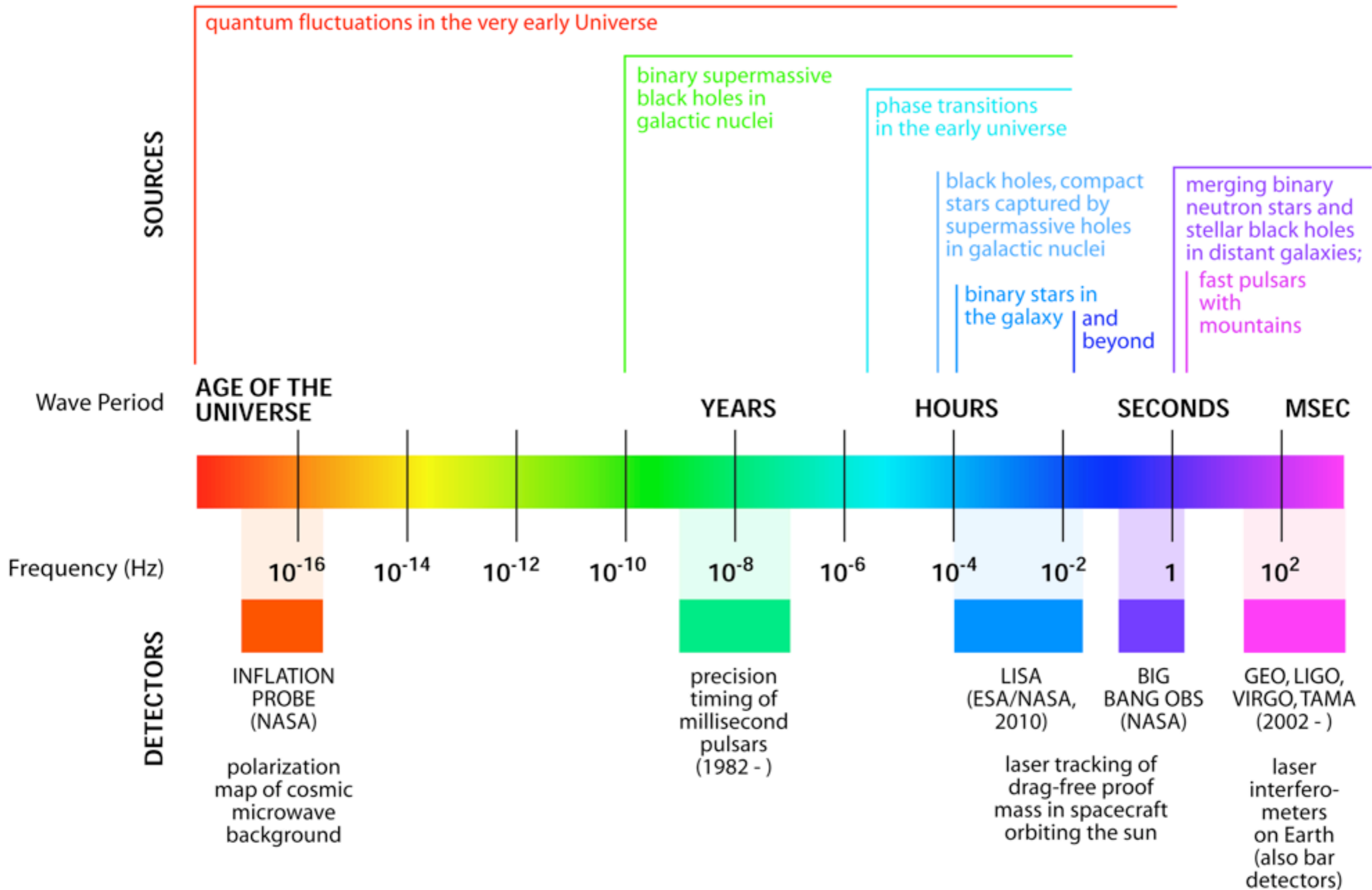
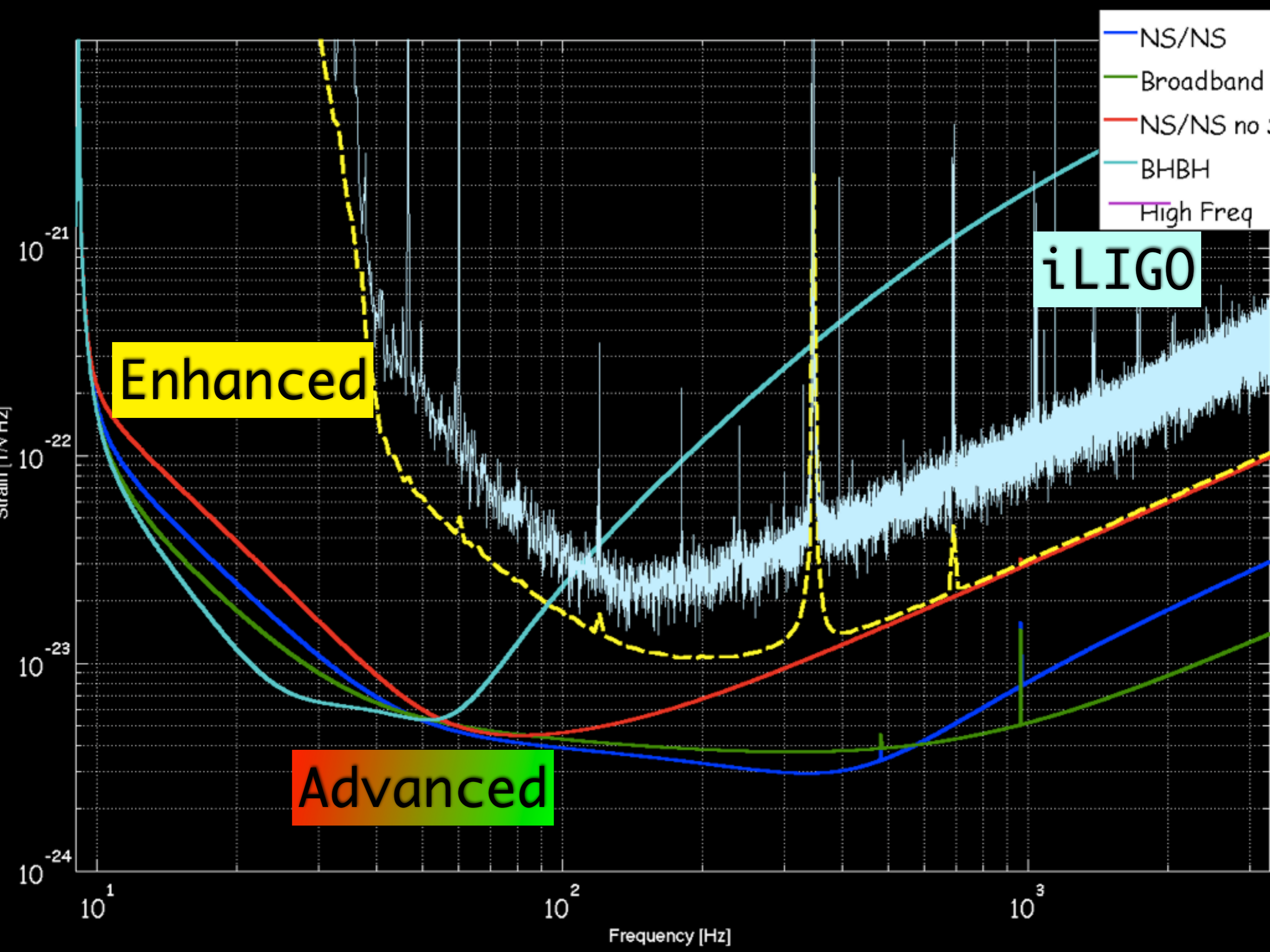


Near Term Upgrades

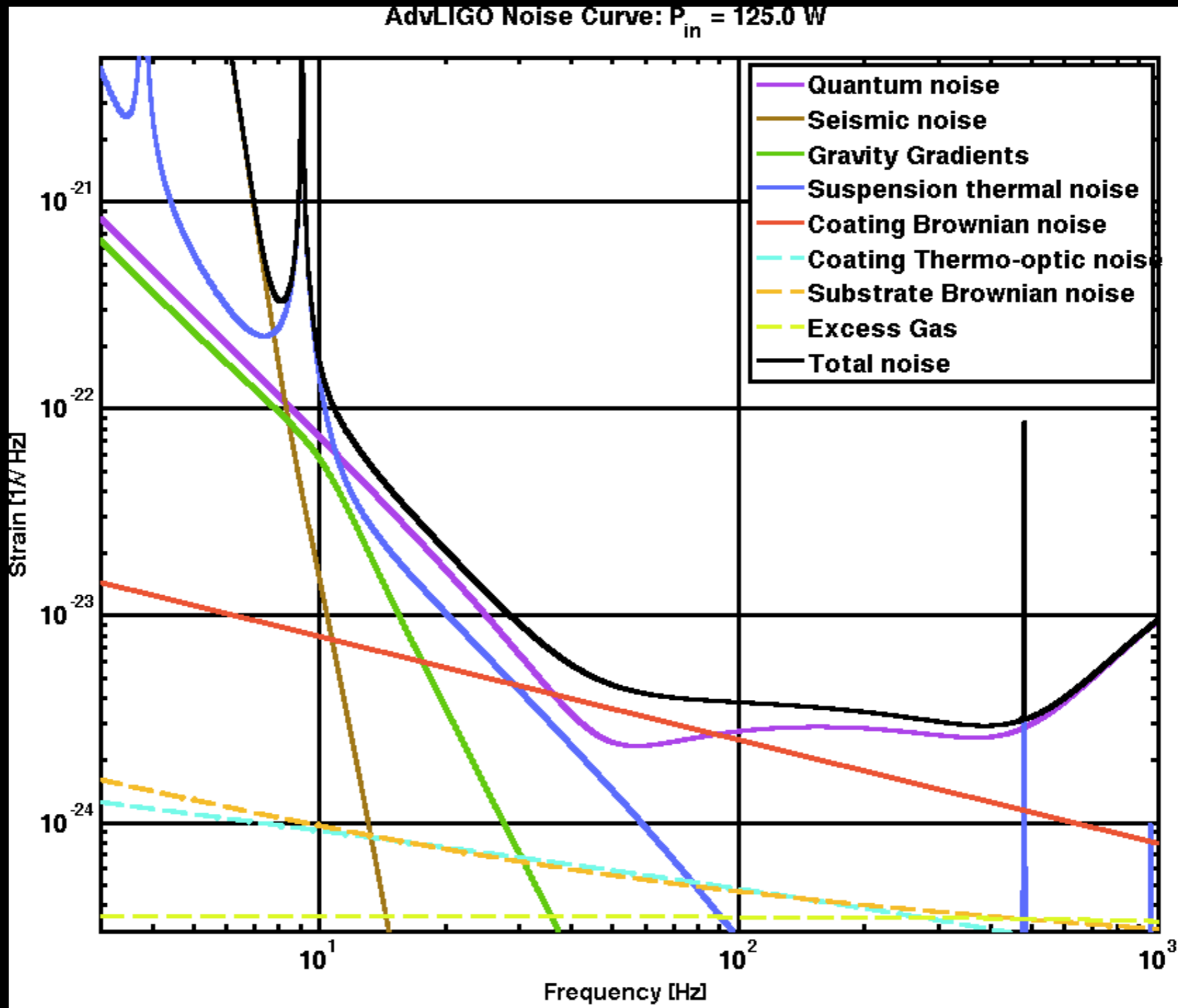


THE GRAVITATIONAL WAVE SPECTRUM

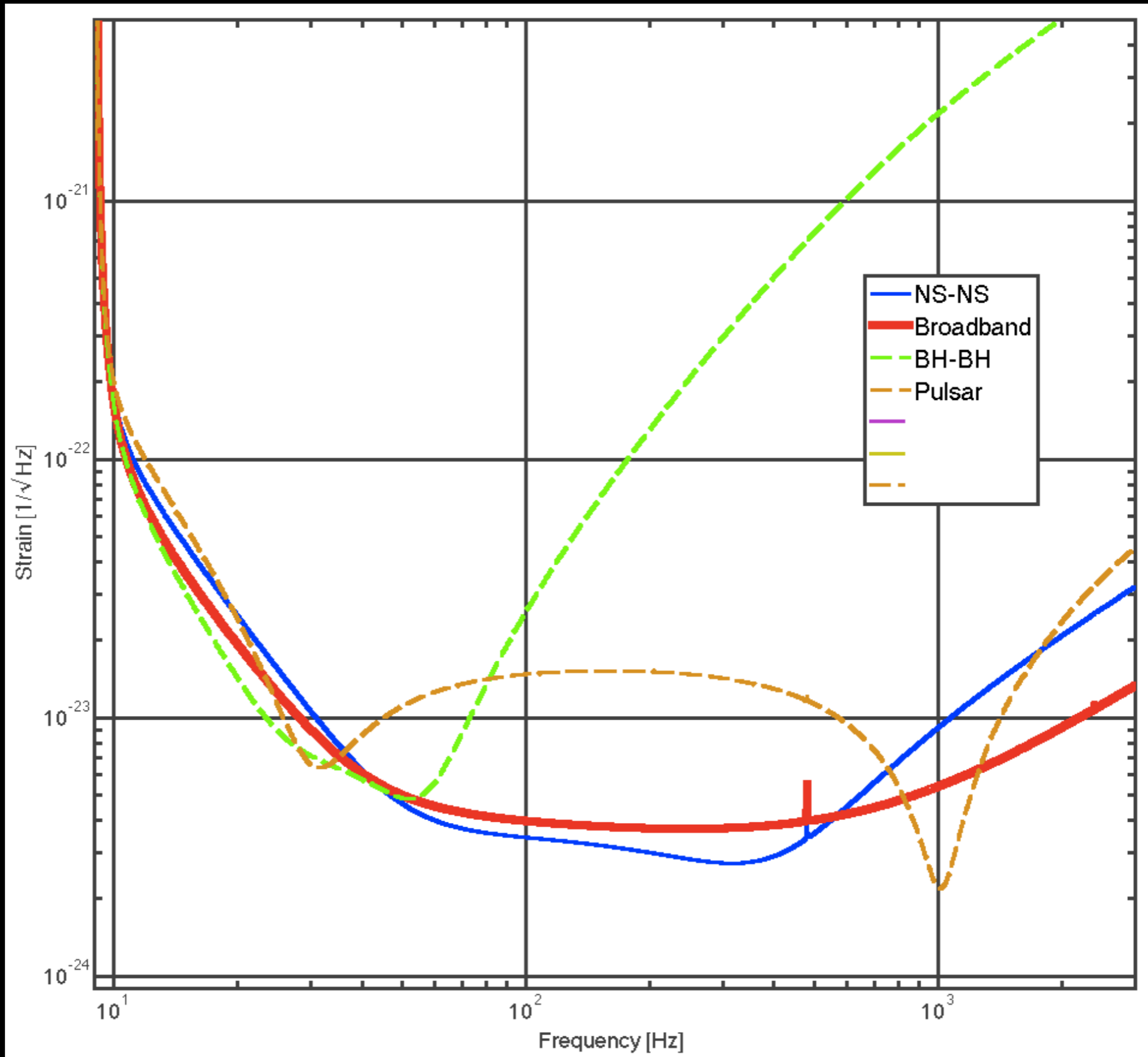




Noise Limits



Tunings



Low F v. High F

○ Seismic Noise

○ Newtonian Noise

○ Angular Controls

○ Low Noise Coatings

○ Suspension Thermal

High Circulating Power

Radiation Pressure

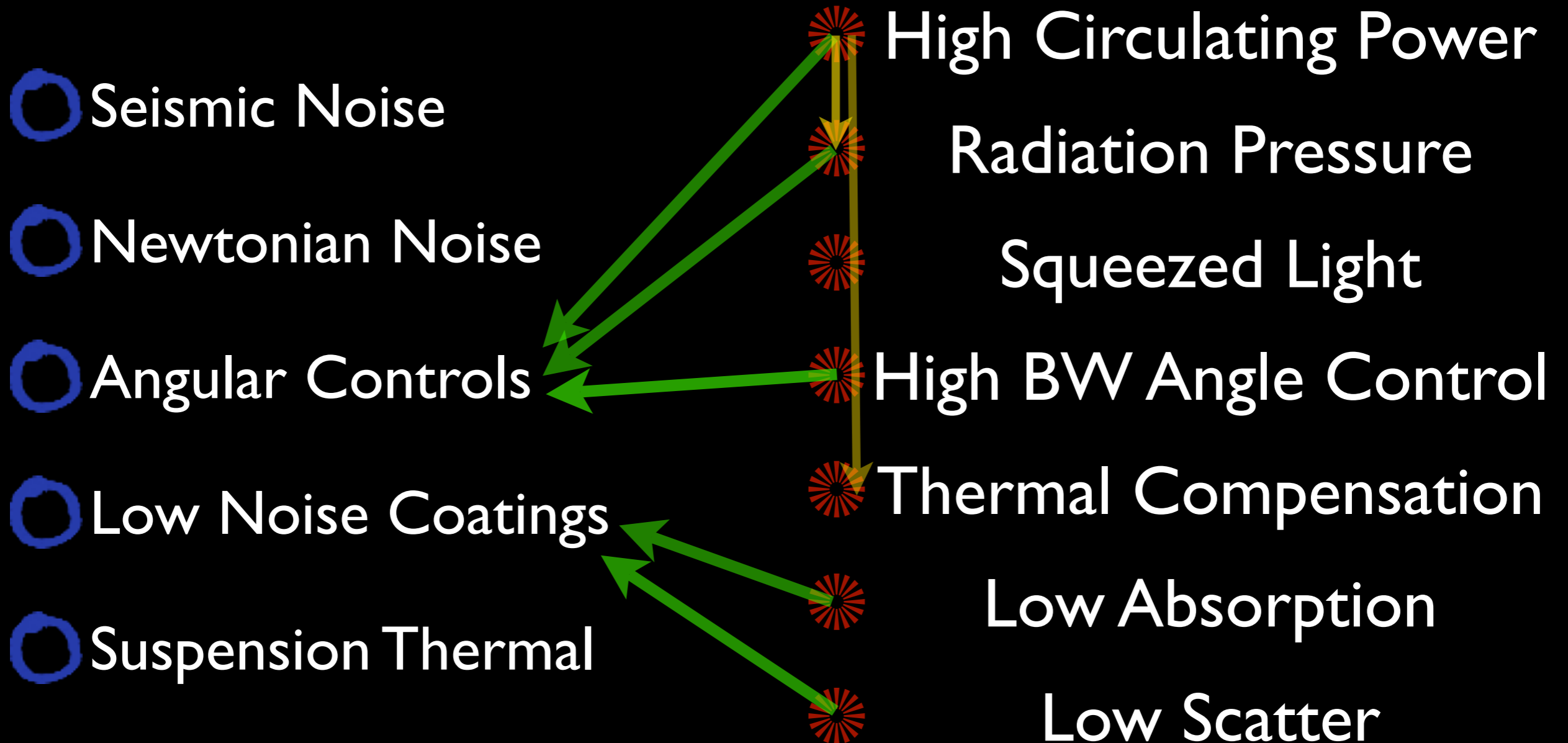
Squeezed Light

High BW Angle Control

Thermal Compensation

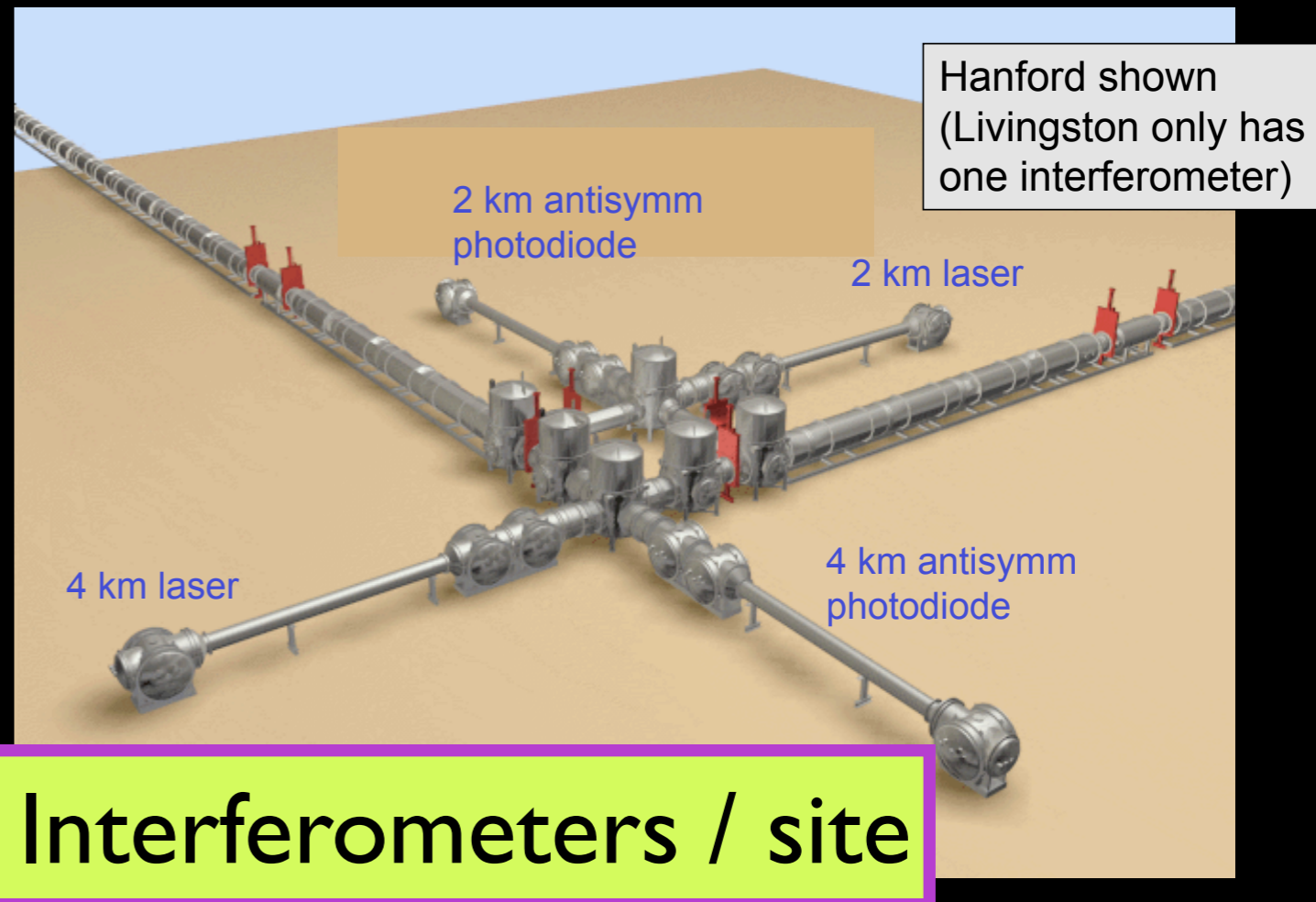
Low Absorption

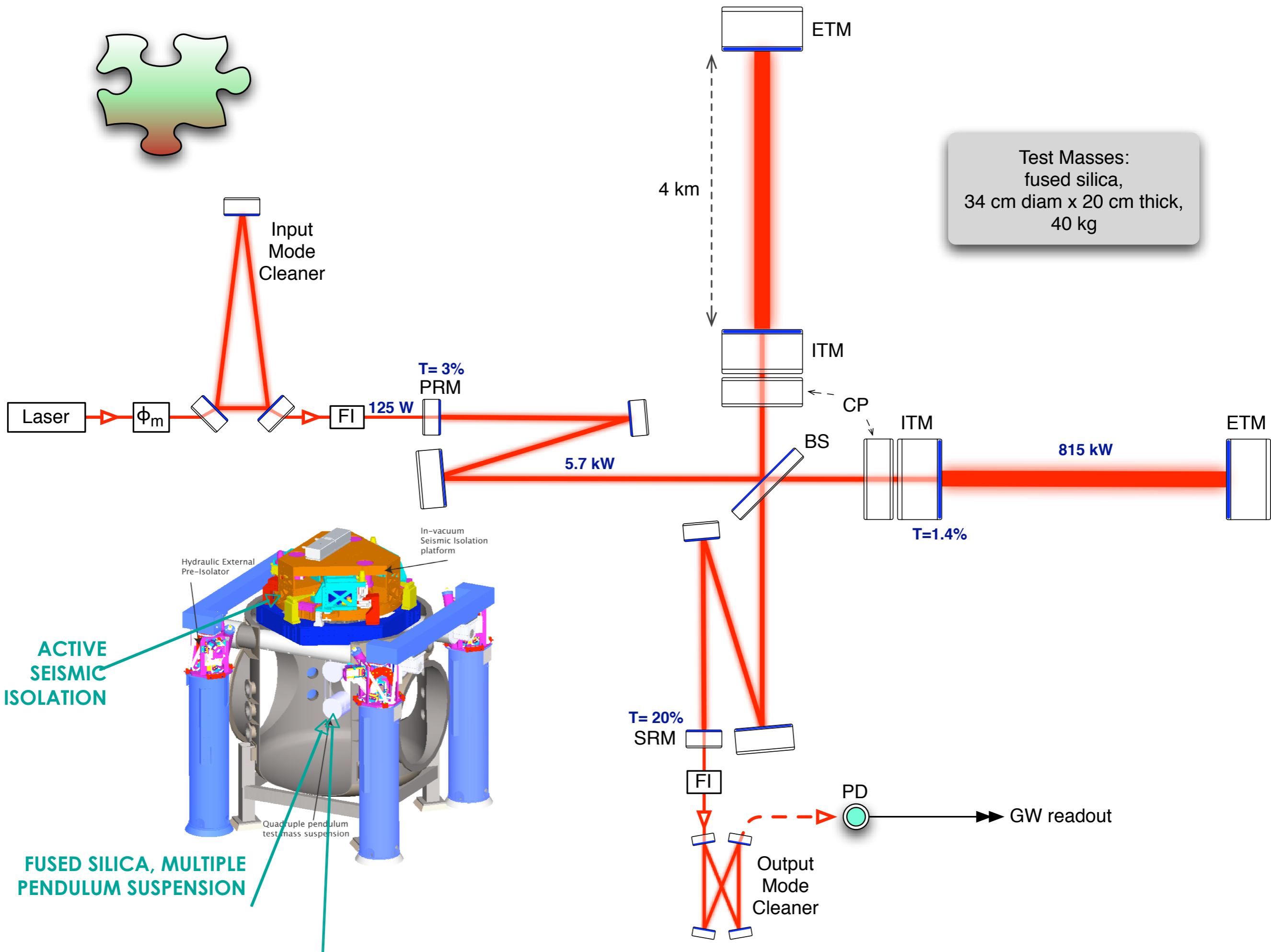
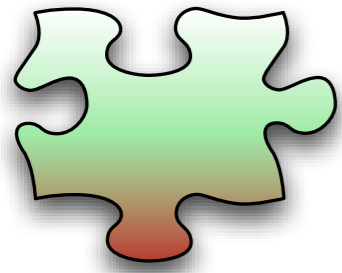
Low Scatter



What to do?

- AdvLIGO/Virgo
~ 2014-2017
- Need Adv+ IFO R&D
now, before 3G.
- Need coincidence for
both low f & high f.
- H1/H2 + L1/L2 + V1/V2

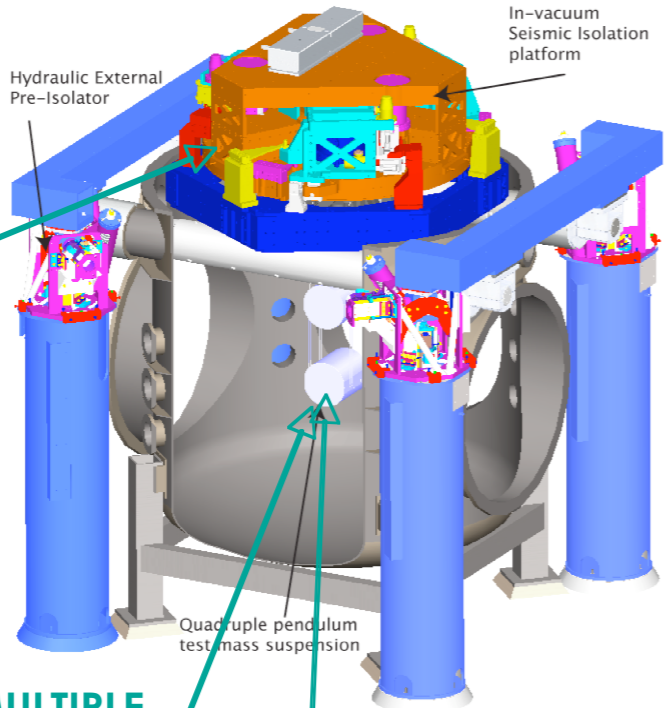




Test Masses:
fused silica,
34 cm diam x 20 cm thick,
40 kg

ACTIVE
SEISMIC
ISOLATION

FUSED SILICA, MULTIPLE
PENDULUM SUSPENSION



T= 20%
SRM

FI

PD

GW readout

Output
Mode
Cleaner

Input
Mode
Cleaner

Laser

ϕ_m

FI

125 W

T= 3%
PRM

5.7 kW

4 km

ETM

ITM

CP

ITM

BS

815 kW

ETM

T=1.4%

In-vacuum
Seismic Isolation
platform

Hydraulic External
Pre-Isolator

Quadruple pendulum
test mass suspension

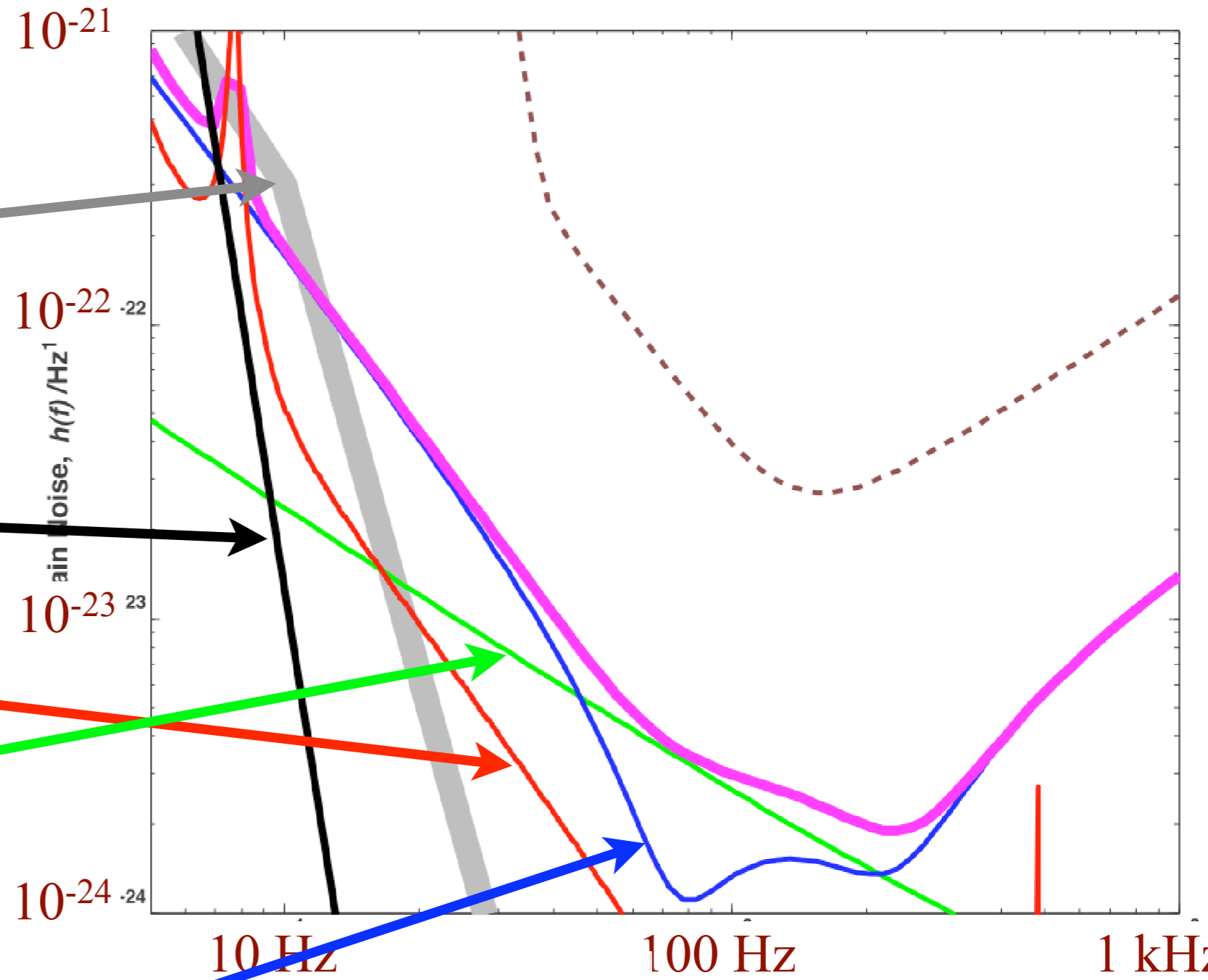
- Newtonian background, estimate for LIGO sites

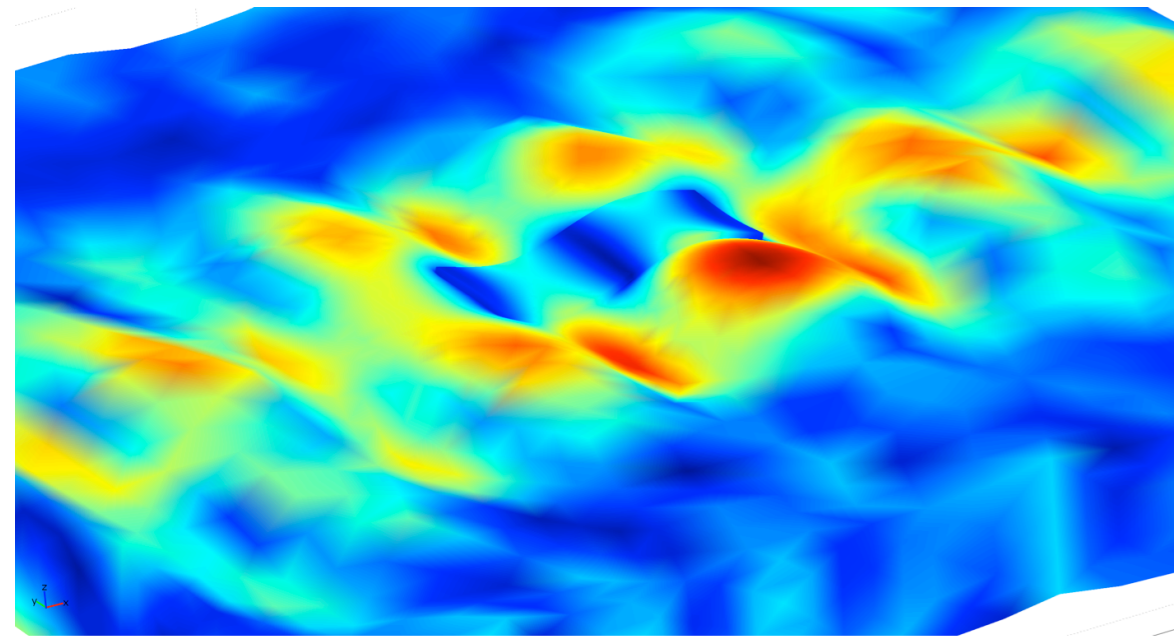
- Seismic 'cutoff' at 10 Hz

- Suspension thermal noise

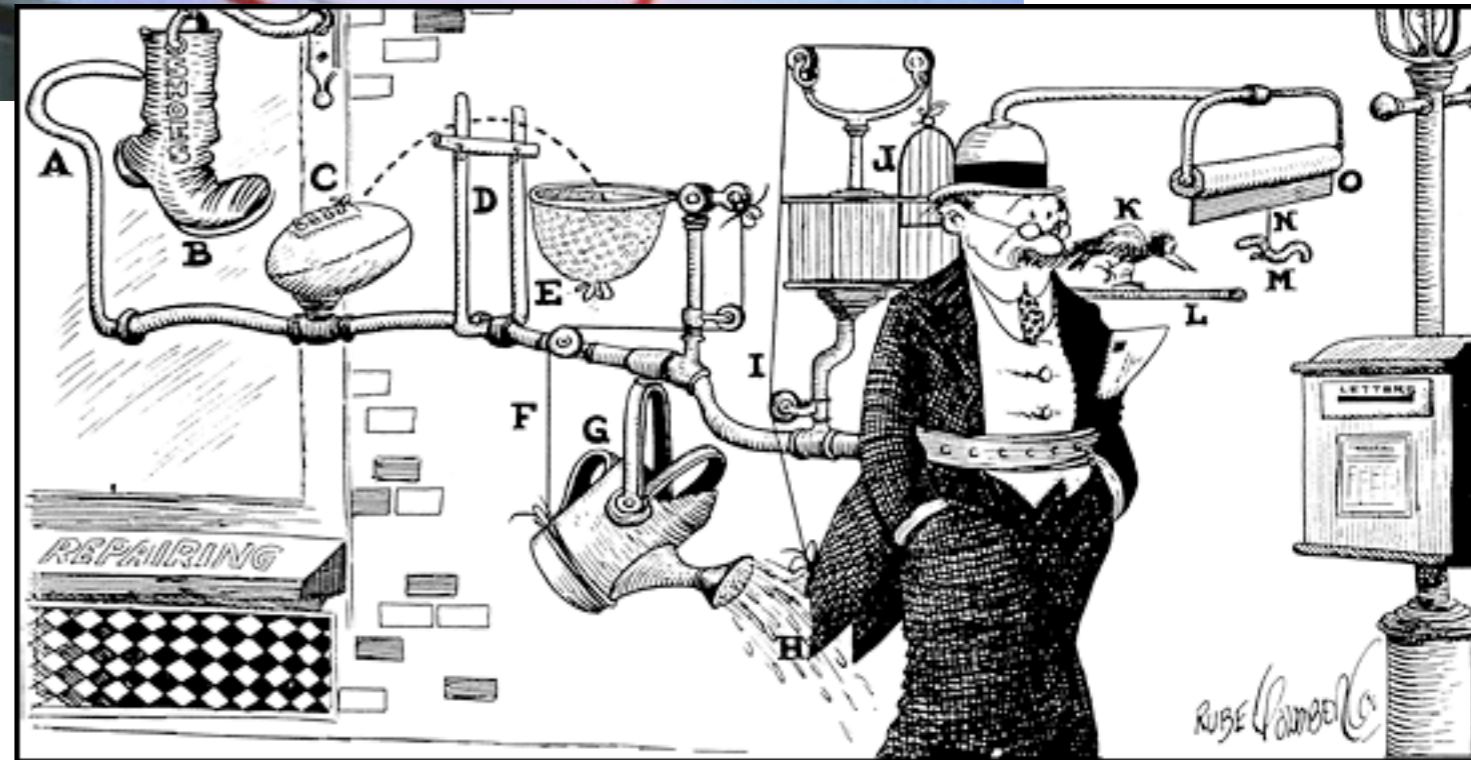
- Test mass thermal noise

- Unified quantum noise dominates at most frequencies for full power, broadband tuning





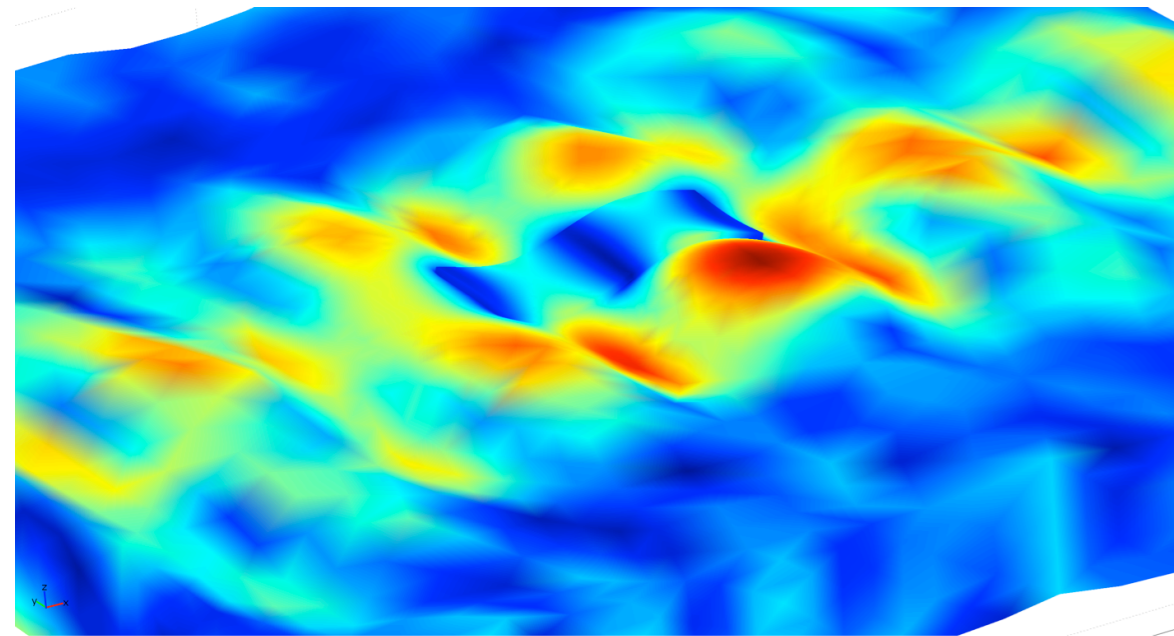
FEA of Concrete Slab



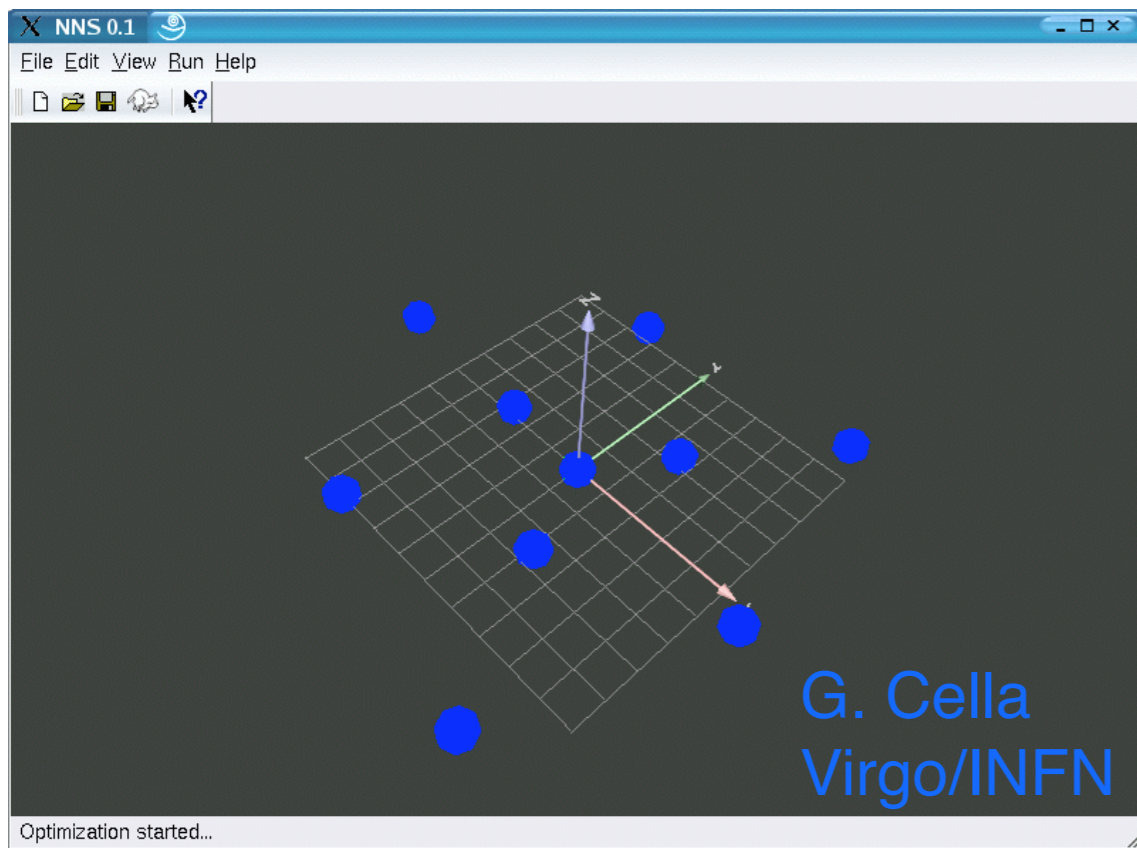
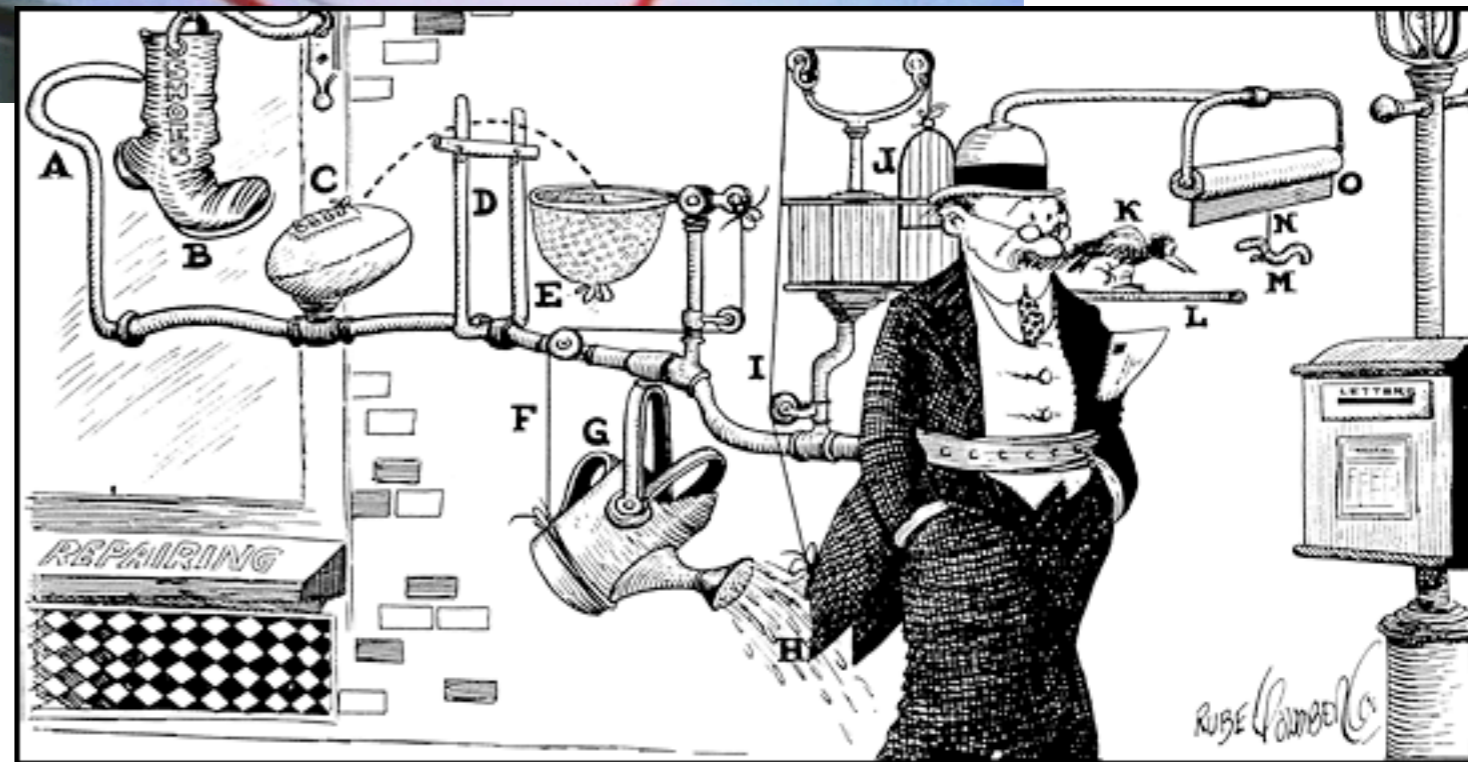
● Noise Cancellation

- Accelerometers measure ground motion
- Adaptive algorithm estimates GG noise
- Subtraction done through software

G. Cella
Virgo/INFN



FEA of Concrete Slab



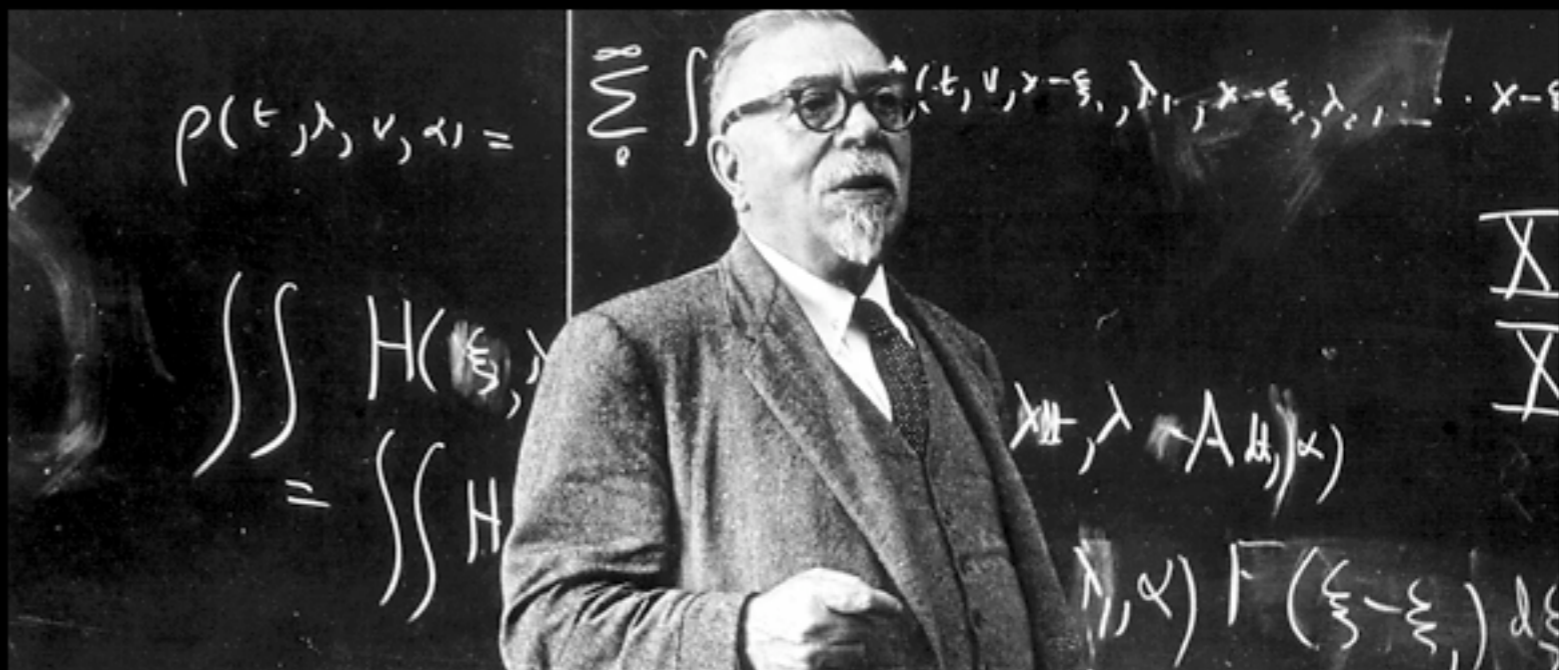
● Noise Cancellation

- Accelerometers measure ground motion
- Adaptive algorithm estimates GG noise
- Subtraction done through software



Wiener Filter

$$x[n] = \sum_{i=0}^N a_i w[n - i]$$



Norbert Wiener, MIT

Block Toeplitz

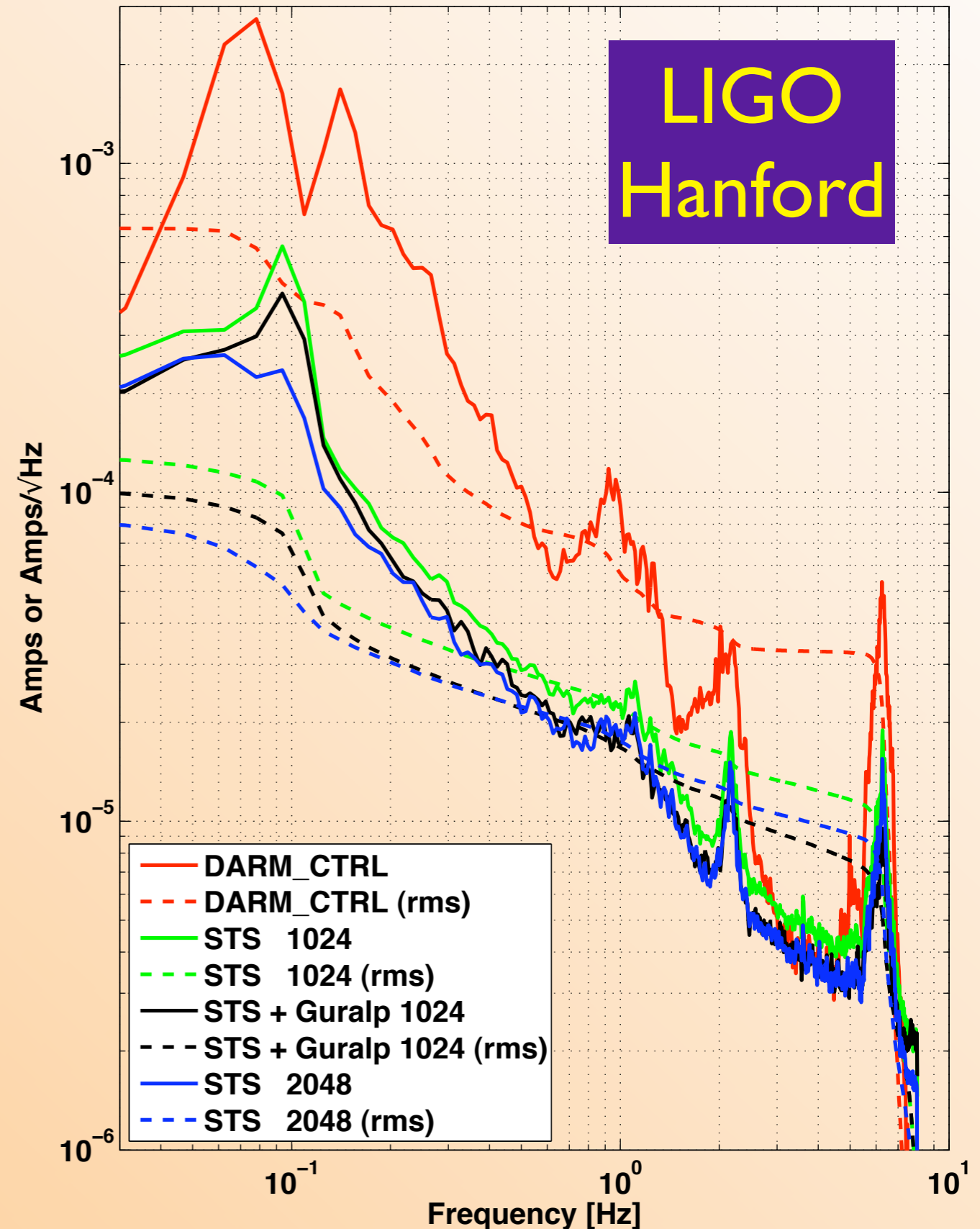
| | | | | | |
|----------|--------------|----------|--------------|---|---|
| $R_w[0]$ | $R_w[1]$ | \dots | $R_w[N]$ | $\begin{bmatrix} a_0 \\ a_1 \\ \vdots \\ a_N \end{bmatrix} = \begin{bmatrix} R_{sw}[0] \\ R_{sw}[1] \\ \vdots \\ R_{sw}[N] \end{bmatrix}$ | Cross Correlation Matrix |
| $R_w[1]$ | $R_w[0]$ | \dots | $R_w[N - 1]$ | | |
| \vdots | \vdots | \ddots | \vdots | | |
| $R_w[N]$ | $R_w[N - 1]$ | \dots | $R_w[0]$ | | |

Input Signal (PEM)
Covariance Matrix

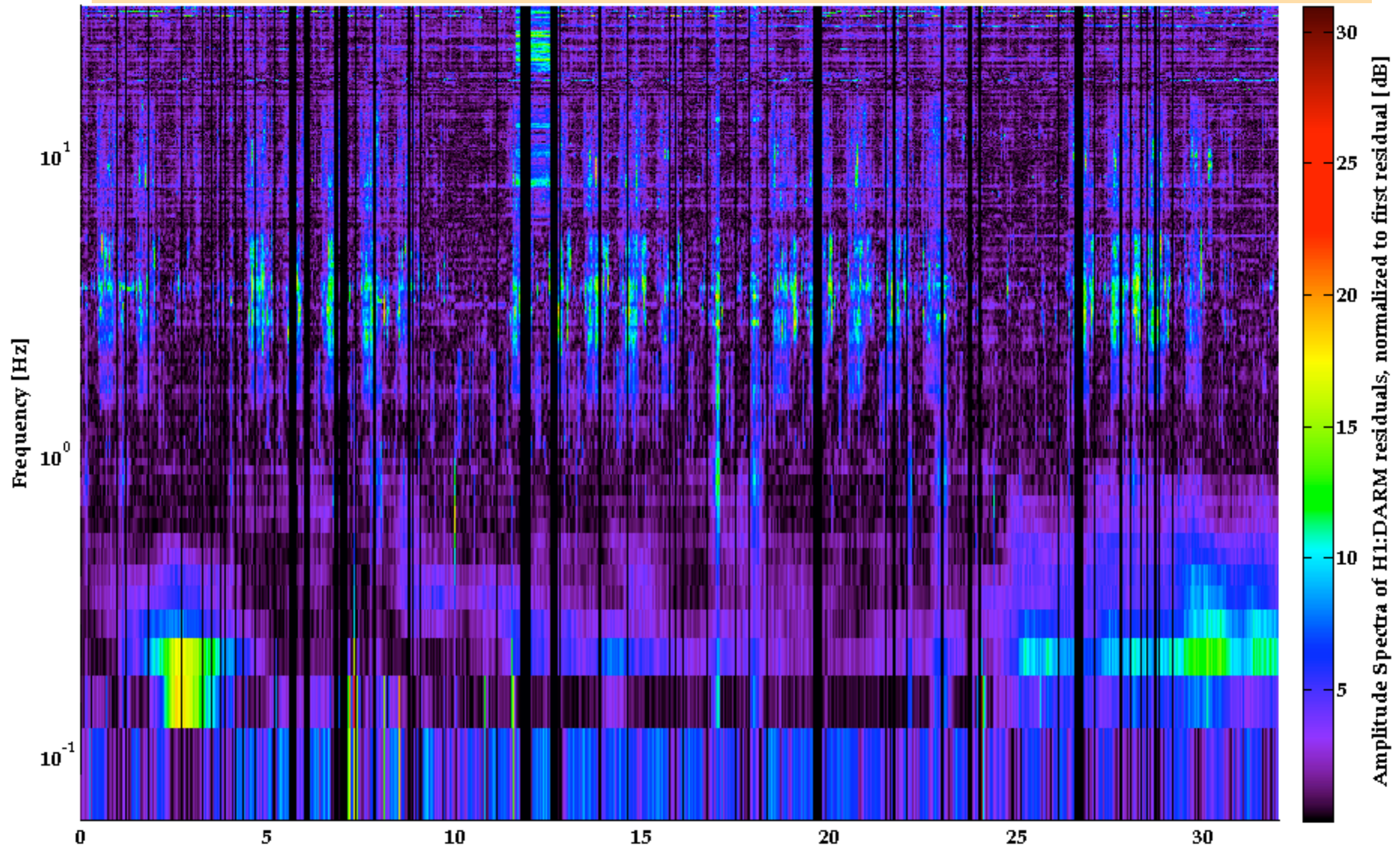
Wiener FF for L- Control

MISO Wiener Filter based subtraction

- * **Simulated** reduction of mirror control force
- * >5x reduction in RMS in addition to the Active Isolation (HEPI)
- * Using ground seismometers only.
- * *No knowledge* of transfer functions required.



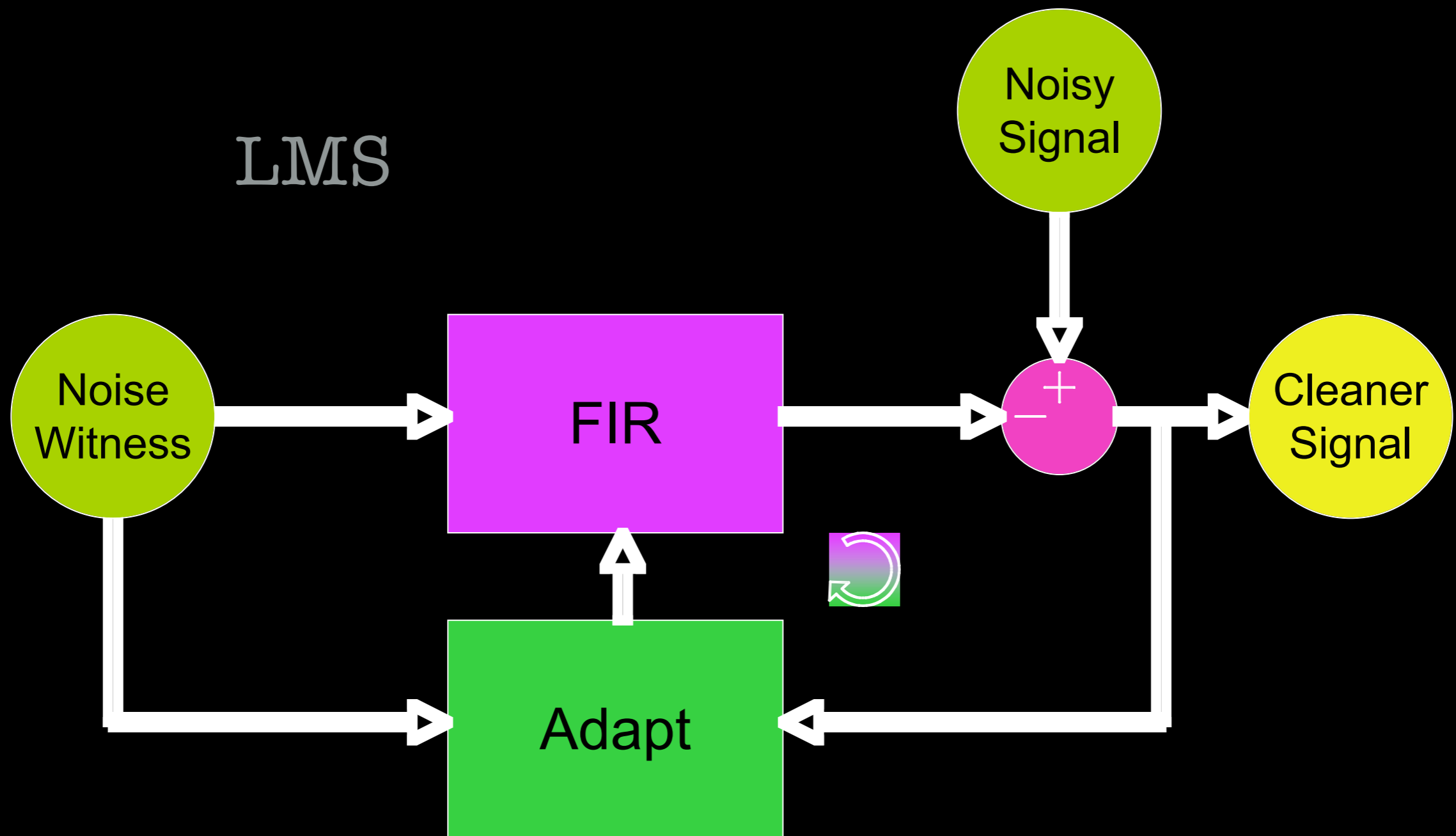
Wiener Performance



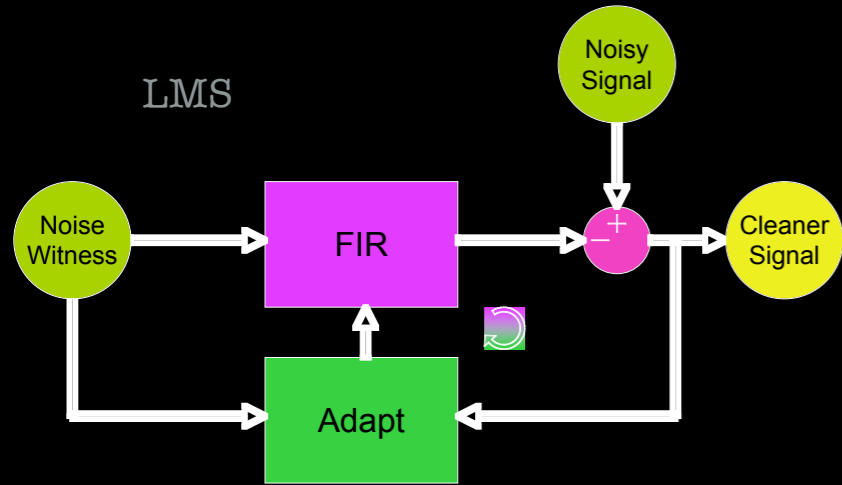
Time [Days]: t0 is 873423660 Black vertical stripe indicates no data available

JENNE DRIGGERS, CIT

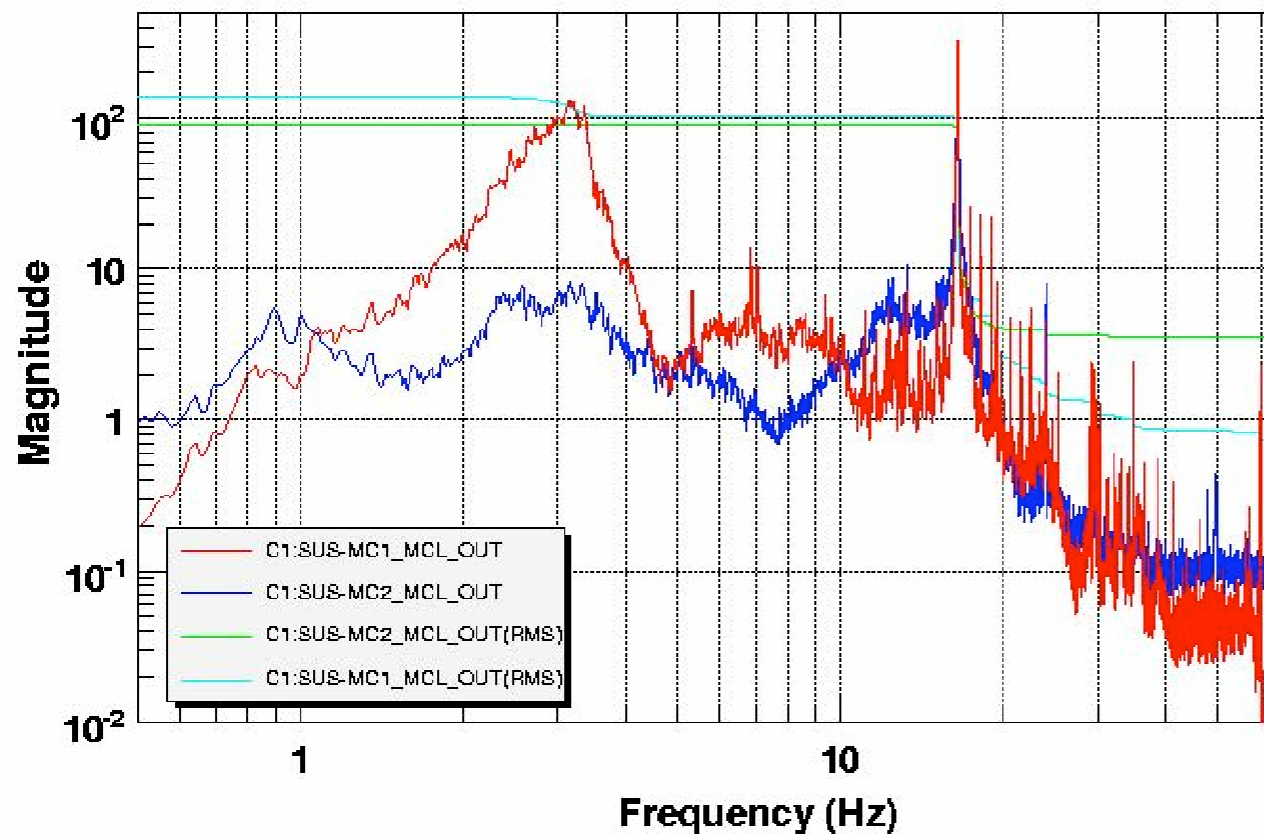
From LMS to Filtered-X



From LMS to Filtered-X



Power spectrum

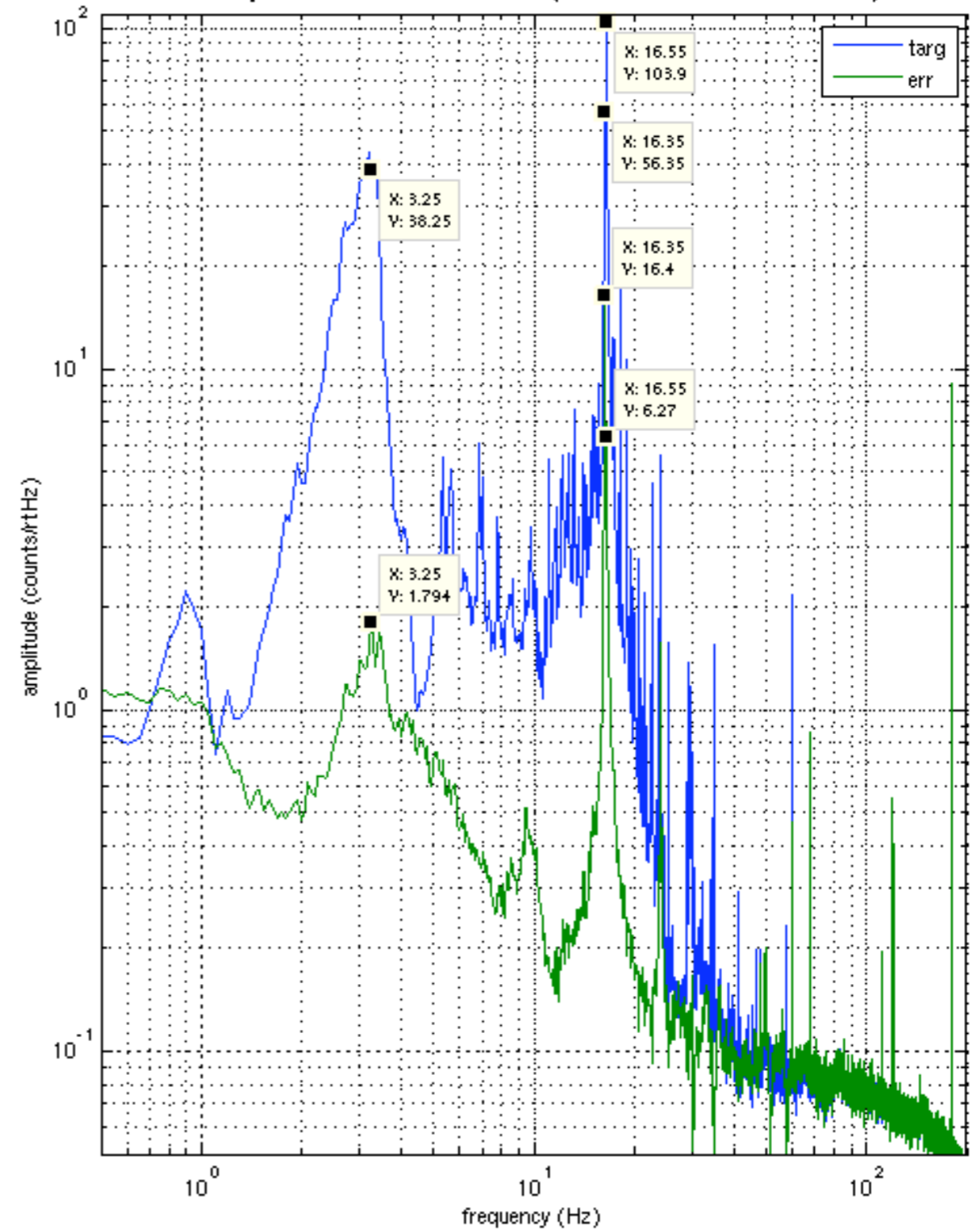


T0=19/03/2008 18:55:13

Avg=20

BW=0.

Adaptive Filter Result (100 runs on 1024 s)



Adaptive FF running online at CIT 40m prototype

Coating Possibilities

- Standard approach: find a **lower loss** material
 - Reached a plateau? Is $\phi \sim 10^{-5}$ possible?
- Use **less layers** (lower Finesse, higher loss)
- Use **thinner layers** ($\lambda \approx 404\text{-}532$ nm)
- Optimize **periodic pattern** (e.g. 1/8-3/8 layers) to minimize high loss material's volume
- Optimize Brownian w/ the **ABC scheme** (see *Yanbei Chen's talk*)
- ABC scheme w/ **multi-color** arm cavity readout

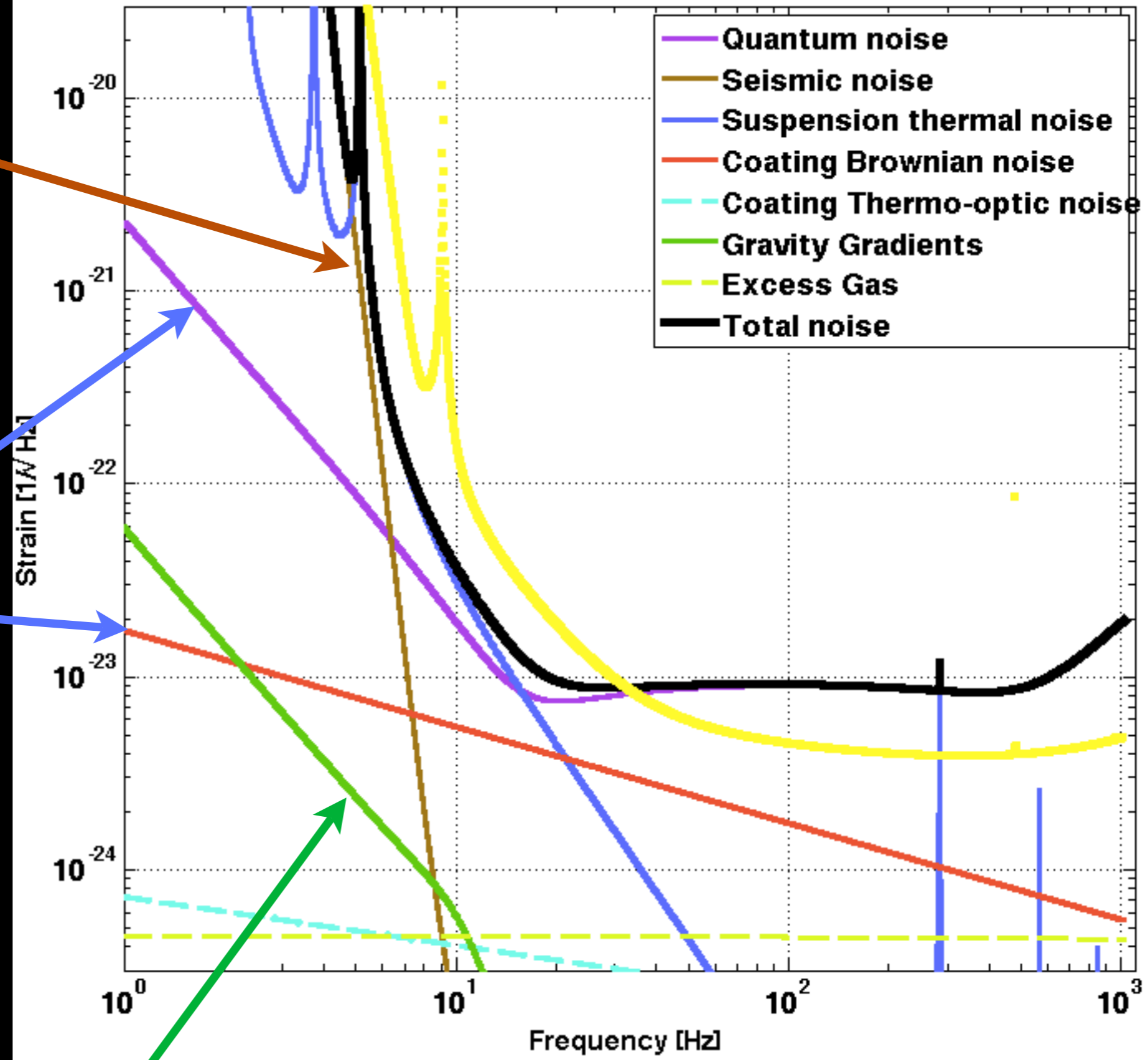
Red v. Blue



CONFIDENCE

IT LETS YOU WIELD ANY
COLOR
LIGHTSABER YOU LIKE.

bLIGO Noise Curve: $P_{in} = 5.0 W$



x30 Seismic FF

405 nm laser

- Absorption
- EO Components
- High QE PDs

x10 Newtonian Subtraction