

### **Contents**

### 1. Invention of DFI [2004-2006]

~ Kawamura, Chen, et al

### 2. R&D experiment [2006-2008]

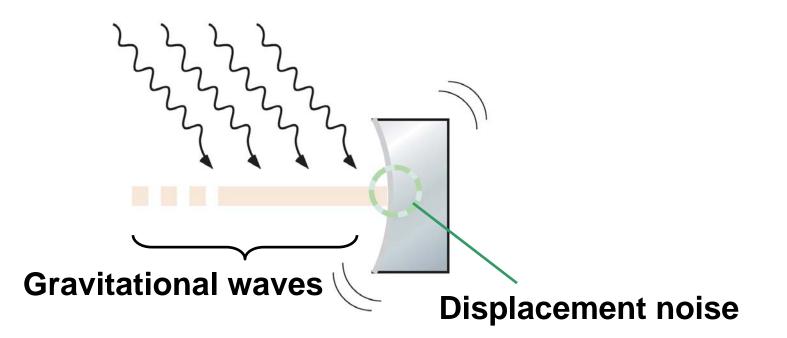
~ Kokeyama, Sato, et al

### 3. Recent study [2008-2009]

~ Nishizawa, Vyatchanin, et al



#### [Kawamura PRL 04]



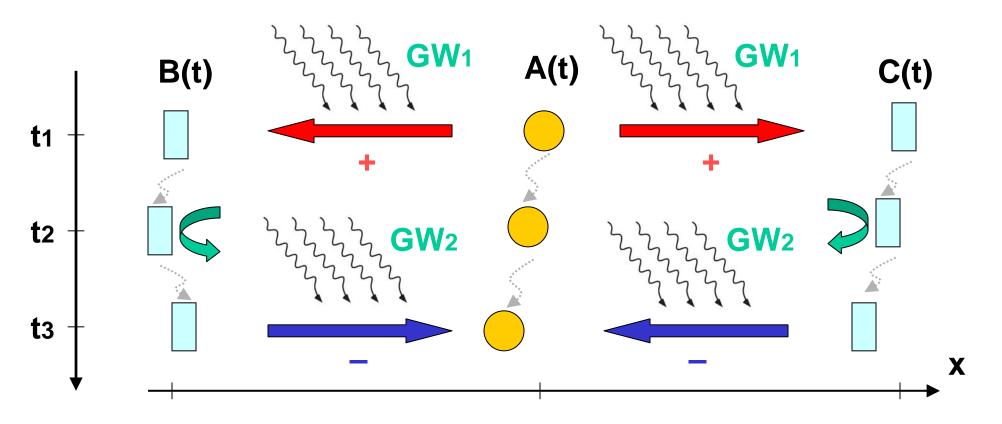
There is a difference between GW and disp. noise.

### **GW : probed during the light travel disp: probed at the reflection**



There is a way to distinguish them!!

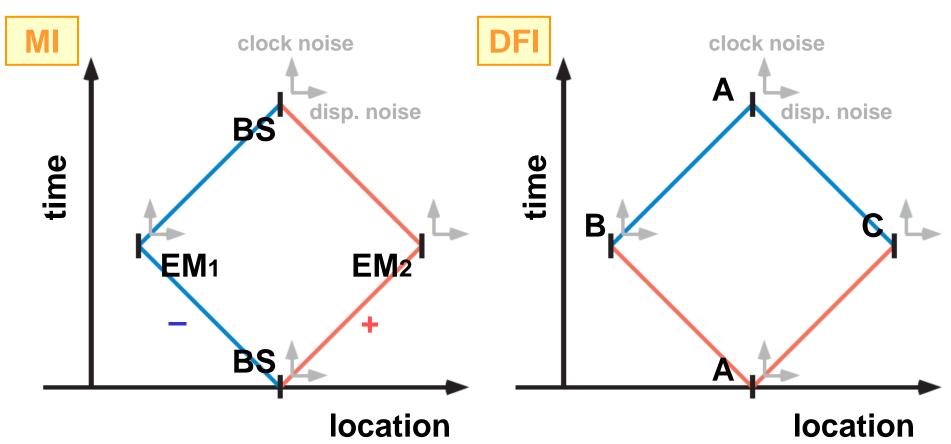
# How can we distinguish GW and disp-noise?



 $[A(t_1)-B(t_2)] + [C(t_2)-A(t_1)] - [A(t_3)-B(t_2)] - [C(t_2)-A(t_3)] + GW_1 - GW_2$ 

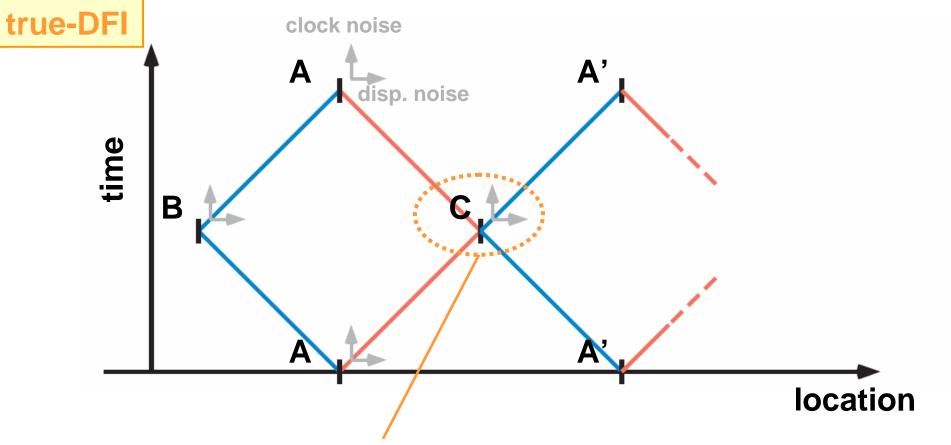
displacement free + GW + clock noise

# **DFI vs Michelson ifo**



MI : no clock noise (freq noise), but disp. noise DFI : no disp. noise, but clock noise Actually we want to cancel both.

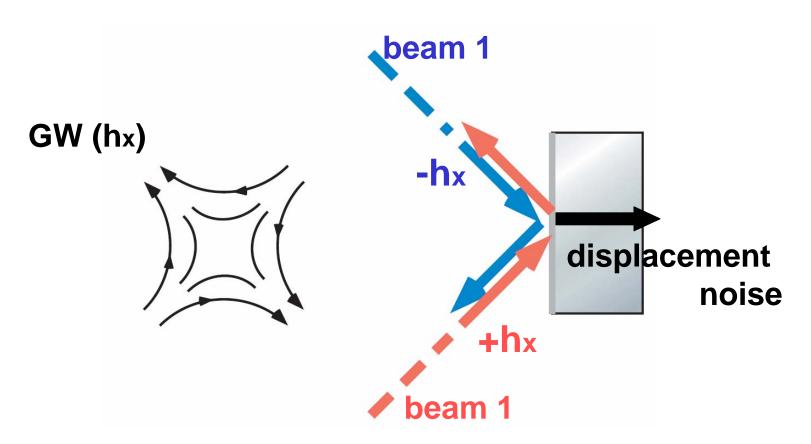
# **Displacement-noise and Frequency-noise free**



Both disp noise and freq noise are cancelled (for C).

- We should do the same for all the components
- GW signal should not be cancelled

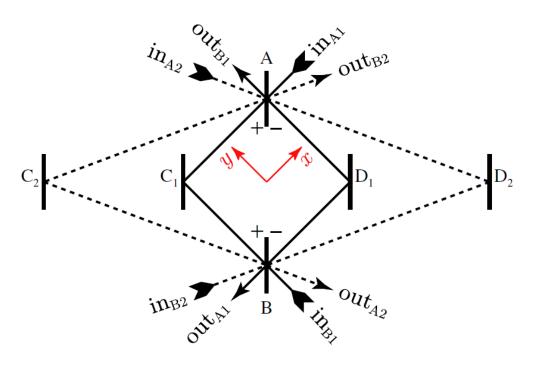
## Separation of GW and disp noise



- Sign of GW is opposite between beam 1 and beam 2
- After subtraction, GW signal is prop to frequency

# **Bi-directional Mach-Zehnder ifo**

[Chen PRL 2006]



1<sup>st</sup> subtraction (mirror disp) bi-directional measurement (ex. AC1B – BC1A)

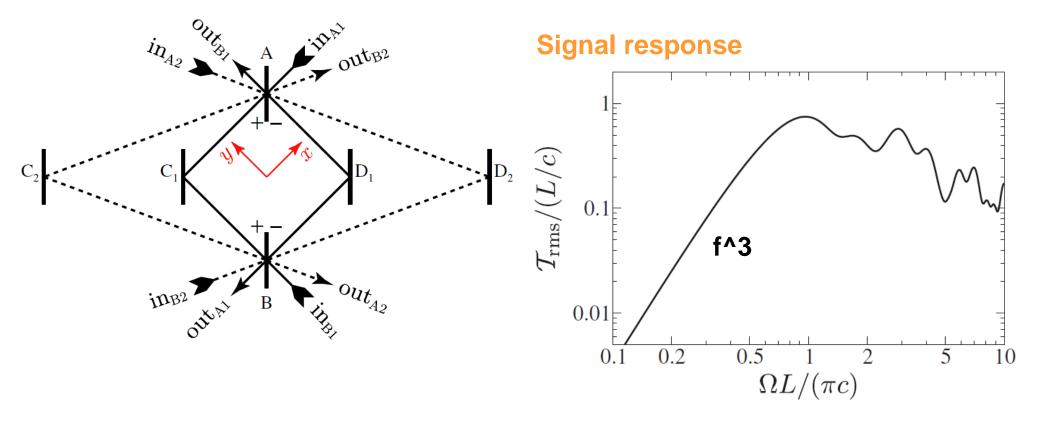
2<sup>nd</sup> subtraction (freq noise) MZ interferometer (ex. AC<sub>1</sub>B – AD<sub>1</sub>B)

3<sup>rd</sup> subtraction (BS disp) 2 MZ combination (ex. AC<sub>1</sub>B – AC<sub>2</sub>B)

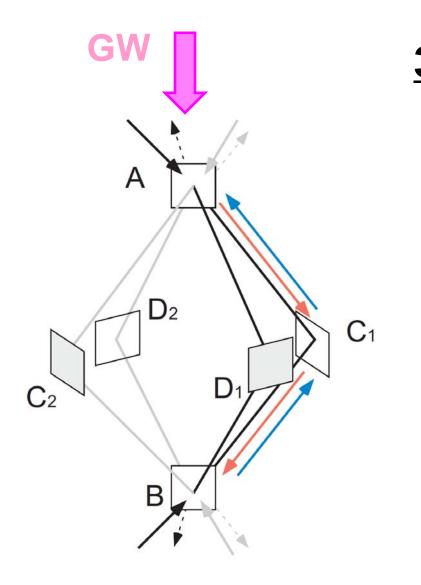
Signal is proportional to f^3 (bad)

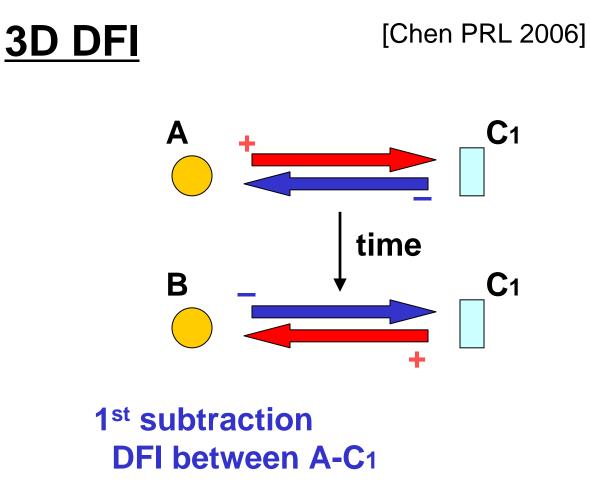
# **Bi-directional Mach-Zehnder ifo**

[Chen PRL 2006]



Signal is proportional to f^3 (bad)

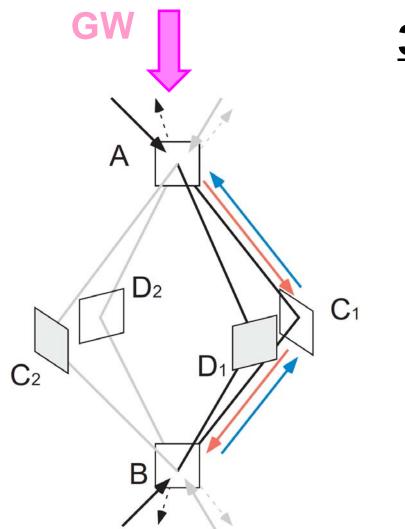




2<sup>nd</sup> subtraction A-C1 and B-C1 at different time



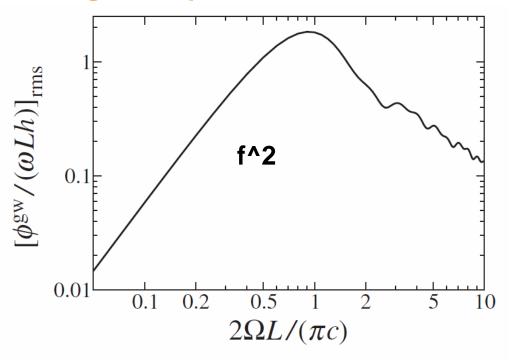
Signal is proportional to f<sup>2</sup> (bad)



### <u>3D DFI</u>

#### [Chen PRL 2006]

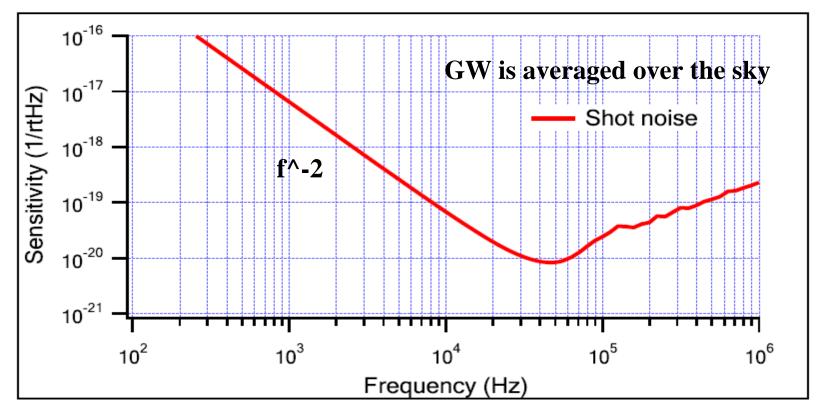
#### **Signal response**



Signal is proportional to f^2 (better)

# **3D-DFI noise spectrum**

#### L=4km, I=10W each



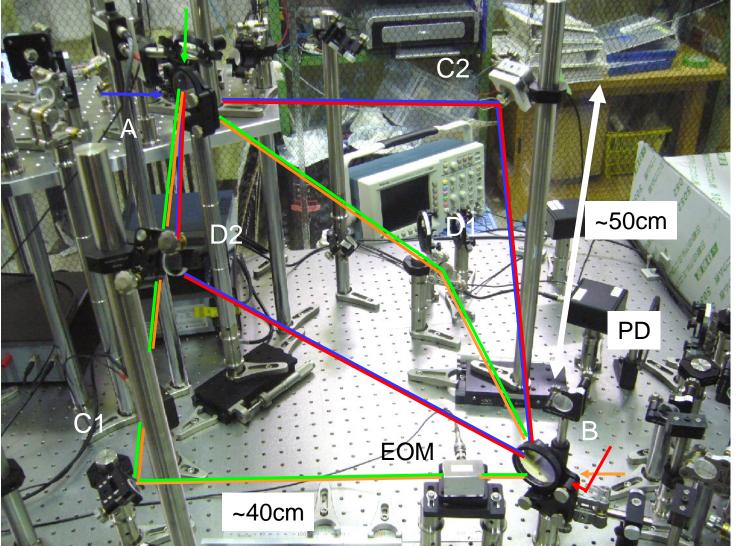
- No displacement noise at any frequency
- Shot noise is bad at low frequency

### After the discovery...

- R&D experiments in Japan
  - 2D DFI
    3D DFI
- Effort to find a better configuration
  - time-delay device
  - narrow-band DFI
  - DFI with rigid platform
  - resonant-type DFI

# **3D DFI table-top experiment**

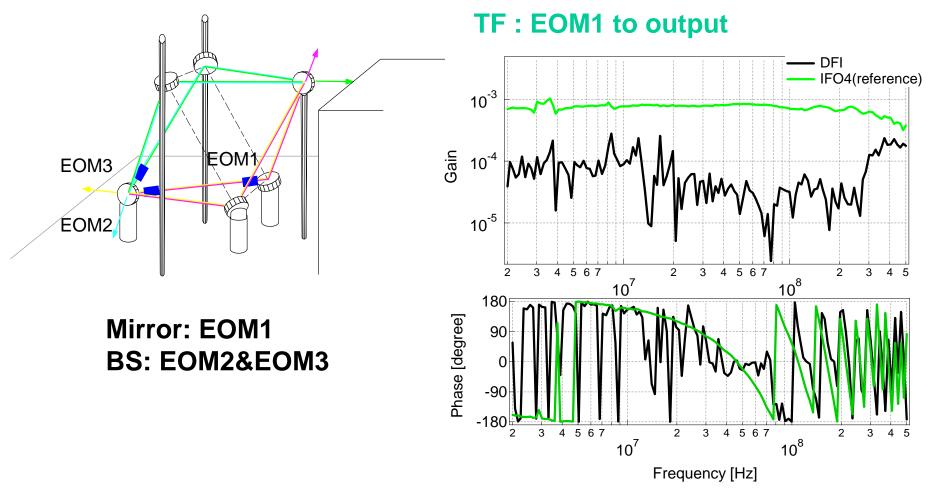
[Kokeyama 2009]



- Alignment in 3D
- High frequency (~MHz)
- TF measurement using EOMs

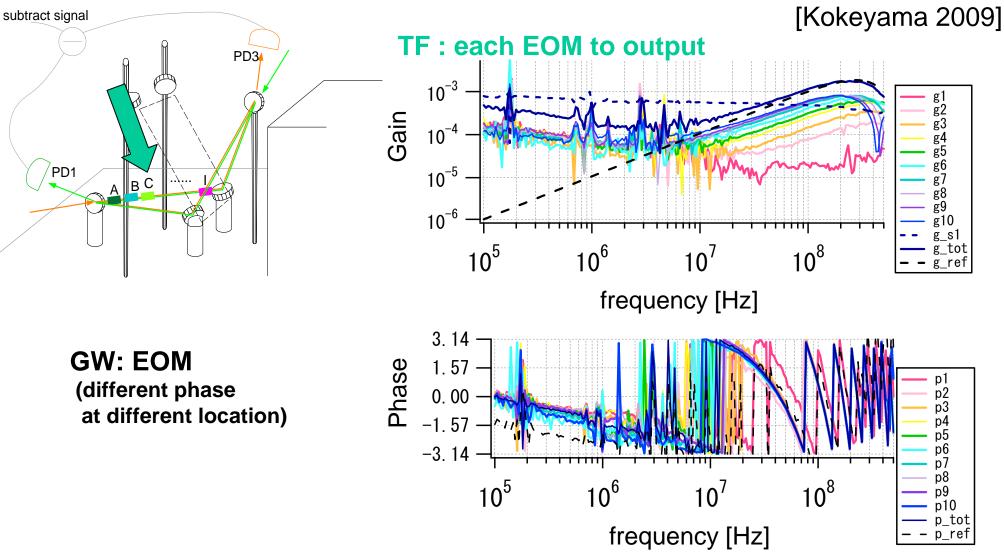
# **Verification of noise cancellation**

[Kokeyama 2009]



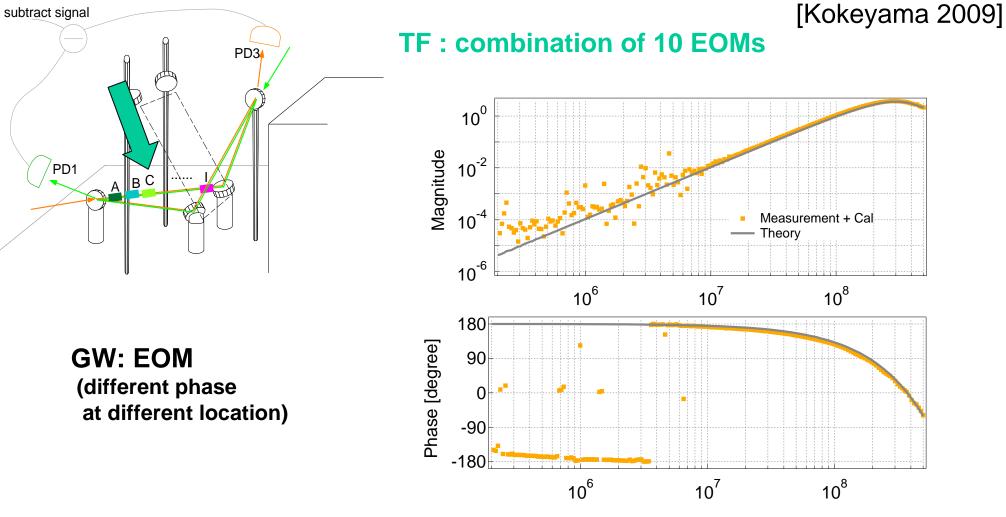
Displacement components is reduced by 20~30dB

# **Verification of GW non-cancellation**



GW signal can be imitated by combining these outputs

# **Verification of GW non-cancellation**



Frequency [Hz]

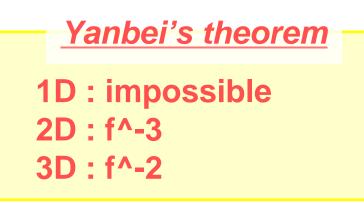
### After the calibration, DFI signal response is obtained.

## Weak-response problem



Even with the 3D DFI, shot noise  $\propto$  f^-2

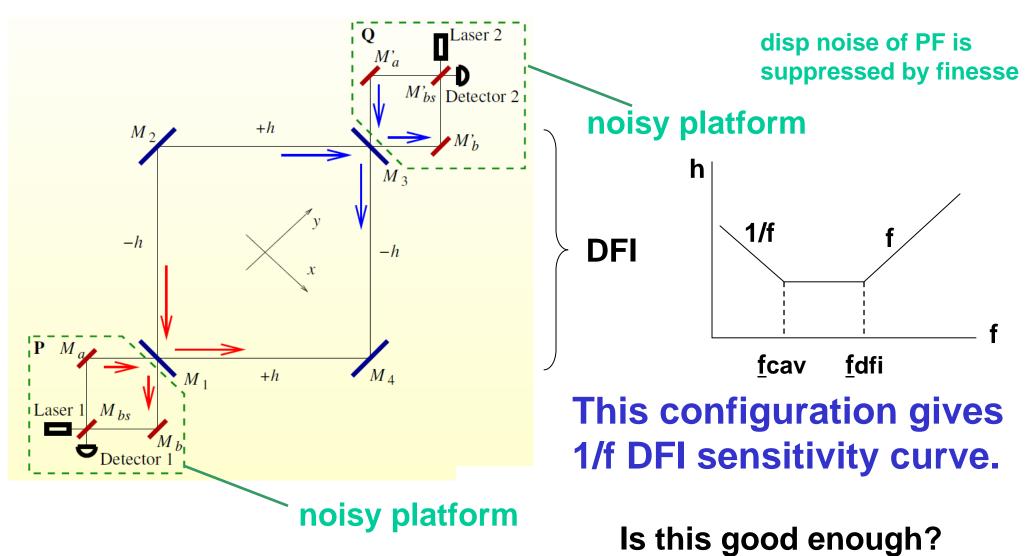
Moreover...



<u>one possible ways to circumvent the problem</u> Make a cavity and ignore disp noise outside the cavity

## **Resonant speed-meter DFI**

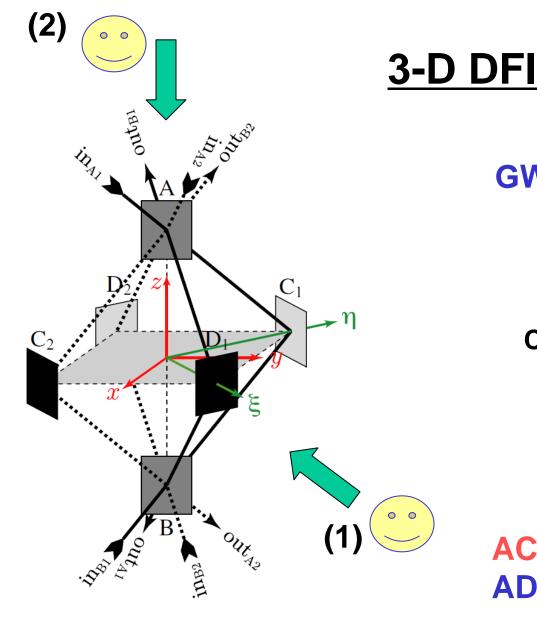
[Vyatchanin 2009]



# **Summary and discussions**

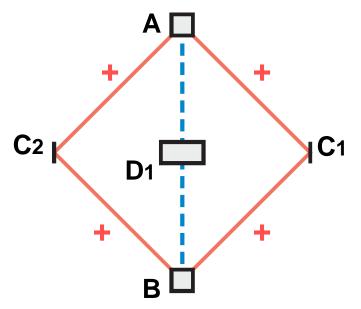
- 3D DFI with f<sup>-2</sup> slope is so far the best for complete DFI
- R&D experiments done
- Resonant partial-DFI seems nice with 1/f slope
- Is it possible to make the slope better?
- DFI might be useful for a space-based GW detector
- Maybe with atomic interferometer...

End of the slides



### [Chen PRL 2006]

GW from (1)'s direction



ACjB is no MZ (no GW). ADjB's GW cancels with BDjA's.

No response to GW from (1)'s direction.

