



# GRB-triggered searches for gravitational waves from compact binary inspirals in LIGO and Virgo data during S5/VSR I

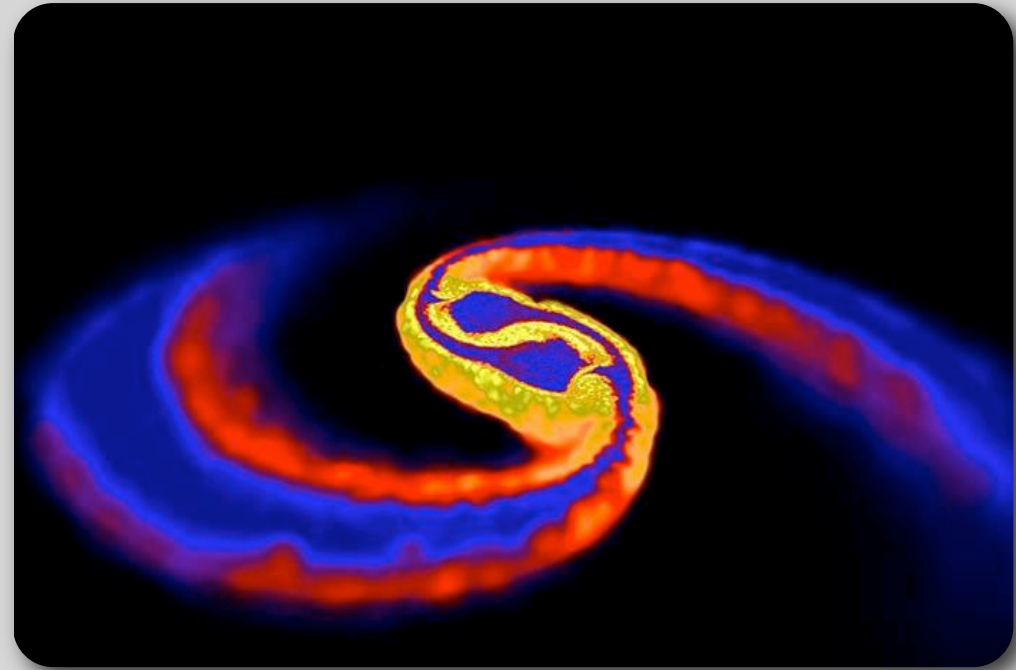
Nickolas Fotopoulos (UWM) for the [LIGO](#) Scientific Collaboration and the Virgo Collaboration

 APS April 2009 meeting, Denver, CO | 2009.05.02

# Short GRBs: ideal targets for GW astronomy (I)

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- Most short GRBs are probably NSs disrupted by compact companions in the final stages of inspiral.
- A detection will constrain component masses and spins.\*
- A high-SNR detection will constrain NS equations of state.†
- Simultaneous EM/GW observations can measure absolute luminosity distance.‡



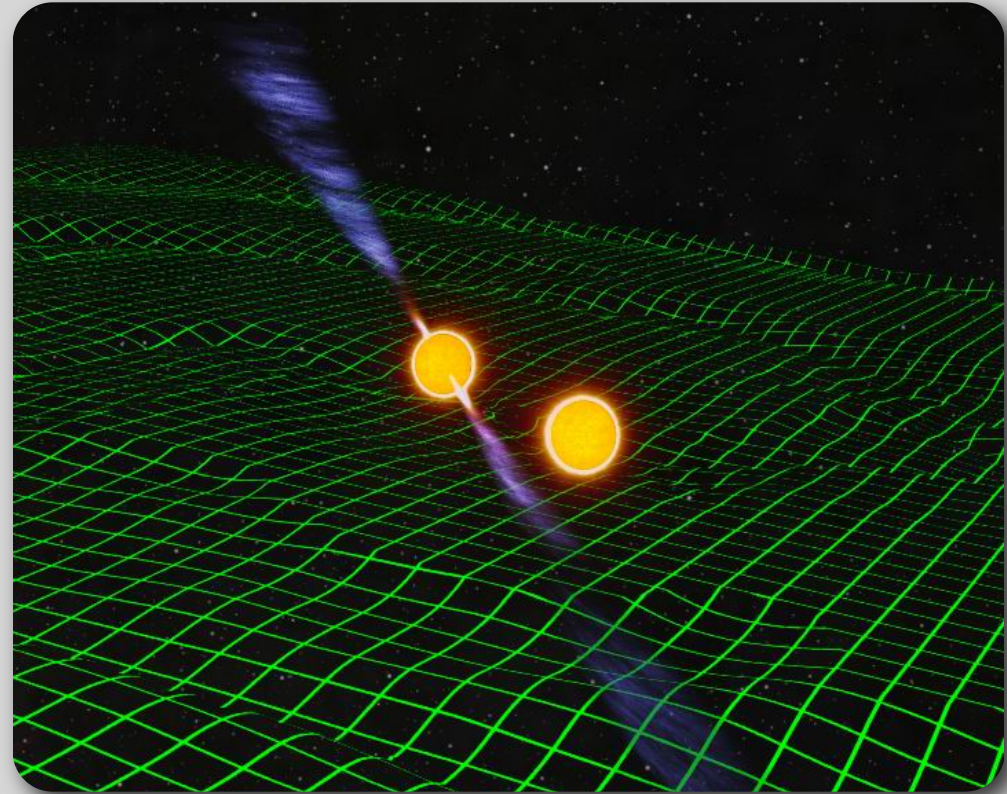
NS-NS merger simulation  
Price and Rosswog

\* Cutler and Flanagan, PRD 1994;  
Finn and Chernoff, PRD 1993;  
Poisson and Will, PRD 1995  
† Flanagan and Hinderer, PRD 2008;  
Read et al, arXiv:0901.3258  
‡ Nissanke et al, arXiv:0904.1017

# Short GRBs: ideal targets for GW astronomy (II)

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- A significant GW candidate with an EM counterpart is a far more compelling detection.
- The GW emission during inspiral is well modeled. This enables matched filtering, which digs deeply into the detector noise.
- A known time and sky location can be searched with significantly lowered thresholds.



NS-NS inspiral depiction  
John Rowe Animation

# S5: Nov 2005 – Nov 2007 | VSR I: May 2007 – Oct 2007

- 213 GRBs
- 33 short GRBs
- 22 short\* GRBs while two+ GW detectors were taking good data (duty cycle, data quality)

051114	070209
051210	070429B
051211	070512
060121	070707
060313	070714
060427B	070714B
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Already published: no  
inspiral in M3 I  
(Abbott et al, ApJ 2008)

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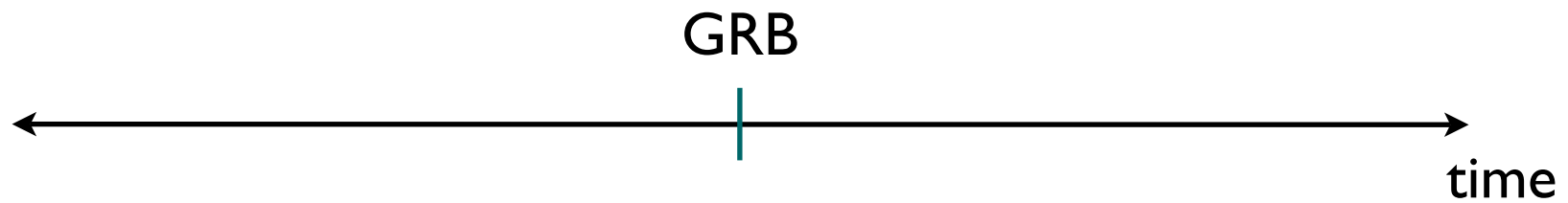
Long duration but other suggestive features

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# Experiment overview

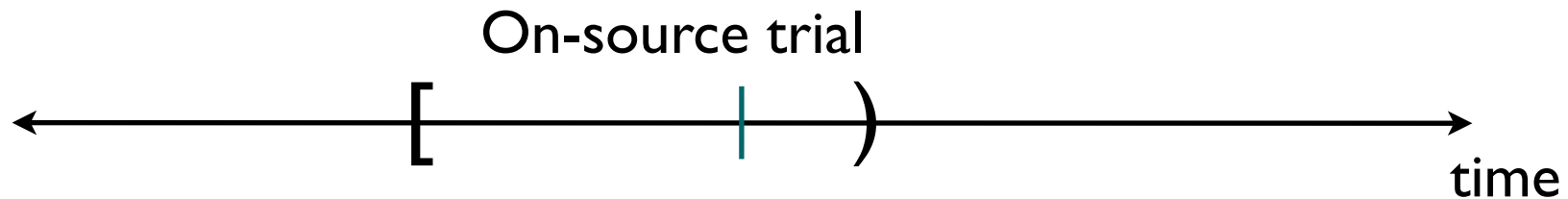
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# Experiment overview

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- We associate GW triggers with GRBs within  $[-5, +1)$  s of the reported GRB time. This is the on-source trial.

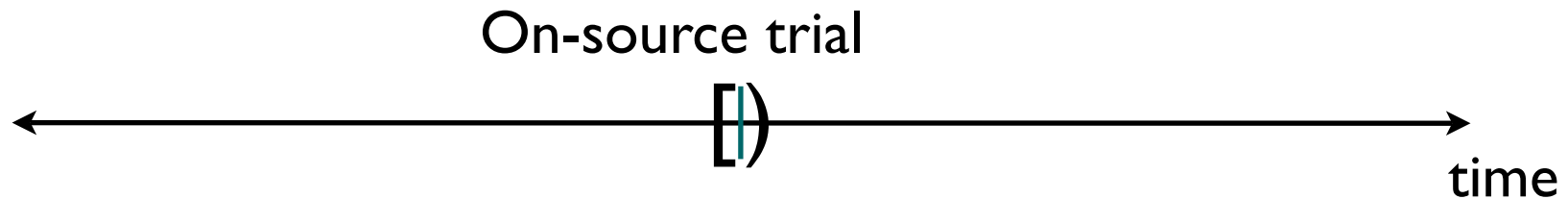




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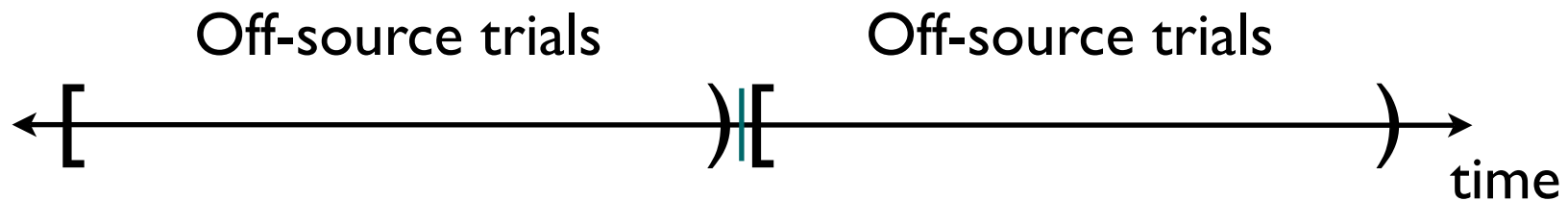
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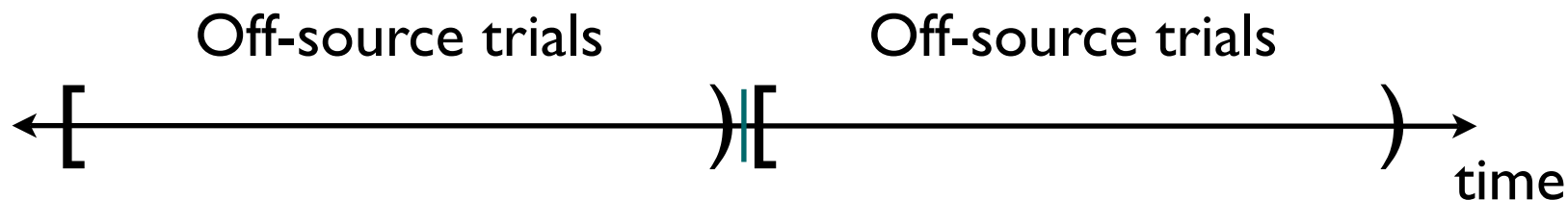
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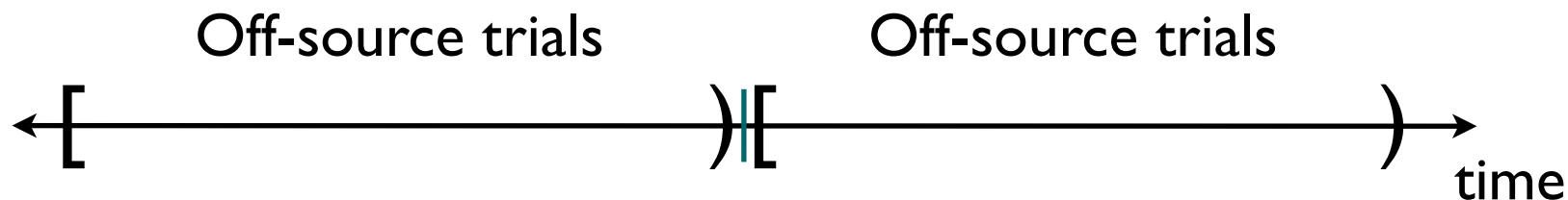
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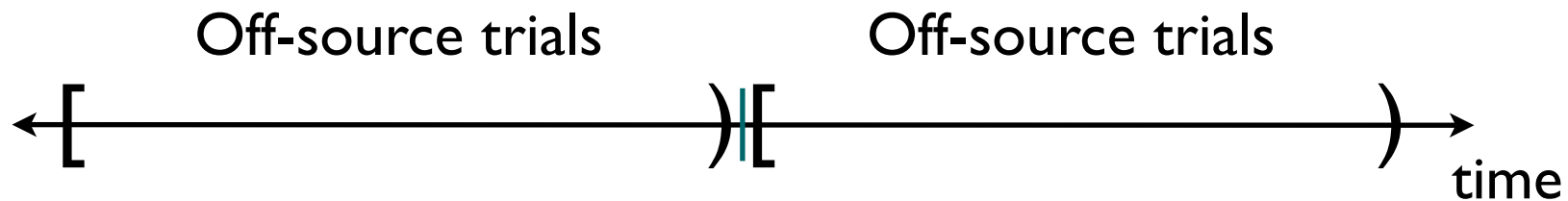


\* Abbott et al, PRD 2006

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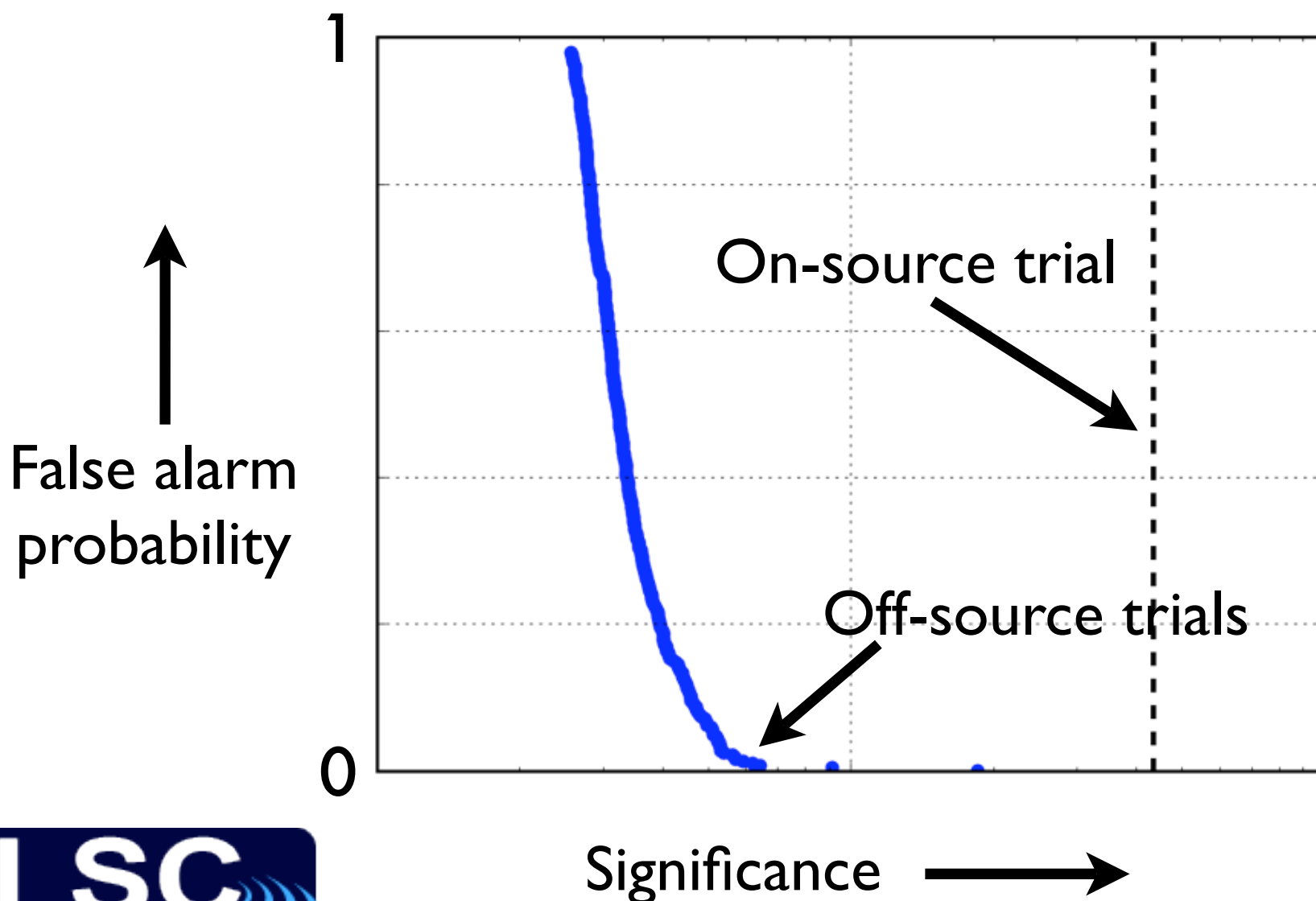


- We **reuse** the hierarchical inspiral search pipeline used in previous LIGO analyses.\*
- We combine injection and off-source trials to form a **likelihood statistic**.

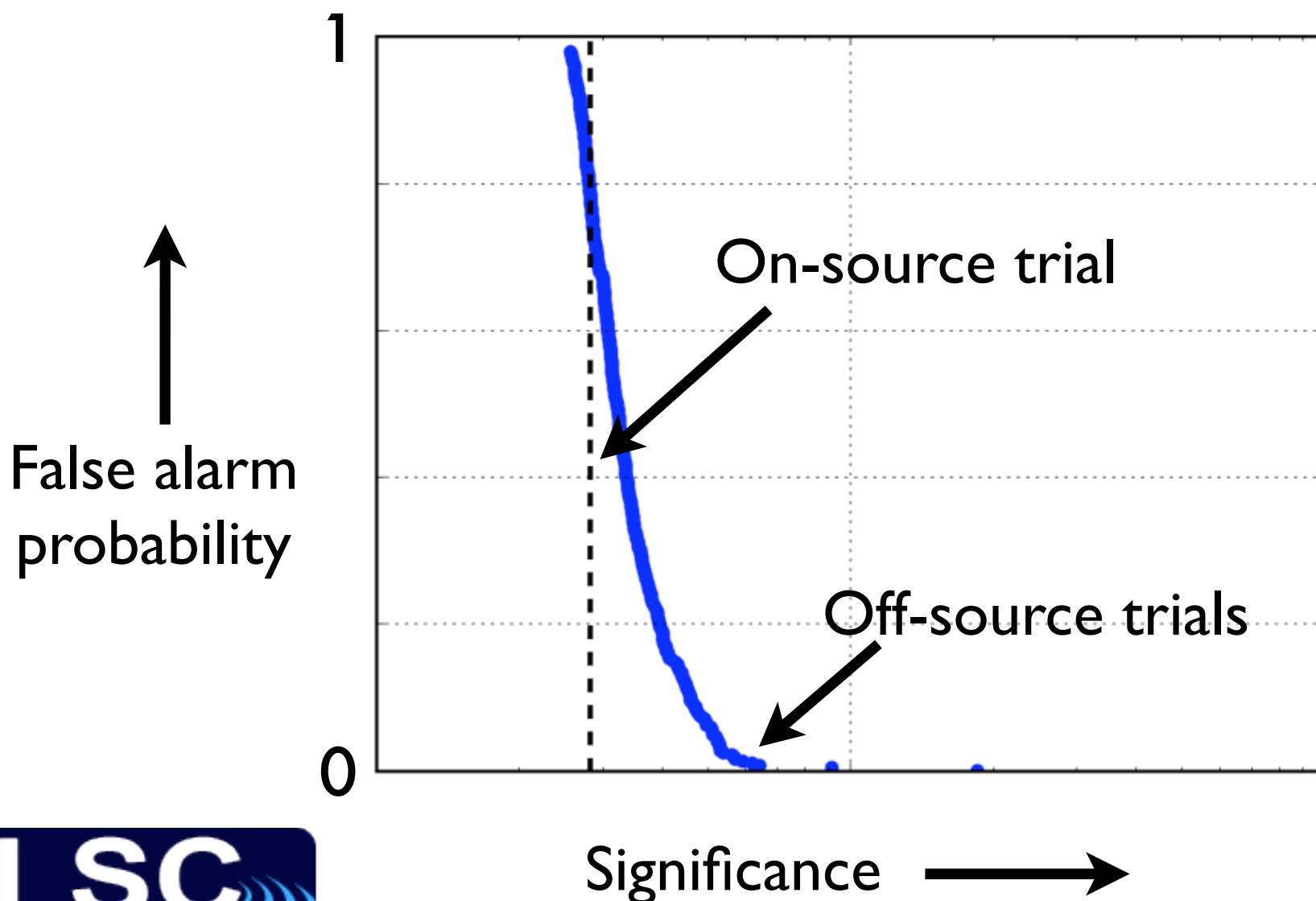


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# What a **detection** might look like

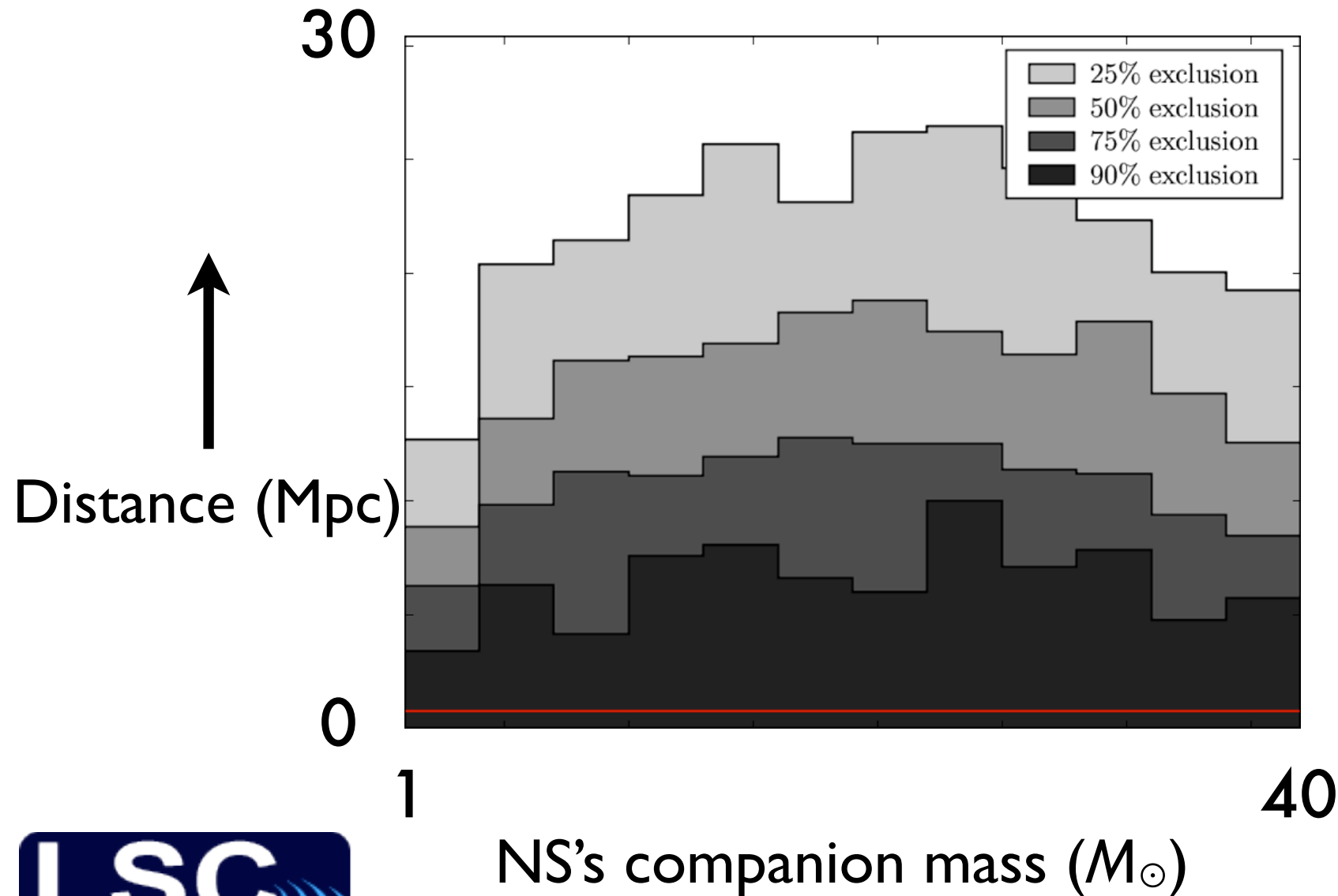


# What a **null result** might look like





# Astrophysical exclusions\* from null results



\* example recycled from GRB 070201 <sup>8</sup>

# Where we are, where we will be

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- LIGO and Virgo are committed to **multi-messenger astronomy**. A coincident detection would provide enormous science.
- We are searching for GWs from **compact binary inspirals** in coincidence with **22 short\* GRBs** in S5/VSR1.
- S6/VSR2 begins soon with enhanced detectors. Advanced detectors should routinely record detectable GWs.

