



LIGO Laboratory / LIGO Scientific Collaboration

LIGO-T0900043-07

ADVANCED LIGO

07/24/09

**Optical Layout and Parameters for the Advanced
LIGO Cavities**

Muzammil A. Arain and Guido Mueller

Distribution of this document:
LIGO Science Collaboration

This is an internal working note
of the LIGO Project.

California Institute of Technology
LIGO Project – MS 18-34
1200 E. California Blvd.
Pasadena, CA 91125
Phone (626) 395-2129
Fax (626) 304-9834
E-mail: info@ligo.caltech.edu

Massachusetts Institute of Technology
MIT NW22-295
185 Albany Street
Cambridge, MA 02139 USA
Phone (617) 253-4824
Fax (617) 253-7014
E-mail: info@ligo.mit.edu

LIGO Hanford Observatory
P.O. Box 1970
Mail Stop S9-02
Richland WA 99352
Phone 509-372-8106
Fax 509-372-8137

LIGO Livingston Observatory
P.O. Box 940
Livingston, LA 70754
Phone 225-686-3100
Fax 225-686-7189

<http://www.ligo.caltech.edu/>

Table of Contents

1	<i>Introduction</i>	3
1.1	Purpose and Scope	3
1.2	Definitions	3
1.3	Acronyms	3
1.3.1	LIGO Documents	3
2	<i>Optical Configuration</i>	4
2.1	Derived Cavity Parameters	6
2.2	BS and Schnupp Asymmetry	7
2.3	Low Power Operation	7
2.4	Changing P(S)R2 Position for Optimizing for Low Power	8
2.5	Adaptive Mode Matching from IOO to IFO at Low Power	9
3	<i>Summary</i>	11

1 Introduction

1.1 Purpose and Scope

This document describes the optical parameters of the various cavities for Advanced LIGO. The lengths between various optical elements, ROC values of the mirrors, and their tolerances are listed. The various cavity parameters like Finesse, linewidths, and transversal mode spacing are calculated. Included also are the higher order modes offsets from resonances. The recycling cavity parameters are picked assuming that the TCS keeps the IFO same for both the cold as well as full power operation. The RC can be matched to a low power IFO state by appropriately changing the recycling cavity mirrors positions.

1.2 Definitions

Finesse: Measure of the selectiveness/build-up of the cavity given by $F = \frac{\pi\sqrt{r_1 r_2}}{1 - r_1 r_2}$

Free Spectral Range: FSR is given by $FSR = \frac{c}{2L}$ where c is the speed of light while L is the length of the cavity. The units we use are Hz.

Linewidth: The point at which the normalized transmission through a cavity becomes 1/2. This is calculated as $Linewidth = \frac{0.5 * FSR}{F} = \text{Half-Width-Half-Max (HWHM)}$.

Transversal Mode Spacing: Transversal mode spacing in the frequency difference between two Gaussian modes. For example, this is the frequency difference between TEM₀₀ mode and TEM₀₁. For any higher order TEM_{nm} mode, the difference between TEM₀₀ and TEM_{nm} mode is given by $\frac{(n+m)FSR * a \cos(\pm\sqrt{g})}{\pi}$ where g is the G-factor of the cavity. Note that we will use Hz as the units of transversal mode spacing.

Sagitta or Sag: For a beam with 1/e² beam size of w incident on a mirror if ROC R, the sag is given by $\frac{w^2}{2R}$ between the center of the beam and the beam radius.

1.3 Acronyms

ROC: Radius of Curvature

PRC: Power Recycling Cavity

SRC: Signal recycling Cavity

1.3.1 LIGO Documents

1. Michael Smith and Dennis Coyne, "Stable Recycling Cavity Mirror Coordinates and Recycling Cavity Lengths," LIGO-T080078-06-D.

2. Muzammil A. Arain and Guido Mueller, "Design of the Advanced LIGO recycling cavities," *Opt. Express* 16, 10018-10032 (2008) <http://www.opticsinfobase.org/abstract.cfm?URI=oe-16-14-10018>.
3. R. Abbott et al., "Advanced LIGO Interferometer Sensing and Control Conceptual Design," LIGO-T070247-00-I.
4. http://ilog.ligo-wa.caltech.edu:7285/advligo/Pickle_results?action=AttachFile&do=get&target=ASC_07May09.ppt

2 Optical Configuration

The optical configuration of the Advanced LIGO cavities is given in Fig. 1 where we include both recycling cavities and the arm cavities.

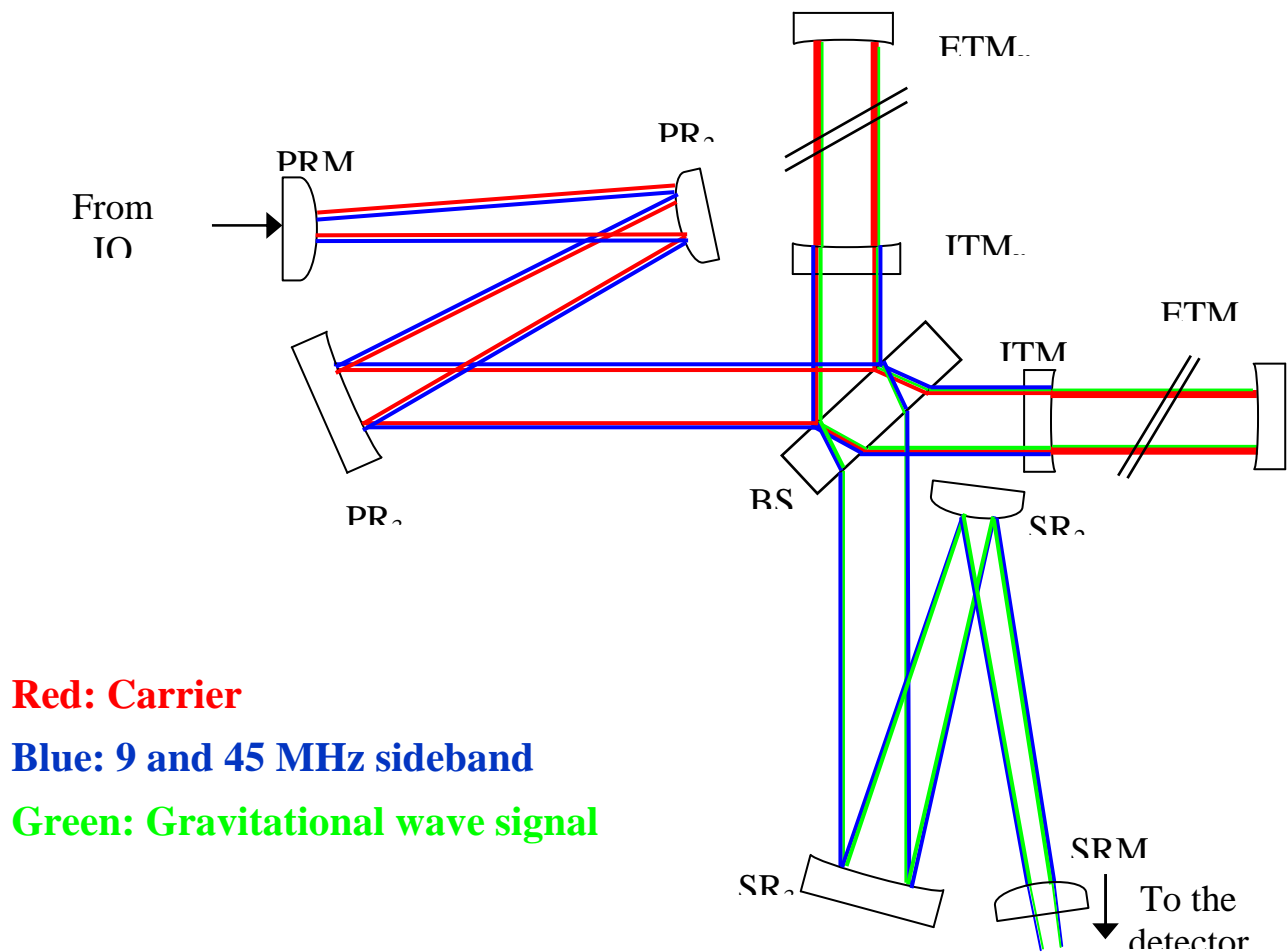


Fig. 1: Optical layout of Advanced LIGO cavities.

The various distances involved between optical elements are taken from Ref. 1 and 2. These are shown in Table 1. Various lengths have been taken from Ref. 1. The criterion of designing the recycling cavities and the related Gouy phase selection has been discussed in Ref. 2.

Table 1: Optical Parameters and Distances in Advanced LIGO Cavities
Recycling Cavity Parameters 25° PRC and 19° SRC Gouy Phase

Definition	Unit	PRC		SRC	
		Straight	Folded	Straight	Folded
P(S)RM radius of curvature	m	-10.997	-8.8691	-5.6938	-11.5806
Distance b/w P(S)RM and P(S)R ₂	m	16.6037	15.7971	15.726	15.941
P(S)R ₂ ROC	m	-4.555	-4.41	-6.427	-4.867
Distance b/w P(S)R ₂ and P(S)R ₃	m	16.1558	15.2067	15.4607	16.0079
P(S)R ₃ ROC	m	36	34	36	36
Distance b/w P(S)R ₃ and BS	m	19.5384	19.4221	19.368	20.1072
BS Effective thickness	mm	0	0	131.5	132
Distance b/w BS and CPy	m	4.8497	9.4767	4.8046	9.4314
Distance b/w CP and ITM	mm	5	5	5	5
ITM ROC	m	1934	1934	1934	1934
Reqd. beam waist size in arm	mm	12.0	12.01	12.0	12.01
Beam Size at ITM*	mm	5.30	5.31	5.30	5.31
Beam waist location from ITM	m	1884.4	1885	1884.4	1885
Arm Cavity Length	m	3994.5	3996.0	3994.5	3996.0
ETM ROC	m	2245	2245	2245	2245
Beam Size at ETM*	m	6.2	6.21	6.2	6.21
Schnupp Asymmetry (X _{SML} -Y _{SML})	mm	50.4	-50	50.4	-50
Angle of Incidence at P(S)R ₂	degree	0.79	0.963	0.87	0.878
Angle of incidence at P(S)R ₃	degree	0.615	1.144	0.785	0.916

SML = Short Michelson Length, * Beam Size mentioned are 1/e² (Intensity) beam radius

Table 2: Component Parameters for Recycling Cavity Mirrors

Optics	ROC (m)		Beam Size (mm)		Sag (μm)		ROC Tolerance in % and mm			Tol. Sag (nm)	
	Straight	Folded	Straight	Folded	Straight	Folded	Both (%)	Straight (mm)	Folded (mm)	Straight	Folded
PRM	-11.00	-8.87	2.2	2.1	-0.23	-0.24	1	-110.0	-88.7	-2.3	-2.4
PR2	-4.56	-4.41	6.2	6.3	-4.18	-4.54	0.5	-22.8	-22.1	-20.8	-22.6
PR3	36.00	34.00	54.0	54.5	40.46	43.62	0.5	180.0	170.0	201.3	217.0
SRM	-5.69	-11.58	2.1	2.6	-0.38	-0.30	1	-56.9	-115.8	-3.8	-3.0
SR2	-6.43	-4.87	8.2	6.6	-5.27	-4.49	0.5	-32.1	-24.3	-26.2	-22.4
SR3	36.00	36.00	54.0	54.2	40.50	40.80	0.5	180.0	180.0	201.5	203.0

Note: Here ‘Sag’ is the sagitta change due to ROC while ‘Tol Sag’ is the change in sagitta between the nominal ROC value and when the ROC is at the end of the tolerance. For example, for PRM, ‘Tol. Sag’ = (Beam size)²/(2*11) - (Beam size)²/{2*(11+0.1)}

Note that the tolerances of P(S)R₃ are based upon our ability to correct any manufacturing tolerance by repositioning P(S)R₂. From layout standpoint, we can reposition P(S)R₂ by ± 10 cm requiring P(S)RM be moved by ± 20 cm. Thus we had to select 0.5% tolerance for P(S)R₃. Any error in ROC of P(S)R₂ and P(S)RM can also be corrected by repositioning the mirrors but the range of motion required for these mirrors is small.

2.1 Derived Cavity Parameters

To derive cavity parameters we have to use some mirror transmittances and distances. These are given in Table 4 and are taken from Ref. [1-3].

Table 3: Derived Cavity Parameters

<i>Quantity</i>	<i>Unit</i>	<i>Straight IFO (Folded)</i>	
		<i>PRC</i>	<i>SRC</i>
ITM Transmittance	%	1.4	
PRM Transmittance	%	3.0	
SRM Transmittance	%	20.0	
ETM Transmittance	ppm	5	
PRC Length	m	57.656 (60.411)	
SRC Length	m	56.008 (62.137)	
Input Mode Cleaner Round Trip Length	m	32.9461(34.513)	
Input Mode Cleaner Finesse		520	
Arm cavity length	m	3994.5 (3996)	
Lower Mod. Frequency= IMC FSR	MHz	9.099471 (8.684428)	
Upper Mod. Frequency	MHz	45.497355(43.42214)	
Arm cavity Finesse		443	
Arm cavity FSR	KHz	37.52	
Arm cavity TMS	KHz	32.453	
Arm cavity Linewidth	Hz	42.33	
Arm cavity G-factor		0.8303	
G-factor PRC		0.8214	
One way Gouy Phase PRC	Radian	25	
G-factor SRC		0.8699	
One way Gouy Phase SRC	Radian	19	
Straight Interferometer		<i>PRC</i>	<i>SRC</i>
Carrier Recycling cavity Finesse		114	26
Recycling cavity FSR	MHz	2.6	2.67
Recycling cavity TMS	MHz	0.3611	0.2825
Carrier Recycling cavity Linewidth	KHz	10.98	52.68

2.2 BS and Schnupp Asymmetry

Note that we have used the HR side of the BS for designing the PRC while for the SRC, the beam passing through the BS AR side is chosen. Thus for the straight cavity, PRC is designed for the Y-arm while SRC is designed for the X-arm. For the folded cavity, PRC is designed for the X-arm while SRC is designed for the Y-arm. The difference in the resulting ROC for the cavity mirrors is very small and well within the proposed tolerance.

One important thing to note that the optical thicknesses play a different role for the cavity length (or phase) locking and the mode matching. For the optical phase or cavity length calculations, when the light beam passes through a substrate, the optical phase accumulated as $n \cdot d$ where n is the refractive index while d is the thickness of the material. For the case of mode matching, a substrate of thickness d is modeled as n/d . So the effective thickness is reduced. This has an important significance when considering the BS thickness and the Schnupp asymmetry. When considering the PRC X-arm, this arm travels through the BS. So the ‘optical thickness’ is larger than the actual thickness of the BS by a factor of n , i.e., the refractive index. In assigning the Schnupp asymmetry, currently X-arm has a longer arm length than the Y-arm when considering the optical phases and the cavity lengths. However, when considering the beam propagation, because of the n/d effect in the BS thickness, the beam propagating through the BS sees n/d as the thickness. This difference more or less makes the mode matching same into the X-arm and the Y-arm. The Schnupp asymmetry is reversed in the case of folded IFO which preserves the advantage because now the Y-arm beam passes through the BS.

2.3 Low Power Operation

Parameters given in Section 2 are basically for the cold IFO state. TCS is supposed to preserve the IFO mode as the IFO is locked at higher power. This is done by using ring heaters on the test masses and by using CO_2 beam on the compensation plate.

However, we can envision that we may want to operate the IFO at low power to take advantage of the various IFO configurations available. As mentioned in Ref. 3, there are significant astronomical interest in operating at or near 25 W. The thermal lens in the ITMs is the main factor that would change the mode matching. It would be desirable to design the RC to match some intermediate power level such that we can operate the IFO without engaging TCS at low powers. Since we want good mode matching at the beginning also for locking purposes, it seems prudent to design the RC for some intermediate power level between the cold and the 25 W operation. We can do so by repositioning the RC mirrors. We have chosen to present the design of the RCs for 12.5 W assuming that we do not engage the TCS. Therefore, we consider that there will be a thermal inside the ITM. The change of mode due to ITM HR ROC change is very insignificant as compared to the ITM substrate thermal lens and therefore can be neglected. From various simulations and thermal modeling, thermal lens of about 5 km at 125 W for 5.3 cm beam size is expected. This translates into a 50 km thermal lens at 12.5 W. Therefore, we can design the RC such that we include a 50 km thermal lens in the ITM substrate. The mode matching between the arm cavity mode and the recycling cavity, and mode matching product of AC mode, RC mode, and IMC mode is presented in Fig. 3.

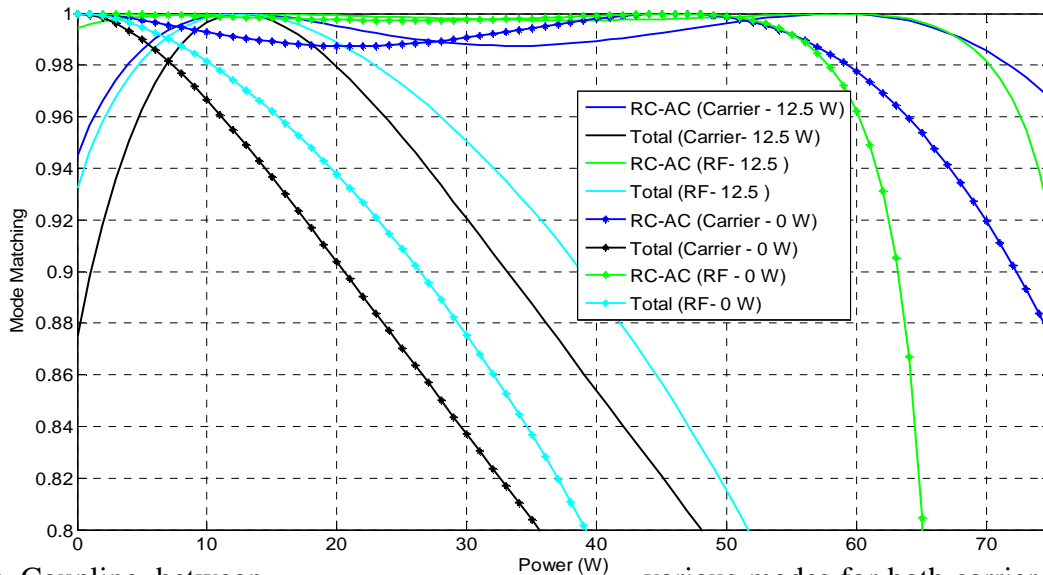


Fig. 3: Coupling between various modes for both carrier and the RF sidebands. Here RC-AC represents the coupling between the arm cavity mode and the recycling cavity mode. Total represents the product of coupling between AC-RC and IMC-RC. The dotted curves represent the mode matching for the IFO at 0 W. The solid lines without markers are for 12.5 W.

As is evident from Fig. 3, there is considerable advantage in designing the system for low power operation. We can gain almost 8% mode matching improvement at 25 W if we design the system for 12.5 W. Note that 12.5 W is chosen such that the mode matching is good for both zero power and 25 W. The data in Fig. 3 has been generated using modal model of the IFO developed at UF. This should be checked against FFT model.

2.4 Changing P(S)R2 Position for Optimizing for Low Power

The idea is to design the RC for cold power and if needed be at the time of commissioning, optimize the distance between PR2-PR3 to mode match for a certain low power. This would help in initial commissioning where TCS may not be needed for low power operation. Here low power means 20-30 W.

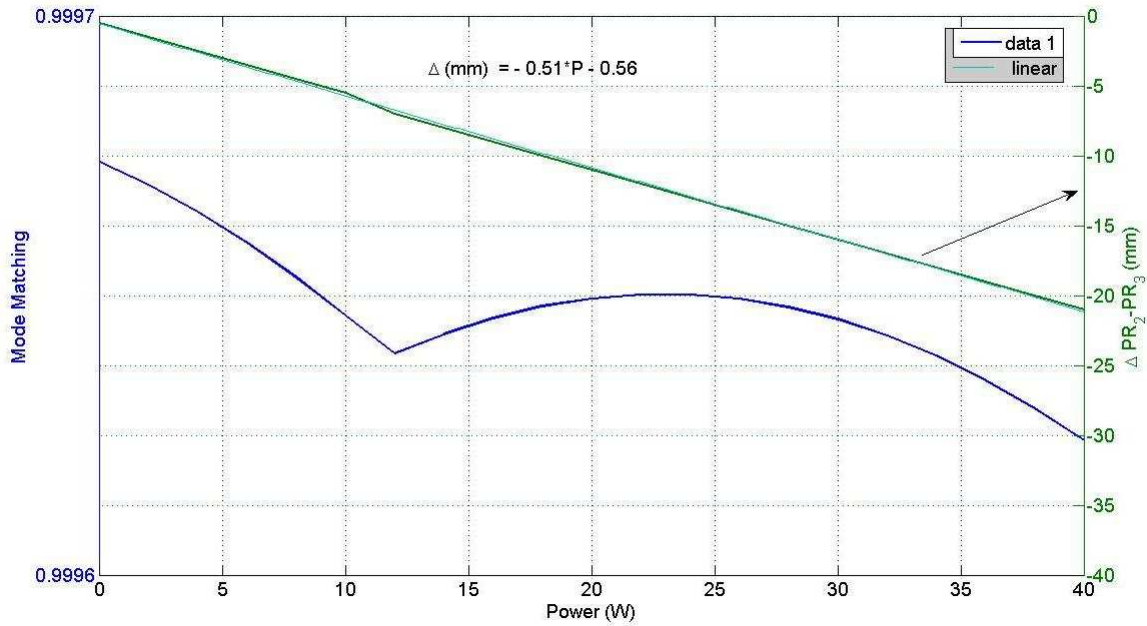


Fig. 4: This is an example of how to use PR2 to optimize the mode matching as a function of Power.

2.5 Adaptive Mode Matching from IOO to IFO at Low Power

IOO plans to install one (possibly two) adaptive lenses in the IOO chain so that the mode matching at all powers can be adjusted for the thermal effects. A new material (SF57 from SCHOTT) is undergoing tests for adaptive mode matching. So far all the results indicate that this would be a very good candidate for the adaptive lenses. If we use two, we can easily mode match perfectly at all power levels. Figure 4 shows the mode matching for various levels of adaptive elements in operation. The black curve shows how the mode-matching would be affected from IOO to the IFO if we have no adaptive lens in the IOO chain. Red curve shows the case for one adaptive element while the blue curve indicates the mode matching if we have two adaptive elements in the IOO chain. The location of the two adaptive elements chosen for this simulation is one after the PMMT₂ and the second just before SM₂. Figure 5 shows the ROC change required from the adaptive elements at these locations.

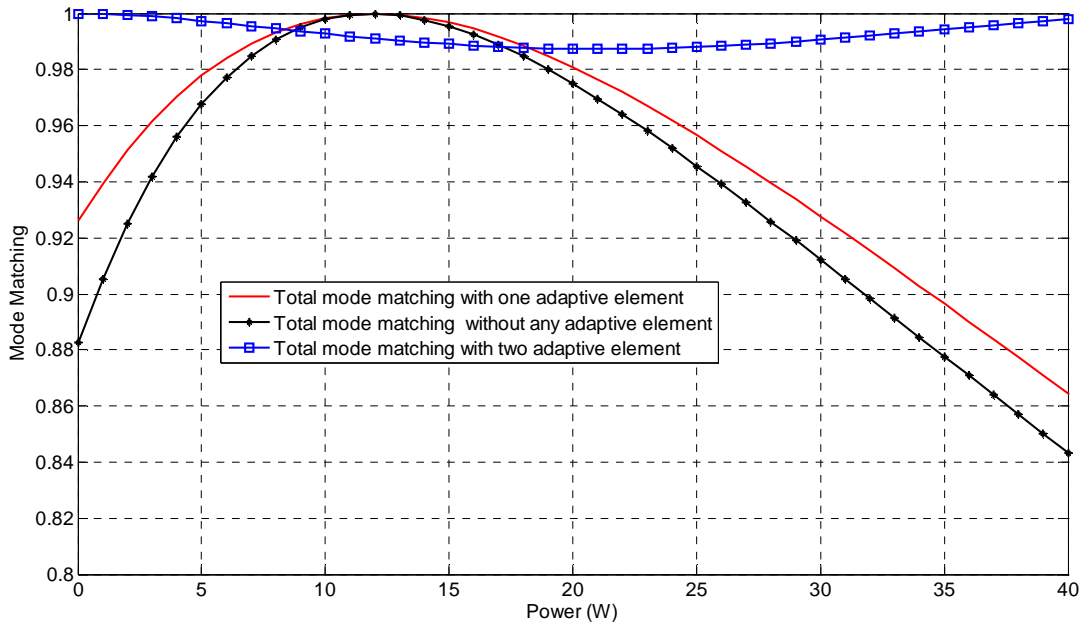


Fig. 4: Total mode matching from IOO to IGO including both IOO-RC and RC-AC mode matching for various adaptive compensation schemes. Mode matching is for the carrier.

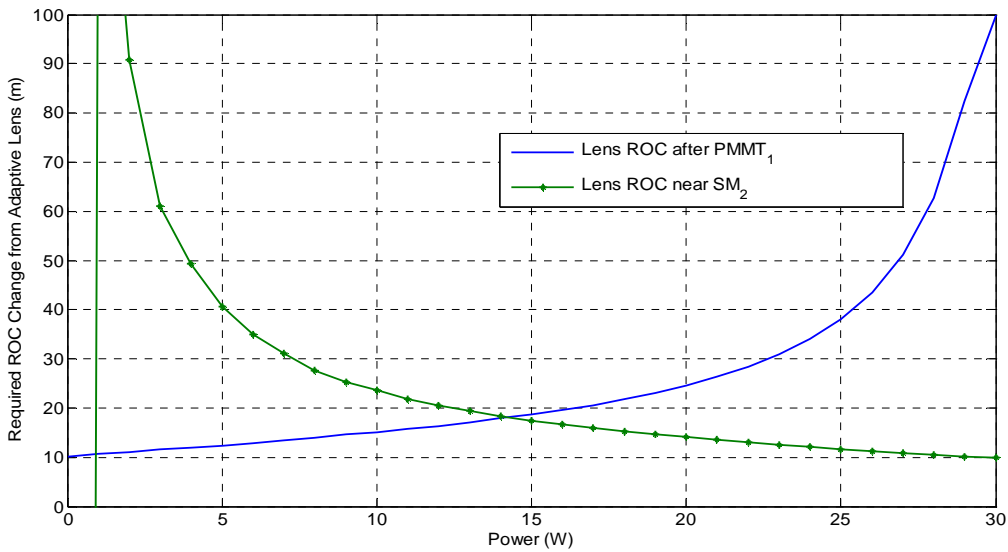


Fig. 5: ROC change required at from two adaptive lenses located after PMMT₂ and just before SM₂ respectively to get the mode matching as indicated by the blue curve in Fig. 4. The values shows are the ROC values at about 2 mm beam size.

3 Summary

We have presented the optical parameters for the various cavities in Advanced LIGO. We have checked the possibility of designing the system for reduced power operation such that we do not have to engage TCS for correction. The choice of PRC and SRC Gouy phase of 25 degree and 19 degree (one way) is inspired by ISC's modeling of ASC.⁴ The proposed values are not optimal for the mode matching performance but rather are a compromise between ASC and thermal performance.